

DISORIENTED CHIRAL CONDENSATE: THEORY AND PHENOMENOLOGY***

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The basic ideas underlying the production dynamics and search techniques for disoriented chiral condensate are described.

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1. Introduction

These notes are an abbreviated version of lectures given at the 1997 Zakopane School. They contain two topics. The first is a description in elementary terms of the basic ideas underlying the speculative hypothesis that pieces of strong-interaction vacuum with a rotated chiral order parameter, disoriented chiral condensate or DCC, might be produced in high energy elementary particle collisions. The second topic is a discussion of the phenomenological techniques which may be applied to data in order to experimentally search for the existence of DCC.

Two other topics were discussed in the lectures but will not be mentioned in these notes other than in this introduction. One was a review of the experimental situation regarding DCC searches. There are so far only two such attempts. One has been carried out at the Fermilab TeVatron collider by test/experiment T864 (MiniMax). Preliminary results, all negative, have been presented at a variety of conferences [1]. No new information is available now, and the interested reader is invited to consult the references [2]. The other experiment, WA98, is in the fixed-target heavy-ion beam at CERN. Again there is no evidence of DCC production [3]. The

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analysis methods are at present being refined by that group and are different than for MiniMax, because they are blessed with a relatively large phase-space acceptance. A recent status report is given by Nayak [4].

The other omitted topic is quite separate, and has to do with the initiative called FELIX. It is a proposal for a full-acceptance detector and experimental program for the LHC dedicated to the study of QCD in all its aspects — hard and soft, perturbative and non-perturbative. Much has happened since Zakopane with respect to FELIX. Especially noteworthy is the production of a lengthy and detailed Letter of Intent [5], which provides much more detail than was possible in the lectures on what it is about, and in any case provides guidelines for all LHC experiments on interesting issues and opportunities in QCD worthy of study. Unfortunately, at this time of writing, the initiative has run into difficulties with the CERN committees and management, with its future at present uncertain.

2. Theory of DCC production

2.1. *Light-quark effective theory (LQET)*

We accept without question that the usual QCD Lagrangian provides a correct description of the strong interactions. Nevertheless, at large distances the spectrum of the theory is that of colorless hadrons, not colorful quarks and gluons. The confinement mechanism responsible for this situation is only part of the story. In addition there is an approximate chiral $SU(2)_L \times SU(2)_R$ flavor symmetry which is spontaneously broken. The pions are the collective modes, or Goldstone bosons, associated with this phenomenon of spontaneous symmetry breaking. In addition, in this low-energy world where hadron resonances are a dominant feature, the constituent quark model works quite well, with an interaction potential which does not involve in any explicit way the gluons; direct gluonic effects seem to be quite muted [6].

There are a variety of low-energy effective Lagrangians which are in use, associated with this regime. And they are quite well-motivated, with a starting point being the fundamental QCD short-distance Lagrangian. The procedure of starting at short distances and ending up with a large-distance effective theory depends strongly upon taking into consideration the effects of instantons [7]. These lectures are not the place to go into what instantons are, and it has to be assumed that the uneducated reader will search elsewhere to find out [8]. It is rather clear on phenomenological grounds, and is supported by detailed variational calculations, that the important instantons have a size characterized by a momentum scale of about 600 MeV and that the size distribution is quite sharply peaked about this value [9]. The

instantons form a dilute liquid in (Euclidean) space-time, with a packing fraction of only 10%. Nevertheless, the interactions between them, and the effects of the instantons on the Fermi sea of light quarks, are very important. There are light-Fermion “zero modes” surrounding the instantons, and their correlations are such as to rearrange the Fermi sea of light quarks in just the right way to create the chiral symmetry breaking.

Assuming that these instantons are indeed the most important gluonic configurations at the 600 MeV scale, their main effect when “integrated out” of the QCD path integrals, is to leave behind an effective action of the Nambu-Jona-Lasinio type between the light quarks. This effective action, to be applied at momentum scales below 600 MeV, does indeed imply spontaneous chiral symmetry breaking and the existence of the pionic Goldstone bosons, which emerge as composites of the quark-antiquark degrees of freedom. It also constitutes a definite starting point in general for constituent-quark spectroscopy. An extensive amount of work utilizing this effective action is quite successful phenomenologically [9].

At still lower momentum scales, or larger distance scales, the constituent-quarks themselves can be “integrated out” of the effective action. They are replaced by the pionic degrees of freedom, comprising the lowest mass scale, or largest distance scale, in the strong interactions. The effective action looks then very much like the one used for the Higgs sector. However the action of this effective theory need not be restricted to be renormalizable. There will be, in addition to the quadratic free-particle term and quartic interaction, terms of higher polynomial order, some with derivative couplings depending upon the choice of description. This is just the purely chiral effective action studied in great detail by Gasser and Leutwyler, among others [10].

Simplified versions of the chiral effective theory are the linear and nonlinear sigma models. The linear sigma model is what is isomorphic to the usual Higgs theory; the nonlinear sigma model essentially “integrates out” the massive sigma degree of freedom (analogous to the massive Higgs boson of standard-model electroweak theory), leaving only the pionic degrees of freedom in the effective action. In what follows, these effective actions are what are relevant for the description of DCC because, as we shall see, the space-time scale for the DCC system is assumed to be “macroscopic”, *i.e.* have dimensions large compared to 1 fermi. However, we should emphasize here that the use of sigma models, be they linear or nonlinear, is only a rough approximation to the true situation at best, because the Yukawa couplings of the Goldstone system of pions and sigma to constituent quarks are large and should not be ignored, except perhaps at the very lowest momentum scales.

2.2. DCC production scenarios

As we have mentioned in the introduction, DCC is by assumption a “macroscopic” region of space-time within which the chiral order parameter is not oriented in the same direction in the internal $O(4) = SU(2) \times SU(2)$ space as the ordinary vacuum. If we use for simplicity the language of the linear sigma model, the sigma and pion fields form a four-vector ϕ in the $O(4)$ space, and in the vacuum there is a nonvanishing value of ϕ in the σ direction.

What we shall assume is that in a high energy collision, there are regions of space-time where $\langle\phi\rangle$, essentially a classical quantity, is rotated away from the σ direction. At late times, of course, this “chirally rotated vacuum” must relax back to ordinary vacuum. The mechanism will clearly be the radiation of the collective modes, the pions, via a semiclassical mechanism. It should be clear that, if indeed this happens, it is quite an interesting phenomenon, because it would provide a rather direct look into the properties of the QCD vacuum itself.

How might the DCC be produced? In hadron-hadron collisions, most models of particle production, be they stringy or partonic, put the bulk of the space-time activity near the light cone. That is, the flow of produced quanta is concentrated in a rather thin shell expanding from the collision point at the speed of light. But what happens within the interior of this shell? If it rapidly relaxes to vacuum, then it is hard to see why the normal vacuum must be chosen, because the interior region is separated from the exterior true vacuum by a “hot” shell of expanding partons. And as we shall see, the time scale for the true vacuum to be selected via the small chiral-symmetry-breaking effects associated with nonvanishing pion mass is quite long.

It is this “Baked-Alaska” scenario which will be the main thrust of this discussion [11]. However, there is a second space-time scenario relevant to heavy ion collisions. The idealized case is that of infinite pancakes colliding with each other at the speed of light, with quark-gluon plasma produced in between the pancakes at early times. But at late times, after this plasma goes through the deconfining/chiral phase transition, the possibility of DCC formation also exists [12] and has been rather extensively studied [13]. The geometry in this idealized situation is that of a boost-invariant 1+1 dimensional expanding system, and is the most tractable case to consider from a calculational point of view. We shall review a simple example studied by Blaizot and Krzywicki in the next section.

Can one actually predict that DCC should be produced under these circumstances? Certainly not. On the other hand can the possibility of DCC production be excluded theoretically? Again, certainly not. The subject

needs to be data driven. This is the reason that both Cyrus Taylor and I decided to go into experimental physics and make an experimental search at the TeVatron collider. While the MiniMax experiment which emerged is very modest, and while the results so far are negative, the experience gained has been invaluable in learning how to construct good search strategies, and in determining what is necessary in order to do a better experimental job in the future.

2.3. Sigma models

The space-time scenario for DCC production begins at an early proper time, *i.e.* near the light cone, when the energy density becomes low enough that the chiral order parameter is nonvanishing. The proper time scale here is plausibly somewhere between 0.3 and 0.8 fermi. At this point one must assume initial conditions for the chiral field, which then evolves, at the simplest level of approximation, according to the classical equations of motion. A natural hypothesis is, in the context of the linear sigma model, that the chiral field initially vanishes (at least on average). This means it is initially on top of the “Mexican hat”, and rolls off the hat into the region of the minimum. During this initial roll, the linear sigma model, at the least, should be used for the proper-time evolution of the vacuum condensate. As we shall see, this evolution can go on for anywhere from one to five fermi of proper time. The formalism is as follows. The chiral fields can be written as

$$\phi = (\sigma, \vec{\pi}) \quad i = 1, \dots, 4 \quad (1)$$

or alternatively as a 2×2 matrix of fields

$$\Phi = \sigma + i \vec{\tau} \cdot \vec{\pi} \quad (2)$$

with the Lagrangian for the former case being

$$\mathcal{L} = \frac{1}{2} |\partial_\mu \phi|^2 - \frac{\lambda}{4} (\phi^2 - f_\pi^2)^2 \quad (3)$$

At late times, when the chiral field is near the minimum of the Mexican-hat potential radially, but is still rolling around azimuthally, the nonlinear sigma model can be used. One writes, in the 2×2 notation

$$\Phi = R e^{i \vec{\tau} \cdot \vec{\pi}} \quad (4)$$

and freezes out the radial degree of freedom

$$R \approx f_\pi \quad (5)$$

If one further writes as a special case

$$\Phi = f_{\pi} e^{i\tau_3 \pi_3(x)} \quad (6)$$

it follows that the pion field which appears in the exponential can be shown to obey free-field equations of motion.

$$\square \pi_3(x) = 0 . \quad (7)$$

This can be extended using a general, global $SU(2) \times SU(2)$ rotation to a class of solutions ("Anselm class" [14]):

$$\Phi \rightarrow U_L \Phi U_R^{\dagger} . \quad (8)$$

It is important to keep in mind in what follows that the linear sigma model leads to nonlinear equations of motion, while the nonlinear sigma model here leads to linear equations of motion.

While this discussion has been only at the classical level, there has been a lot of work, especially by those interested in the heavy-ion DCC, in going beyond this approximation. The state of the art is to incorporate quantum corrections to the linear sigma model in the mean-field (or Hartree, or large-N, or "random-phase") approximation [15]. This amounts essentially to inclusion of tadpole contributions to self-energy parts or geometrical-series bubble summations to appropriate propagators. While this level of computation is simple in translationally-invariant systems, it is not at all simple in cases like this, where the space-time geometry of sources is quite nontrivial. So far, the inclusion of quantum effects is important quantitatively but has not seriously changed the qualitative properties. In this discussion we shall remain within the classical approximation.

2.4. The Blaizot-Krzywicki model

A very easy application can be made of the above considerations [16]. Consider boost-invariant solutions of the nonlinear sigma-model equations within infinitely large heavy-ion disks receding from each other at the speed of light after a central collision. The equation of motion

$$\left(\frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial z^2} \right) \pi = 0 \quad (9)$$

leads to the solution

$$\pi = \alpha \ln \tau + \text{const.} \quad (10)$$

with

$$\tau^2 = t^2 - z^2 \quad (11)$$

and

$$\Phi = f_\pi \left(U_L e^{i\alpha\tau_3 \ln \tau} U_R^\dagger \right) . \quad (12)$$

We see that in general the chiral angle in isospin space precesses with increase of proper time, provided the left-handed global rotation differs from the right-handed one. This curious feature depends on the choice of geometry, however, and does not generalize to the Baked-Alaska scenario.

These calculations have been extended to the case of the linear sigma model, with qualitatively similar results.

2.5. Baked Alaska

The simplest case of the Baked-Alaska scenario again utilizes the nonlinear sigma model. Inside the future light cone one assumes spherical symmetry. In addition, it is assumed that for $0 < t < T$, with T some “decoupling time”, there is chirally rotated vacuum within the light cone, *i.e.*

$$\phi = (0, f_\pi, 0, 0) . \quad (13)$$

We have taken, without much loss of generality, the chiral orientation to be purely in the π^0 direction, because other cases can be obtained via global chiral rotations. With this hypothesis, the free equation of motion for the pion field can be used for times $t > T$, using the boundary conditions for the field at $t = T$. When the pion mass is neglected, the form of the solution must be that of a right-moving plus left-moving pulse:

$$r\pi(r, t) = f(r - t) + g(r + t) \Rightarrow f(r - t) - f(r + t) \quad (14)$$

and the shape of the function f is easily found to be a triangular pulse of width $2T$ and maximum height $f_\pi T$.

$$r\pi(r, t) = \frac{f_\pi}{2} [(2T - t + r)\theta(2T - t + r)\theta(t - r) - (2T - t - r)\theta(2T - t - r)] . \quad (15)$$

Thus for times $t > 2T$ there is only a triangular pulse of pion field (actually $r\pi$) radiated outward at the speed of light, which comprises the decay products of the initially formed DCC.

The calculations and formalism for this example have been laid down in considerable detail in a recent paper [17], to which the interested reader is referred for details. In particular, the calculations have been redone for nonvanishing pion mass, as well as for the linear sigma model. An important feature of this class of solutions is that, when one generalizes them away from the production of π^0 field by applying a global $SU(2) \times SU(2)$ rotation, the

right-handed and left-handed SU(2) rotations (*c.f.* Eq. (8)) must be the same in order that the vacuum at large times inside the light cone is the same as the ordinary vacuum.

Another way of looking at all this is in terms of sources. The free wave equation does not hold on the light cone for times $t < T$. It is in that region that the hot shell of expanding partons or other quanta separates the DCC vacuum on the inside from the true vacuum on the outside. So one has, everywhere in spacetime

$$\square \pi(x) = J(x) \; , \tag{16}$$

with the source J having support only on the light cone. A careful look at the discontinuity of the vacuum fields across the light cone leads to the expression

$$J(x) = \frac{\pi f_\pi}{r} \delta(t-r) \theta(T-t) \; . \tag{17}$$

Now there is a relation, well-known from the Bloch-Nordsieck treatment of semiclassical electromagnetic radiation, between the source which creates a classical field and the spectrum of quanta produced by the source. In this case [18] it is

$$2E \frac{dN}{d^3p} = \frac{1}{(2\pi)^3} \left| \tilde{J}(p) \right|^2 \; , \tag{18}$$

where $\tilde{J}$ is the Fourier transform of J , put on mass shell.

$$\tilde{J}(p) = \int d^4x \, e^{ip \cdot x} J(x) \qquad p^2 = m_\pi^2 \; . \tag{19}$$

It follows that, with pion mass neglected, the expression for $\tilde{J}$ is

$$\tilde{J}(p) = \frac{4\pi^2 f_\pi}{|p|} \int_0^T dt \, e^{ipt} \sin pt \; . \tag{20}$$

One readily checks that this agrees with the spectrum calculated explicitly from the equations of motion. Note that the spectrum has a high-momentum tail coming from the delta-function support on the light cone, but that the typical momentum scale is of order $1/T$.

It is an important feature of this radiation that it is coherent, which means that for a specific choice of source the multiplicity distribution of produced particles is Poisson-distributed, and that there is no Bose-Einstein enhancement. However, upon averaging over sources, in particular their chiral orientations, these simple features undergo essential complications.

2.6. Radially boost-invariant Baked Alaska

Another simple and interesting example of DCC production has been considered by Lampert, Dawson, and Cooper [19] (hereafter LDC). They assume a Baked-Alaska scenario, within the linear sigma model, which has full boost invariance. In other words, the chiral fields existing within the light cone depend only upon the proper time. This scenario cannot be regarded as realistic, because the particle distribution must look the same in all reference frames, and hence cannot have a limited spectrum of energies. They argue that one can extract the physics by sampling the distribution at large proper times within a small space-time region. However, this is not a faithful description of what a real piece of experimental apparatus would see.

A better strategy in my opinion is to use their solutions up to time T , and then assume, as was done in the previous section, that there is at later times no explicit source on the light cone. In other words the LDC solution at time T is evolved via the source-free linear sigma-model equations of motion, in order to generate the asymptotic fields. Some work has been done along these lines [20], but not enough to report here.

However, let us return to the simplified version of LDC. At the classical level the field equations become ordinary coupled differential equations, and can be easily solved numerically:

$$\left[\frac{1}{\tau^3} \frac{d}{d\tau} \tau^3 \frac{d}{d\tau} + \lambda(\sigma^2 + \pi^2 - f_\pi^2) \right] \begin{pmatrix} \sigma \\ \pi \end{pmatrix} = \begin{pmatrix} f_\pi m_\pi^2 \\ 0 \end{pmatrix}. \quad (21)$$

However LDC do better, and include the mean-field quantum corrections. It has been found, however, that the quantum effects do not qualitatively or even quantitatively make a big difference [17].

What is noteworthy about the solutions is that it takes a rather long proper time for the initial chiral fields to “roll” into the minimum of the potential; the time taken for the sigma field to settle down to near its vacuum value is about 5 fermi. This long proper-time interval, even in the presence of pion mass, would indicate the credibility of scenarios which create DCC from deep within the light cone. However, the finite extent of the source on the light cone needs to be investigated, as well as the effect of considering a statistical ensemble of initial conditions for the initial “roll”, before drawing serious conclusions.

2.7. Doing better

All the theoretical attempts we have described are extremely idealized. There does exist some numerical simulation work which attempts less idealized scenarios [21], but I think it fair to say that none of it is yet very near to what is needed for, say, Monte-Carlo input appropriate to real experimental searches. There are many issues to be addressed.

Perhaps the most important deficiency is that the idealized cases very likely have far too much symmetry. Looking at the Baked-Alaska scenario from the source point of view, it is probably the case that in a given event not only the source strength but also the chiral orientation of the source depends upon where in the lego phase-space one is. Since the chiral dynamics is spin-zero, the correlations in the lego phase-space are most probably short-range, of order one to two units. This means that on the sphere, near 90 degrees relative to the beam, only a steradian at a time may have the rather symmetrical structure of the classical solutions we have considered.

There is another way of looking at this. Suppose DCC is produced and observed. It almost by definition will consist of a cluster of pions of almost identical momenta. This cluster will have a rest frame, and in that rest frame the classical radiation field associated with this cluster will in fact have approximate spherical symmetry. However in general, this frame of reference is related to the observer's frame of reference by not only a longitudinal Lorentz boost, but probably also by a transverse boost. It is in fact reasonable to assume a distribution of transverse velocities of DCC such that the mean is semirelativistic, say somewhere between 0.4 and 0.8. If this is the case, and the internal relative velocities of the pions within the DCC cluster are smaller or at least no larger, then it will often be the case that the DCC in the laboratory frame will look like a coreless minijet. In fact DCC searches within minijets, using the techniques sketched in the next section, might be very fruitful.

In any case, these pieces of transversely boosted DCC, in reference frames where the longitudinal velocity is zero or small, will typically occupy of order a steradian of solid angle, indicating again that perhaps the natural correlation length for DCC is of order 1–2. However, there is an unsolved theoretical issue here. Suppose one has a piece of DCC centered at rapidity of +1, and another centered at –1 with a different chiral order parameter. They will have some small overlap at rapidity 0. How do the two pieces interact? Will there be a tendency to create a common alignment, or will they form independent domains? I believe this to be an important fundamental question; it has not yet received serious attention by theorists. The classical linear sigma model should be sufficient as the fundamental theoretical tool; the problem is the introduction of a realistic collision-geometry scenario which is still computationally tractable.

There is even a remote possibility that the nonlinearities of the linear sigma model are strong enough to promote long-range correlations in rapidity, *i.e.* to create a semiclassical structure which produces the phenomenology of a soft Pomeron. For example it looks not at all out of the question that a ladder built from pions on the sides and sigmas on the rungs, with couplings dictated by the linear sigma model, could produce a Regge intercept of unity or larger for forward hadron-hadron scattering. But I do not know how such a Reggeon ladder could be related to the classical DCC scenarios we have discussed.

2.8. Charged DCC?

Peter Lipa asked this question and, together with Brigitte Buschbeck, we have begun to look into this issue [22]. The idea is even more speculative than ordinary DCC. But if it makes any sense at all, it has the advantage that one can make the search using data sets containing charged-particle information only. This is not the case for ordinary DCC, where, as discussed in the next section, the relationship between charged and neutral pion production is what is examined. For charged DCC it is the relationship between positive and negative pion production which is the object of study.

Charged pion fields are to the real, Cartesian pion degrees of freedom ϕ_1 and ϕ_2 as circularly polarized light is to linearly polarized light. And certainly there are classical sources of circularly polarized light, either associated with vorticity of the source, or with a 90° phase difference of the sources of the two Cartesian components (or both). So from this point of view it seems not totally out of the question to imagine a similar possibility for the pions.

We are presently modeling a space-time scenario where DCC in the ϕ_1 direction is produced at positive eta, and DCC in the ϕ_2 direction is produced at negative eta, with a “domain wall” in between. We find, in the spectrum of produced DCC pions, a dipole layer in the lego plot, with an average positive charge per pion on one side of the domain wall, compensated by negative charge on the other. It is to be emphasized that this is an average charge *per pion* at the quantum level. The width of the dipole layer is 1–2 units of rapidity, as might be anticipated on general grounds. We need to do a little more work before reporting on its strength and momentum dependence.

3. DCC phenomenology

3.1. The inverse square root distribution

A basic signature of DCC production is the presence of very large event-by-event fluctuations in the fraction of produced pions which are neutral [11, 12, 14, 23]. Generic production models will give a distribution in the neutral fraction, defined as

$$f = \frac{N_{\pi^0}}{N_{\pi^0} + N_{\pi^+} + N_{\pi^-}} \quad (22)$$

which is binomial and peaked at $f = 1/3$. There will be in the classical large N limit a very small probability that, for example, all the pions are neutral. But if the pions are DCC decay products this probability is not at all so small.

The simplest estimate for the distribution of neutral fraction in DCC production assumes that the chiral orientation in (Cartesian) isospin space is random. Then a totally elementary calculation gives the result that the distribution of neutral fraction is inverse-square-root. With $f = \cos^2 \theta$ it follows that

$$dn \approx d(\cos \theta) = \frac{1}{2 \cos \theta} d(\cos^2 \theta) = \frac{df}{2\sqrt{f}}. \quad (23)$$

It is probable however that this component is to some extent immersed in generic background. Cuts in p_t , for example, may be useful in enhancing the signal.

It is also the case that it appears to be better to use an indirect technique to test for the presence of the inverse-square-root component [24]. This utilizes the machinery of multiparticle production dynamics and multiplicity distributions, to which we now turn.

3.2. Multiplicity distributions and generating functions

In the MiniMax experiment, the phase-space coverage is small, about 1.0 lego-area units. For that analysis it is reasonable to assume that within the acceptance the chiral order parameter takes a fixed value.

The raw information relevant to DCC physics is the multiplicity distribution of produced pions. Ideally they should be momentum analyzed, with the data binned in intervals of p_t . However this so far has been not done experimentally, and we simplify by ignoring the momentum degree of freedom. Let the probability of producing N pions be $P(N)$, and introduce the generating function

$$G(z) = \sum_N z^N P(N) \quad (24)$$

which contains all information about the multiplicity distribution. For a Poisson distribution, the generating function is an exponential

$$G(z) = e^{\mu(z-1)} \quad P(N) = \frac{\mu^N e^{-\mu}}{N!} . \quad (25)$$

In the general case it is a superposition of Poisson distributions with positive semidefinite weight function ρ

$$G(z) = \int_0^\infty d\mu \rho(\mu) e^{\mu(z-1)} . \quad (26)$$

For the two species of charged and neutral pions, the generalization is a generating function of two variables, again a superposition of Poisson distributions. The definition of generic pion production is that the only correlation is produced by the aforementioned ρ for the total pion multiplicity distribution, so that

$$\begin{aligned} G(z_{ch}, z_0) &= \sum P(N_{ch}, N_0) z_{ch}^{N_{ch}} z_0^{N_0} \\ &= \int_0^\infty d\mu \rho(\mu) e^{\mu[f(z_0-1)+(1-f)(z_{ch}-1)]} \end{aligned} \quad (27)$$

with the neutral fraction f approximately $1/3$.

By expanding things out, one sees that the partition into charged and neutral pions is governed by a binomial distribution; indeed this could have been the common-sense starting point. It is also the case that existing Monte-Carlo codes have the property that the distribution of the neutral fraction is approximately binomial.

For pure DCC, all one needs do is introduce the inverse square-root distribution in f as another weight factor:

$$\langle \rangle \equiv \int_0^\infty d\mu \rho(\mu) \Rightarrow \int_0^\infty d\mu \rho(\mu) \int_0^1 \frac{df}{2\sqrt{f}} . \quad (28)$$

One sees that the basic difference between generic production and DCC production is that in the former case the generating function depends only upon one variable, while for DCC it depends nontrivially upon two. A very good way of testing for the distinction is via factorial moments [25]. These are just the derivatives of the generating function with respect to the z 's at $z = 1$. In the case of only one variable, one has

$$f_1 = \langle n \rangle = \left. \frac{\partial G}{\partial z} \right|_{z=1} \quad f_2 = \langle n(n-1) \rangle = \left. \frac{\partial^2 G}{\partial z^2} \right|_{z=1} . \quad (29)$$

The normalized factorial moments are

$$F_n = \frac{1}{\langle n \rangle^n} \left. \frac{\partial^2 G}{\partial z^n} \right|_{z=1} = \frac{f_n}{(f_1)^n} . \quad (30)$$

In the case of interest, there is a two-dimensional array of normalized factorial moments which captures the information contained in the joint multiplicity distribution:

$$F_{ij} = \frac{1}{\langle n_{ch} \rangle^i \langle n_0 \rangle^j} \frac{\partial^{i+j} G}{\partial z_{ch}^i \partial z_0^j} . \quad (31)$$

However for generic production there are many relations between these, because the generating function depends upon only one variable:

$$F_{ij} = F_{(i+j),0} . \quad (32)$$

Therefore many ratios of the F_{ij} are expected to be unity

$$r_{ij} = \frac{F_{ij}}{F_{(i+j),0}} = 1 \quad (\text{Generic Production}) \quad (33)$$

while for pure DCC they can also be explicitly computed, and are far from unity:

$$r_{ij} = \frac{F_{ij}}{F_{(i+j),0}} = \frac{\langle (1-f)^i f^j \rangle \langle (1-f)^j \rangle}{\langle (1-f)^{i+j} \rangle \langle f^j \rangle} = \frac{i!(2j-1)!!}{(i+j)!} . \quad (34)$$

There are also experimental reasons why these ratios are useful. First of all, one can go from π^0 production to photon production via convolution, and the factorial moment method remains robust: one simply replaces the neutral-pion fugacity z_0 by the generating function for the gammas, a second-order polynomial in the photon fugacity. In addition, if efficiencies are not 100 percent, but are uncorrelated with total multiplicity or other global parameters, they can be incorporated in terms of modified fugacities. The properties of the resultant factorial moments, which are the direct observables experimentally, essentially do not change. That is, the bivariate factorial moment ratios directly extracted from data on production of charged hadrons and of gammas, as in Eqs. (33) and (34), still will be unity for generic production and far from unity in the case of DCC production.

3.3. Doing better: correlation functions and generating functionals

Once the experimental acceptance becomes large, or if one investigates, *e.g.* the p_t dependence of the presumed DCC fraction (large at low p_t , small at high p_t ?), then the correlation structure of the DCC order parameter, and even the generic particle distributions themselves, becomes of paramount importance. Clearly the formalism of generating functions should be preserved as closely as possible, and this is in principle a straightforward matter of replacing generating functions with generating functionals.

In the case of DCC production, the bridge from the classical calculations to phenomenology is fairly clear. For a given classical solution of, say, the linear sigma model there will be a source function

$$(\square + \mu^2) \phi(x) \equiv J(x) . \quad (35)$$

As we discussed, the multiplicity distribution is built from the squared Fourier transform of the source function, put on mass shell,

$$2E \frac{dN}{d^3p} = \frac{1}{(2\pi)^3} \left| \tilde{J}(p) \right|^2 . \quad (36)$$

The fluctuations for the classical case are Poissonian, and so the generating functional is again just an exponential in all the continuous number of fugacity variables, now parametrized by the particle momenta. Finally one should average over the choice of source function, which at the least is parametrized by choices of initial conditions for the classical field configurations (including chiral orientation). This leads to a DCC generating functional of the form

$$G_{\text{DCC}} = \int \mathcal{D}J_1 \mathcal{D}J_2 \mathcal{D}J_3 \mathcal{P}(J) \exp \sum_i \frac{d^3p}{2E(2\pi)^3} \left| \tilde{J}_i(p) \right|^2 (z_i(p) - 1) . \quad (37)$$

A generalized DCC distribution of the inverse square root type follows if

$$\mathcal{P}(J) \Rightarrow \mathcal{P}(|\vec{J}|^2) . \quad (38)$$

Thus far this generalization is reasonably straightforward, and leaves the unknown issues mostly at the level of the choice of boundary conditions and structure of the classical field equations considered in the previous section. However it is not reasonable to assume that the totality of particle production originates as DCC. And there is no consensus on what generating functional to use to describe generic particle production. Many correlation phenomena exist, some from minijets and perturbative QCD, undoubtedly some associated with the fluctuations in the number of “wounded” constituent quarks per collision, still others associated with impact parameter

dependence, and more associated with resonance production. In addition, one can consider various ways of combining the generic particle production with DCC production. Three extreme cases are as follows:

1)

$$G = G_{\text{generic}} + G_{\text{DCC}} . \quad (39)$$

This means that in a given event either generic particles are produced or DCC, but not both.

2)

$$G = G_{\text{generic}} \cdot G_{\text{DCC}} . \quad (40)$$

In this case the amount of DCC produced in a given event is not correlated at all with the amount of generic particles.

3)

$$z = z_{\text{generic}} + z_{\text{DCC}} . \quad (41)$$

In this case the amount of DCC produced is, up to binomial-distribution fluctuations, in proportion to the amount of generic particle production.

The phenomenological consequences of this distinction are very strong; a small DCC admixture of the first type is much easier to isolate experimentally than a small admixture of the third type. The case of independent production is intermediate.

3.4. DCC vs. Bose-Einstein, etc.

There is an interesting question of whether observation of the inverse square-root distribution implies observation of Baked-Alaska DCC. This is not at all clear. The inverse square-root distribution was discovered in a different context, namely the production of a cluster of pions in a maximally symmetric state. The inverse square-root behavior is also a consequence of the well-known Andreev, Plumer, Weiner (APW) description [26] of Bose-Einstein correlations. In addition, it has been shown that in the tree-level expansion of the chiral effective theory, the inverse-square root distribution emerges [27]. But none of these descriptions follows the same line as the DCC description above. In particular, the part of the APW scenario which leads to DCC-like behavior assumes a random Gaussian distribution of source functions, except with respect to the isospin degrees of freedom. In the DCC description, one computes the sources via the sigma-model equations of motion and (most probably) a random set of initial conditions. This would for sure make the distribution of DCC sources non-random functionals. And the tree-level chiral expansion does not seem to have the possibility of incorporating the Baked-Alaska physics.

It is probably the case that one should view DCC as a specific mechanism of Bose-Einstein enhancement. There seems to me nothing wrong with this point of view, because there really is an ongoing problem of describing Bose-Einstein effects from the most general point of view. In any case, for me this is not a very important issue until there is concrete experimental evidence for DCC-like behavior. First of all something needs to be seen. Thereafter there will be plenty of opportunity for figuring out what it means via further interaction between follow-on observations and the theory.

3.5. Charged DCC again

The phenomenological description of the charged DCC production discussed in Section 2.8 is similar to that of the conventional DCC. It is being worked out by Peter Lipa, Brigitte Buschbeck [22], and myself.

The major change is that the source function in momentum space is taken to be complex

$$J_i = u_i + iv_i \quad (42)$$

and the real and imaginary parts u and v are each taken as random variables. In the usual case there is assumed to be complete correlation between u and v . One finds

Before:

$$G = \int \mathcal{D}^3 J f(|J|^2) e^F \Rightarrow \int \mathcal{D} J |J|^2 f(|J|^2) \int_0^1 \frac{df}{2\sqrt{f}} e^F. \quad (43)$$

After:

$$G = \int \mathcal{D}^3 u \mathcal{D}^3 v f(|J|^2) e^F \Rightarrow \int \mathcal{D} J |J|^2 f(|J|^2) \int_0^1 df (1-f) \int_{-1}^1 dq e^F \quad (44)$$

with

$$q = \frac{\langle (N_+ - N_-) \rangle}{\langle (N_+ + N_-) \rangle}. \quad (45)$$

For the “action” of the generating function one has

$$F = \mu \left[(z_+ - 1) \left(\frac{1+q}{2} \right) (1-f) + (z_- - 1) \left(\frac{1-q}{2} \right) (1-f) + (z_0 - 1) f \right]. \quad (46)$$

Again the MiniMax-like ratios of factorial moments can be constructed, and compared with charged particle data. It appears from a cursory first look

that when averaged over all p_t the data will be of opposite sign to the charged-DCC expectation. Charge tends to be locally neutralized in the lego plot, more than it would from a random throwing of charge into phase space. This to be expected in both the string and parton cascade pictures; there is very little charge separation and flow in the space-time evolution.

However when the p_t 's are low, the experimental situation changes. It is here that the charged-DCC hypothesis is not ruled out, and the detailed phenomenology of this region of phase-space may turn out to be interesting.

The assumption of complete randomness of u and v is an extreme one. There are evidently interpolations between the extremes of no correlation and complete correlation which can be easily constructed. Study of the structure of the classical models as outlined in Section 2.8 will help in choosing a reasonable starting hypothesis.

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