THE ENERGY DEPENDENCE OF THE SATURATION SCALE *

Masaaki Kuroda

Institute of Physics, Meiji-Gakuin University, Yokohama 244, Japan

DIETER SCHILDKNECHT

Fakultät für Physik, Universität Bielefeld Universitätsstrasse 25, 33615 Bielefeld, Germany Dieter.Schildknecht@physik.uni-bielefeld.de

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At low $x\cong Q^2/W^2\ll 1$, in deep inelastic scattering, the photon fluctuates into a $q\bar{q}$ vector state that interacts via two gluons with the proton. The energy dependence is determined by the saturation scale, in our approach given by $A_{\rm sat}^2(W^2)\sim (W^2)^{C_2}$. Imposing DGLAP evolution, we find $C_2^{\rm theory}=0.27$ in agreement with the model-independent analysis of the HERA data. Different values of the exponent C_2 are correlated with different ratios of the longitudinal to transverse structure function. This stresses the need for experiments to separate longitudinal and transverse contributions in deep inelastic scattering.

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In deep inelastic scattering (DIS) at low $x \cong Q^2/W^2 < 0.1$ the photon fluctuates into on-shell quark–antiquark vector states (e.g. [1]) which interact with the proton via coupling of two gluons [2] to the $q\bar{q}$ color–dipole [3]. The cross section depends on the effective transverse three momentum of the gluon, \vec{l}_{\perp} , absorbed by the $q\bar{q}$ pair in the imaginary part of the virtual-photon forward Compton amplitude. The effective transverse momentum of the absorbed gluon gives rise to a novel scale, the "saturation scale", characteristic for DIS in the $x \to 0$ limit. Since the photon fluctuates into an on-shell $q\bar{q}$ pair, the γ^*p energy, W, besides Q^2 , is the appropriate variable

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for the virtual photon-absorption cross section to depend on. The color–dipole cross section, $\sigma_{q\bar{q}p}(\vec{r}, W^2)$, and the saturation scale, $\Lambda_{\rm sat}^2(W^2)$, become functions of W^2 .

In the high-energy limit, with

$$\Lambda_{\rm sat}^2(W^2) = \frac{1}{6} \langle \vec{l}_{\perp}^2 \rangle = \frac{1}{6} B' \left(\frac{W^2}{1 \text{ GeV}^2} \right)^{C_2},$$
 (1)

the fit to the DIS data gave [4]

$$C_2^{\text{exp}} = 0.27 \pm 0.01,$$

 $B' = 0.340 \pm 0.063 \text{ GeV}^2.$ (2)

In terms of $\Lambda^2_{\rm sat}(W^2)$, the cross section of this QCD-based generalized vector dominance — color–dipole picture (GVD-CDP) in the limit of $Q^2 \to 0$ and $Q^2 \to \infty$ reads

$$\sigma_{\gamma^* p}(W^2, Q^2) = \frac{\alpha}{3\pi} R_{e^+ e^-} \sigma^{(\infty)} \begin{cases} \ln \frac{\Lambda_{\text{sat}}^2(W^2)}{Q^2 + m_0^2}, & (Q^2 \ll \Lambda_{\text{sat}}^2(W^2)), \\ \frac{\Lambda_{\text{sat}}^2(W^2)}{2Q^2}, & (Q^2 \gg \Lambda_{\text{sat}}^2(W^2)). \end{cases}$$
(3)

The cross section depends on the single scaling variable [4, 5]

$$\eta \equiv \frac{Q^2 + m_0^2}{\Lambda_{\text{sat}}^2(W^2)},\tag{4}$$

and at HERA, $\Lambda_{\rm sat}^2(W^2)$ varies in the range of $2~{\rm GeV^2} \lesssim \Lambda_{\rm sat}^2(W^2) \lesssim 7~{\rm GeV^2}$. In (3), $\sigma^{(\infty)}=27.5$ mb and $R_{e^+e^-}=3\sum_f Q_f^2=10/3$. The cross section (3), at any fixed $Q^2\gg \Lambda_{\rm sat}^2(W^2)$ contains the spectacular strong rise of the structure function

$$F_2(x, Q^2) = \frac{Q^2}{4\pi^2 \alpha} \sigma_{\gamma^* p}(\eta(W^2, Q^2))$$
 (5)

with energy at HERA [6]. For $\Lambda^2(W^2) \gg Q^2$, the saturation limit of weak energy dependence, as seen in photoproduction, will be reached. Except for exceedingly small Q^2 , this only happens at energies far beyond the HERA energy range.

A fundamental question concerns the magnitude of the exponent C_2 in (1) that determines the rise of the cross section at large Q^2 . It turned out [7] that an examination of the expression for the proton structure function (5) in the large- Q^2 limit in (3) in terms of the dual language of sea-quark and gluon distributions, upon applying DGLAP evolution [8], provides an answer to this very important question on the magnitude of C_2 .

The structure function in the large- Q^2 limit of (3) takes the form [7]

$$F_2(x,Q^2) = \frac{R_{e^+e^-}}{36\pi^2} \left(T(W^2) + \frac{1}{2} L(W^2) \right), \tag{6}$$

where $T(W^2)$ and $L(W^2)$ are integrals over the first moments of the gluon transverse momentum formed with the dipole cross sections in momentum space. The longitudinal part of $F_2(x, Q^2)$ in (6) is expressible in terms of the saturation scale,

$$L(W^2) = \frac{\sigma^{(\infty)}}{\pi} \Lambda_{\text{sat}}^2(W^2). \tag{7}$$

The (successful) representation of the experimental results according to (3) was based on the assumption of $T(W^2) = L(W^2)$ in (6), *i.e.* equal magnitude of transverse and longitudinal $(q\bar{q})^{J=1}$ scattering cross sections.

We turn to the sea-quark and gluon distributions corresponding to the structure function (6) and the saturation scale in (7). The structure function measures the sea-quark distribution,

$$(qq^{-})_{\text{sea}} \equiv x\Sigma(x, Q^{2}) = \frac{1}{3\pi^{2}} \left(T(W^{2}) + \frac{1}{2}L(W^{2}) \right),$$
 (8)

and the gluon distribution is related to the longitudinal part of F_2 , *i.e.* with (7)

$$\alpha_{\rm s}(Q^2)xg(x,Q^2) = \frac{1}{8\pi}L\left(W^2 = \frac{Q^2}{x}\right) = \frac{1}{8\pi}\frac{\sigma^{(\infty)}}{\pi}\Lambda_{\rm sat}^2(W^2).$$
 (9)

So far, our results have been based on a general analysis of the generic structure of the two-gluon exchange from QCD. We adopt the assumption that the sea-quark and gluon distributions in good approximation have identical dependence on the kinematic variables, in our case $W^2 \cong Q^2/x$. They are assumed to be proportional to each other,

$$x\Sigma(x,Q^2) = \frac{8}{3\pi} \left(r + \frac{1}{2}\right) \alpha_{\rm s}(Q^2) x g(x,Q^2),$$
 (10)

with $r = \text{const} \ge 0$. Substitution of the sea-quark and the gluon distribution from (8) and (9) into (10) yields

$$T(W^2) = rL(W^2) \tag{11}$$

and F_2 in (6) becomes (with (7)),

$$F_2(x, Q^2) = \frac{R_{e^+e^-}}{36\pi^2} T(W^2) \left(1 + \frac{1}{2r} \right) \cong \begin{cases} T, & (r \gg 1), \\ \frac{3}{2}T, & (r = 1), \\ L, & (r \to 0). \end{cases}$$
 (12)

The above-mentioned (successful) representation of the experimental data corresponds to r = 1.

In the range of sufficiently large Q^2 , we are concerned with, the evolution of $F_2(x, Q^2)$ is in very good approximation determined by the gluon structure function alone, [8]

$$\frac{\partial F_2(\frac{x}{2}, Q^2)}{\partial \ln Q^2} = \frac{R_{e^+e^-}}{9\pi} \alpha_s(Q^2) x g(x, Q^2). \tag{13}$$

Upon substitution of (12) and (9), with (11), the evolution in Q^2 is converted into a derivative with respect to W^2 that determines the W^2 dependence

$$(2r+1)\frac{\partial}{\partial \ln W^2} \Lambda_{\text{sat}}^2(2W^2) = \Lambda_{\text{sat}}^2(W^2). \tag{14}$$

Substituting the power law (1), into (14), we obtain

$$(2r+1)2^{C_2}C_2 = 1, (15)$$

or

$$C_2^{\text{theor}} = \begin{cases} 0, & (r \gg 1), \\ 0.276, & (r = 1), \\ 0.65, & (r = 0). \end{cases}$$
 (16)

According to (16), the exponent C_2 determining the rise of $\sigma_{\gamma^*p}(W^2,Q^2)$ at fixed $Q^2 \gg \Lambda_{\rm sat}^2(W^2)$, or, equivalently, the rise of $F_2(x,Q^2)$ as a function of $x = Q^2/W^2$ at fixed Q^2 , is uniquely connected with the relative magnitude of sea *versus* gluon distributions. For a given sea distribution, a very small gluon distribution ($r \gg 1$ according to (10)) and a purely transverse structure function (according to (12)) implies a very weak energy dependence. In contrast, in the limit of $r \to 0$, where the relative magnitude of gluon-to sea distribution is maximal, we must have an extraordinarily strong energy dependence, $C_2 = 0.65$. The previous fit to the experimental data was based on r = 1. The result of the fit given in (2) is in excellent agreement with the theoretical result (16).

The foregoing intimate connection between a strong rise of $F_2(x, Q^2)$ with decreasing x at fixed Q^2 and a large longitudinal photoabsorption cross section, and a weak rise for a small longitudinal contribution, can be further illuminated by measurements allowing to separate longitudinal and transverse cross sections. Such measurements are in fact urgently needed.

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