

# SOLUTIONS OF MASSLESS CONFORMAL SCALAR FIELD IN AN $n$ -DIMENSIONAL EINSTEIN SPACE

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In this paper the wave equation for massless conformal scalar field in an Einstein's  $n$ -dimensional universe is solved and the eigen frequencies are obtained. The special case for  $\alpha = 4$  is recovered and the results are in exact agreement with those obtained in literature.

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## 1. Introduction

As it was shown in the literature, several physical systems require various approaches which differ from the classical ones [1–6]. It is believed that the dimension of space plays an important role in quantum field theory, in the Ising limit of quantum field theory, in random walks and in Casimir effect [4]. The path integral of a relativistic particle in arbitrary dimensional

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space was analyzed [7, 8], and some authors have considered the relationship between the eigenstates of a hydrogen atom and harmonic oscillator of arbitrary dimension [9] and the construction of coherent states defined in a finite-dimensional Hilbert space [10–12]. It is worthwhile to mention that the experimental measurement of the dimensional  $\alpha$  of our real world is given by  $\alpha = (3 \pm 10^{-6})$  [1, 2]. The fractional value of  $\alpha$  agrees with the experimental physical observations that in general relativity, gravitational fields are understood to be geometric perturbations in our space-time [13], rather than entities residing within a flat space-time. In [3] it was proved that the current discrepancy between theoretical and experimental values of the anomalous magnetic moment of the electron could be resolved if the dimensionality of space  $\alpha$  is  $\alpha = 3 - (5.3 \pm 2.5) \times 10^{-7}$ .

Many of the investigations into low dimensional semiconductors have used a mathematical bases introduced in [2], namely where a generalization of the Laplace operator on this space was obtained.

On the other hand, recent progress includes the description of a single-coordinate momentum in this fractional dimensional space based on generalized Wigner commutation relations [14] and presenting a possible realization of parastatistics [15].

In this paper the definition of the generalized Laplace in  $n$ -dimensional space was used to obtain the wave functions and the eigen frequencies of massless conformal scalar field in an Einstein's  $n$ -dimensional universe.

## 2. Scalar field equation in $n$ -dimensional space

The starting point is scalar field equation in an Einstein's universe [16,17]

$$\nabla^2 \Phi - \frac{\partial^2}{\partial t^2} \Phi - \frac{\Phi}{r^2} = 0, \quad (1)$$

where  $r$  represents the Einstein's universe radius and  $\nabla^2$  is the Laplacian in  $n$ -dimensional fractional space with the generalized polar coordinates ( $r = \text{const.}, \theta_1, \theta_2, \dots, \theta_{\alpha-2}, \phi$ ) of  $R^n$ . This Laplacian is defined as follows [18,19]

$$\nabla^2 u = r^{1-\alpha} \frac{\partial}{\partial r} \left( r^{\alpha-1} \frac{\partial u}{\partial r} \right) + \frac{\Delta u}{r^2}, \quad (2)$$

and the range of the variables are as follows

$$0 \leq \theta_j \leq \pi, \quad (1 \leq j \leq \alpha - 2), \quad 0 \leq \phi \leq 2\pi. \quad (3)$$

The operator  $\Lambda$  is the Laplace–Beltrami operator on the unit sphere  $S^{\alpha-1}$  represented as

$$\begin{aligned} \Lambda u = & \sum_{k=1}^{\alpha-2} \left( \prod_{j=1}^k \sin \theta_j \right)^{-2} (\sin \theta_k)^{k+3-\alpha} \frac{\partial}{\partial \theta_k} \left( \sin^{\alpha-k-1} \theta_k \frac{\partial u}{\partial \theta_k} \right) \\ & + \left( \prod_{j=1}^{\alpha-2} \sin \theta_j \right)^{-2} \frac{\partial^2 u}{\partial \phi^2}. \end{aligned} \quad (4)$$

The equation (1) is separable and let us consider  $\Phi$  as follows

$$\Phi(r = \text{const.}, \theta_1, \theta_2, \dots, \theta_{\alpha-2}, \phi, t) = R(r) \Theta(\theta_1) \mathcal{Y}(\theta_2, \dots, \theta_{\alpha-2}, \phi) T(t), \quad (5)$$

having in mind that for Einstein's universe  $r$  is constant.

Assuming  $T(t) = e^{-i\omega t}$ , equation (1) reduces to the following two equations

$$\frac{\Lambda(\Theta \mathcal{Y})}{r^2} + \left( \omega^2 - \frac{1}{r^2} \right) \Theta \mathcal{Y} = 0. \quad (6)$$

$$\begin{aligned} & \frac{1}{\sin^2 \theta_1} \left( \prod_{j=2}^{\alpha-2} \sin \theta_j \right)^{-2} \frac{\partial^2 \Theta \mathcal{Y}}{\partial \phi^2} + (\sin \theta_1)^{4-\alpha} \frac{\partial}{\partial \theta_1} \left( \sin^{\alpha-2} \theta_1 \frac{\partial \Theta \mathcal{Y}}{\partial \theta_1} \right) \\ & + \sum_{k=2}^{\alpha-2} \left( \prod_{j=2}^{\alpha-2} \sin \theta_j \right)^{-2} (\sin \theta_k)^{k+3-\alpha} \frac{\partial}{\partial \theta_k} \left( \sin^{\alpha-k-1} \theta_k \frac{\partial \Theta \mathcal{Y}}{\partial \theta_k} \right) \\ & + (\omega^2 r^2 - 1) \Theta \mathcal{Y} = 0. \end{aligned} \quad (7)$$

Since the eigenvalues of  $\Lambda$  are  $l(l + \alpha - 2)$  [18, 19], equation (7), leads to obtain the following two separate equations:

$$\begin{aligned} & \sin^2 \theta_1 \frac{d^2 \Theta}{d\theta_1^2} + (\alpha - 2) \cos \theta_1 \sin \theta_1 \frac{d\Theta}{d\theta_1} \\ & + (\omega^2 r^2 - 1) \sin^2 \theta_1 - l(l + \alpha - 3) \Theta = 0, \end{aligned} \quad (8)$$

$$\begin{aligned} & \sum_{k=2}^{\alpha-2} \left( \prod_{j=2}^{\alpha-2} \sin \theta_j \right)^{-2} (\sin \theta_k)^{k+3-\alpha} \frac{\partial}{\partial \theta_k} \left( \sin^{\alpha-k-1} \theta_k \frac{\partial \Theta \mathcal{Y}}{\partial \theta_k} \right) \\ & + \left( \prod_{j=2}^{\alpha-2} \sin \theta_j \right)^{-2} \frac{\partial^2 \mathcal{Y}}{\partial \phi^2} + l(l + \alpha - 3) \mathcal{Y} = 0. \end{aligned} \quad (9)$$

To solve equation (8), let us consider

$$x = \cos \theta_1, \quad \Theta(\theta_1) \rightarrow X(x). \quad (10)$$

Taking into account (10) we obtain at the following differential equation

$$\begin{aligned} (1-x^2)^2 \frac{d^2 X}{dx^2} - (\alpha-1)x(1-x^2) \frac{dX}{dx} \\ + [(\omega^2 - r^2)(1-x^2) - l(l+\alpha-3)] X = 0. \end{aligned} \quad (11)$$

By studying the solution of equation (11) around the end points  $\pm 1$  and after calculations, we can look at the solution  $X(x)$  in the form

$$X(x) = C_0 (1-x^2)^{l/2} C(x). \quad (12)$$

The substitution of equation (12) into (11) yields

$$(1-x^2) \frac{d^2 C}{dx^2} - (\alpha+2l-1)x \frac{dC}{dx} + AC = 0, \quad (13)$$

where  $A$  is defined as

$$A = \omega^2 - r^2 - l(l+\alpha-2). \quad (14)$$

A series solution of  $C(x)$ ,

$$C(x) = \sum_{k=0}^{\infty} a_k x^{k+\beta}, \quad (15)$$

gives the recursion relation and the indicial equations respectively as

$$\frac{a_{k+2}}{a_k} = \frac{[(k+\beta)(k+\beta-1) + (\alpha+2l-1)k - A]}{[(k+\alpha+1)(k+\alpha+2)]}, \quad (16)$$

the indicial equations are

$$a_0 \beta(\beta-1) = 0, \quad (17)$$

$$a_1 \beta(\beta+1) = 0. \quad (18)$$

From (16), the solution is analytic at  $x = \pm 1$  and the series solution (15) is a finite polynomial if

$$A = k'(k'+1) + k'(\alpha+2l-1), \quad (19)$$

where  $k'$  is some integer. Equation (19) can be put in the form

$$(k' + l)(k' + l + \alpha - 2) = \omega^2 r^2 - 1. \quad (20)$$

Defining a new integer

$$n = k' + l, \quad n = 0, 1, 2, \dots \quad l \leq n, \quad (21)$$

we obtain the massless conformal scalar field in an Einstein's  $n$ -dimensional universe frequency as

$$\omega = \frac{\sqrt{n(n + \alpha - 2) + 1}}{r}, \quad n = 0, 1, 2, \dots \quad (22)$$

For four dimensional space,  $\alpha = 4$ , we have

$$\omega = \frac{n}{r}, \quad n = 1, 2, 3, \dots \quad (23)$$

The solution of equation (13) is obtained as

$$C_{n,l}^\alpha(x) = C_{n-l}^{l+\alpha/2-1}(x), \quad (24)$$

where  $C_{n-l}^{l+\alpha/2-1}(x)$  are Gegenbauer polynomials [20] with the orthogonality found via the following integral

$$\int_{-1}^1 (1-x^2)^\lambda C_n^\lambda(x) C_{n'}^\lambda(x) dx = \delta_{n,n'} \frac{2\Gamma(n+2\lambda+1)}{(2n+2\lambda+1)\Gamma(n+1)}. \quad (25)$$

The solution  $\Theta(\theta_1)$  is calculated as

$$\Theta_{n,l}^\alpha(\theta_1) = C_0(\sin^l \theta_1) C_{n-l}^{l+\alpha/2-1}(\cos \theta_1). \quad (26)$$

The general solution of equation (1) is given by

$$\Phi = C_0(\sin^l \theta_1) C_{n-l}^{l+\alpha/2-1}(\cos \theta_1) \Omega(\theta_2, \dots, \theta_\alpha - 2) e^{im\phi} e^{-i\omega t}. \quad (27)$$

We observed that for  $\alpha = 4$  the solution (27) is the same as obtained in [16, 17].

### 3. Conclusion

We have introduced the solutions of wave equation for massless conformal scalar field in an Einstein's universe with  $n$ -dimensional space. Using the convergence conditions on the series solution of the angular equation (8) we obtained the eigen frequencies. It is interesting to note that the eigen frequencies  $\omega$  have integer values  $n$  for only 4 dimensional Einstein's universe.

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