# MOMENTS OF WISHART-LAGUERRE AND JACOBI ENSEMBLES OF RANDOM MATRICES: APPLICATION TO THE QUANTUM TRANSPORT PROBLEM IN CHAOTIC CAVITIES 

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We collect explicit and user-friendly expressions for one-point densities of the real eigenvalues $\left\{\lambda_{i}\right\}$ of $N \times N$ Wishart-Laguerre and Jacobi random matrices with orthogonal, unitary and symplectic symmetry. Using these formulae, we compute integer moments $\tau_{n}=\left\langle\sum_{i=1}^{N} \lambda_{i}^{n}\right\rangle$ for all symmetry classes without any large $N$ approximation. In particular, our results provide exact expressions for moments of transmission eigenvalues in chaotic cavities with time-reversal or spin-flip symmetry and supporting a finite and arbitrary number of electronic channels in the two incoming leads.

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## 1. Introduction

We consider the Wishart-Laguerre (L) and Jacobi (J) ensembles of random matrix theory characterized by the following joint probability density (j.p.d.) of real eigenvalues $[1,2,3]$ :

$$
\begin{equation*}
\mathcal{P}_{\beta}^{(\mathrm{L})}\left(\lambda_{1}, \ldots, \lambda_{N}\right):=C_{N, \beta, \nu}^{(\mathrm{L})} \prod_{j<k}\left|\lambda_{j}-\lambda_{k}\right|^{\beta} \prod_{j=1}^{N} \lambda_{j}^{\frac{\beta}{2}(\nu+1)-1} e^{-\frac{1}{2} \lambda_{j}}, \tag{1}
\end{equation*}
$$

$$
\begin{equation*}
\mathcal{P}_{\beta}^{(\mathrm{J})}\left(\lambda_{1}, \ldots, \lambda_{N}\right):=C_{N, \beta, a, b}^{(\mathrm{J})} \prod_{j<k}\left|\lambda_{j}-\lambda_{k}\right|^{\beta} \prod_{j=1}^{N}\left(1-\lambda_{j}\right)^{a}\left(1+\lambda_{j}\right)^{b} \tag{2}
\end{equation*}
$$

In the above equations, $C_{N, \beta, \nu}^{(\mathrm{L})}$ and $C_{N, \beta, a, b}^{(\mathrm{J})}$ are normalization constants, while the index $\beta=1,2,4$ characterizes the symmetry class of the ensemble (orthogonal, unitary and symplectic respectively).

The Wishart-Laguerre ensemble contains covariance matrices of the form $\mathcal{W}=\mathcal{X} \mathcal{X}^{\dagger}$, where $\mathcal{X}$ is a $N \times M(M-N:=\nu \geq 0)$ matrix with i.i.d. Gaussian entries (real, complex or quaternionic variables). The matrix $\mathcal{W}$ is symmetric and positive semidefinite, so its $N$ real eigenvalues are nonnegative $\left(\left\{\lambda_{j}\right\} \geq 0\right)$. Originally introduced by Wishart [1], matrices from this ensemble have been extensively used in multivariate statistical data analysis $[4,5]$ with applications in various fields ranging from meteorological data [6] to finance [7, 8]. They are also useful when analyzing the capacity of channels with multiple antennae and receivers [9], in nuclear physics [10], chiral quantum chromodynamics [11] and also in statistical physics such as in a class of $(1+1)$-dimensional directed polymer problems [12]. Recently, they have also appeared in the context of knowledge networks [13] and new mathematical results for the case $\nu<0$ have also been lately obtained $[14,15]$. Large deviation properties of the eigenvalues have been investigated in $[16,17,18]$, while for an excellent review we refer to [19].

The Jacobi ensemble contains combinations of two $N \times N$ WishartLaguerre matrices $\mathcal{W}_{1}$ and $\mathcal{W}_{2}$ of the form:

$$
\begin{equation*}
\mathcal{J}=\left(\mathcal{W}_{1}-\mathcal{W}_{2}\right)\left(\mathcal{W}_{1}+\mathcal{W}_{2}\right)^{-1} \tag{3}
\end{equation*}
$$

and its eigenvalues are real and lie on the support $-1 \leq\left\{\lambda_{j}\right\} \leq 1$. Matrices distributed according to the Jacobi weight arise as (i) truncations of Haar orthogonal, unitary or symplectic matrices (for the case of unitary matrices, an important application arises in the theory of electronic transport in mesoscopic systems at low temperatures as detailed in Appendix A); (ii) as composition of projection matrices [20].

In order to proceed, we first define a shifted version of the Jacobi ensemble with eigenvalues between 0 and 1 :

$$
\begin{equation*}
\mathcal{P}_{\beta}^{(\mathrm{sJ})}\left(\lambda_{1}, \ldots, \lambda_{N}\right):=C_{N, \beta, \mathfrak{a}, \mathfrak{b}}^{(\mathrm{sJ})} \prod_{j<k}\left|\lambda_{j}-\lambda_{k}\right|^{\beta} \prod_{j=1}^{N} \lambda_{j}^{\mathfrak{a}}\left(1-\lambda_{j}\right)^{\mathfrak{b}} \tag{4}
\end{equation*}
$$

which appears more frequently in physical applications (see Appendix A) and numerical algorithms $[21,22]$. By changing variables $\lambda_{j}=1-2 y_{j}$ in (2),
it is easy to see that

$$
\begin{equation*}
C_{N, \beta, a, b}^{(\mathrm{J})}=\frac{C_{N, \beta, a, b}^{(\mathrm{sJ})}}{2^{N(a+b+1)+\frac{\beta}{2} N(N-1)}} \tag{5}
\end{equation*}
$$

We also define the average spectral densities $\left(\rho(x)=\left\langle\frac{1}{N} \sum_{i=1}^{N} \delta\left(x-\lambda_{i}\right)\right\rangle\right)$ for the three ensembles above as the marginals of their respective j.p.ds.:

$$
\begin{align*}
\rho_{N, \beta, \nu}^{(\mathrm{L})}\left(\lambda_{1}\right)= & \int_{[0, \infty]^{N-1}} d \lambda_{2} \cdots d \lambda_{N} \mathcal{P}_{\beta}^{(\mathrm{L})}\left(\lambda_{1}, \ldots, \lambda_{N}\right),  \tag{6}\\
\rho_{N, \beta, a, b}^{(\mathrm{J})}\left(\lambda_{1}\right)= & \int_{[-1,1]^{N-1}} d \lambda_{2} \cdots d \lambda_{N} \mathcal{P}_{\beta}^{(\mathrm{J})}\left(\lambda_{1}, \ldots, \lambda_{N}\right),  \tag{7}\\
\rho_{N, \beta, \mathfrak{a}, \mathfrak{b}}^{(\mathrm{sJ})}\left(\lambda_{1}\right)= & \int_{[0,1]^{N-1}} d \lambda_{2} \cdots d \lambda_{N} \mathcal{P}_{\beta}^{(\mathrm{sJ})}\left(\lambda_{1}, \ldots, \lambda_{N}\right) \tag{8}
\end{align*}
$$

It follows immediately from the definition that the above densities are all normalized to 1 , and that the following relation holds between the Jacobi and the shifted-Jacobi densities:

$$
\begin{equation*}
\rho_{N, \beta, a, b}^{(\mathrm{J})}(1-2 x)=\frac{1}{2} \rho_{N, \beta, a, b}^{(\mathrm{sJ})}(x) . \tag{9}
\end{equation*}
$$

The purpose of this paper is twofold:

- to collect and present in a user-friendly way (well-known, but somehow scattered throughout the literature) explicit formulae for the above densities for finite $N$ and all symmetry classes;
- to use these formulae to compute integer moments of the eigenvalues $\tau_{n}=\left\langle\sum_{i=1}^{N} \lambda_{i}^{n}\right\rangle$ (where the average is taken w.r.t. any of the three j.p.ds. above). More precisely, we define:

$$
\begin{align*}
\tau_{n}^{(\mathrm{L})}(N, \beta, \nu) & :=\int_{[0, \infty]^{N}} d \lambda_{1} \cdots d \lambda_{N} \mathcal{P}_{\beta}^{(\mathrm{L})}\left(\lambda_{1}, \ldots, \lambda_{N}\right)\left(\sum_{i=1}^{N} \lambda_{i}^{n}\right),  \tag{10}\\
\tau_{n}^{(\mathrm{J})}(N, \beta, a, b) & :=\int_{[-1,1]^{N}} d \lambda_{1} \cdots d \lambda_{N} \mathcal{P}_{\beta}^{(\mathrm{J})}\left(\lambda_{1}, \ldots, \lambda_{N}\right)\left(\sum_{i=1}^{N} \lambda_{i}^{n}\right),  \tag{11}\\
\tau_{n}^{(\mathrm{sJ})}(N, \beta, \mathfrak{a}, \mathfrak{b}) & :=\int_{[0,1]^{N}} d \lambda_{1} \cdots d \lambda_{N} \mathcal{P}_{\beta}^{(\mathrm{sJ})}\left(\lambda_{1}, \ldots, \lambda_{N}\right)\left(\sum_{i=1}^{N} \lambda_{i}^{n}\right) \tag{12}
\end{align*}
$$

One application to the case of quantum transport in chaotic cavities is detailed in Appendix A. Other interesting mathematical results for the Laguerre case can be found e.g. in [23] and references therein. We are also aware that formulae for the Wishart-Laguerre and Jacobi moments for $\beta=1,2,4$ and finite $N$ have been derived by Mezzadri and Simm [24]. These are different from the ones provided here and obtained via a different method, but equivalent.

It is easy to see that the average of any linear statistics (i.e. a quantity of the form $\left.\mathcal{A}=\sum_{i=1}^{N} f\left(\lambda_{i}\right)\right)$

$$
\begin{equation*}
\langle\mathcal{A}\rangle=\int d \lambda_{1} \cdots d \lambda_{N} \mathcal{P}_{\beta}\left(\lambda_{1}, \ldots, \lambda_{N}\right)\left(\sum_{i=1}^{N} f\left(\lambda_{i}\right)\right) \tag{13}
\end{equation*}
$$

can be computed as a one-dimensional integral over the corresponding spectral density of the ensemble as:

$$
\begin{equation*}
\left\langle\sum_{i=1}^{N} f\left(\lambda_{i}\right)\right\rangle=N \int d x \rho(x) f(x) \tag{14}
\end{equation*}
$$

The technical achievement we report in this paper is an explicit computation of this integral valid for finite matrix dimension $N$ and all three $\beta$ s for the case $f(x)=x^{n}, n \in \mathbb{N}$. The integral (14) also elucidates a possible strategy to evaluate a regular asymptotic expansion of moments for large $N$, which has been recently highlighted [25] as a problem of current interest in the context of electronic transport in chaotic cavities (see Appendix A): all one has to do is to seek for a regular $(1 / N)$ expansion of the macroscopic spectral density of the form:

$$
\begin{equation*}
\rho(x)=\rho^{(\infty)}(x)+\frac{1}{N} \rho^{(1)}(x)+\frac{1}{N^{2}} \rho^{(2)}(x)+\ldots \tag{15}
\end{equation*}
$$

in the spirit of high genus correlator expansions [26], and then integrate term by term. While this program is far from completion and is thus left for future work, we show in Appendix B that at least the leading order of the asymptotic expansion of moments is well reproduced for the case of quantum transport in cavities with broken time-reversal symmetry $(\beta=2)$.

The plan of this paper is as follows: in Section 2 we consider the WishartLaguerre ensemble. We first summarize the densities for all three $\beta \mathrm{s}$, then use the integral formula (14) to compute integer moments for $\beta=1,2,4$ in the three subsections. Then in the last subsection we collect results from numerical simulations for the density and moments. The same thing is done for the Jacobi ensemble in Section 3. A summary and outlook is provided
in Section 4. In Appendix A we provide a detailed introduction to the problem of quantum transport in chaotic cavities which constitutes the main motivation for this study, while in Appendix B we show that the leading order term in the expansion of the moments via spectral density is correctly reproduced.

## 2. Wishart-Laguerre ensemble

### 2.1. Spectral densities

The spectral density for the Laguerre ensemble is known for all three symmetry classes $[27,28,29,30]$ and, after tedious algebraic manipulations can be cast in the form $\rho_{N, \beta, \nu}^{(\mathrm{L})}(x)=\frac{1}{2 N} \mathcal{R}_{N, \beta, \nu}^{(\mathrm{L})}(x / 2)$, where

$$
\begin{align*}
\mathcal{R}_{N, 1, \nu}^{(\mathrm{L})}(x)= & 2 \mathcal{R}_{N, 2, \nu}^{(\mathrm{L})}(2 x)-\frac{\Gamma((N+1) / 2)}{\Gamma((N+\nu) / 2)} L_{N-1}^{(\nu)}(2 x)\left\{\phi_{1}(x)-\phi_{2}(x)\right\},  \tag{16}\\
\mathcal{R}_{N, 2, \nu}^{(\mathrm{L})}(x)= & x^{\nu} e^{-x} \sum_{m=0}^{N-1} \frac{\Gamma(m+1)}{\Gamma(m+\nu+1)}\left(L_{m}^{(\nu)}(x)\right)^{2}  \tag{17}\\
\mathcal{R}_{N, 4, \nu}^{(\mathrm{L})}(x)= & \frac{1}{2} \mathcal{R}_{2 N, 2,2 \nu}^{(\mathrm{L})}(x)-x^{2 \nu} e^{-x} L_{2 N}^{(2 \nu)}(x) \frac{\Gamma(N+1)}{2^{2 \nu+1} \Gamma(N+\nu+1 / 2)} \\
& \times \sum_{m=0}^{N-1} \frac{\Gamma(m+1 / 2)}{\Gamma(m+\nu+1)} L_{2 m}^{(2 \nu)}(x) \tag{18}
\end{align*}
$$

and

$$
\begin{align*}
\phi_{1}(x) & =(2 x)^{\nu} e^{-2 x} \sum_{m=0}^{(\kappa+N-2) / 2} \mathfrak{d}_{m} L_{2 m+1-\kappa}^{(\nu)}(2 x)  \tag{19}\\
\phi_{2}(x) & =x^{(\nu-1) / 2} e^{-x}\left[(1-\kappa) \frac{2 \Gamma((\nu+1) / 2, x)}{\Gamma((\nu+1) / 2)}+2 \kappa-1\right]  \tag{20}\\
\mathfrak{d}_{m} & =\frac{\Gamma(m+1-\kappa / 2)}{2^{\nu-1} \Gamma(m+(\nu-1) / 2+2-\kappa / 2)} \tag{21}
\end{align*}
$$

In the above equations, $\kappa=N \bmod 2, L_{N}^{(\nu)}(x)$ is a generalized Laguerre polynomial defined by the sum:

$$
\begin{equation*}
L_{n}^{(\lambda)}(z)=\sum_{k=0}^{n} c_{k}(n, \lambda) z^{k} \tag{22}
\end{equation*}
$$

where

$$
\begin{equation*}
c_{k}(n, \lambda)=\frac{\Gamma(\lambda+n+1)(-n)_{k}}{n!k!\Gamma(\lambda+k+1)} \tag{23}
\end{equation*}
$$

and $\Gamma(a, x)=\int_{x}^{\infty} t^{a-1} e^{-t} d t$ is the incomplete Gamma function, while $(x)_{n}=$ $\Gamma(x+n) / \Gamma(x)$ is the Pochhammer symbol.

In the following subsections, we present the results for integer moments $\tau_{n}^{(\mathrm{L})}(N, \beta, \nu)=\left\langle\sum_{j=1}^{N} \lambda_{j}^{n}\right\rangle$ obtained by integration (Eq. (14)) of the densities given above. The task is most easily accomplished by first defining the following auxiliary function:

$$
\begin{equation*}
\mathcal{Q}(r ; m, \ell ; \alpha):=\int_{0}^{\infty} d x x^{r} e^{-x} L_{m}^{(\alpha)}(x) L_{\ell}^{(\alpha)}(x) \tag{24}
\end{equation*}
$$

which is easily evaluated using (22) as

$$
\begin{equation*}
\mathcal{Q}(r ; m, \ell ; \alpha)=\sum_{k=0}^{m} \sum_{k^{\prime}=0}^{\ell} c_{k}(m, \alpha) c_{k^{\prime}}(\ell, \alpha) \Gamma\left(1+r+k+k^{\prime}\right) \tag{25}
\end{equation*}
$$

Note that, by the orthogonality relation of Laguerre polynomials, one has

$$
\begin{equation*}
\mathcal{Q}(\alpha ; m, \ell ; \alpha)=\frac{(\ell+\alpha)!}{\ell!} \delta_{m \ell} \tag{26}
\end{equation*}
$$

### 2.2. Moments $\beta=1$

For $\beta=1$, the final result of the integration reads as follows:

$$
\begin{equation*}
\tau_{n}^{(\mathrm{L})}(N, 1, \nu)=2^{-n} \tau_{n}^{(\mathrm{L})}(N, 2, \nu)+\mathcal{I}_{2}(n)+\mathcal{I}_{3}(n) \tag{27}
\end{equation*}
$$

where:

$$
\begin{align*}
\mathcal{I}_{2}(n)= & -\frac{\Gamma((N+1) / 2)}{2 \Gamma((N+\nu) / 2)} \sum_{m=0}^{(\kappa+N-2) / 2} \mathfrak{d}_{m} \mathcal{Q}(n+\nu ; N-1,2 m+1-\kappa ; \nu)  \tag{28}\\
\mathcal{I}_{3}(n)= & \frac{\Gamma((N+1) / 2)}{2 \Gamma((N+\nu) / 2)}\left\{(1-\kappa)\left[\frac{2^{(3-\nu) / 2}}{\Gamma((\nu+1) / 2)} \mathcal{Y}_{1}(n)-2^{(1-\nu) / 2} \mathcal{Y}_{2}(n)\right]\right. \\
& \left.+\kappa 2^{(1-\nu) / 2} \mathcal{Y}_{2}(n)\right\}  \tag{29}\\
\mathcal{Y}_{1}(n)= & \sum_{m=0}^{N-1} \frac{c_{m}(N-1, \nu) 2^{n+(\nu+1) / 2+m} \Gamma(n+\nu+1+m)}{n+(\nu+1) / 2+m} \\
& \times{ }_{2} F_{1}(n+(\nu+1) / 2+m, n+\nu+1+m ; n+(\nu+3) / 2+m ;-1)  \tag{30}\\
\mathcal{Y}_{2}(n)= & \sum_{m=0}^{N-1} c_{m}(N-1, \nu) 2^{n+(\nu+1) / 2+m} \Gamma(n+m+(\nu+1) / 2) \tag{31}
\end{align*}
$$

In (30), we have used the following hyper geometric function:

$$
\begin{equation*}
{ }_{2} F_{1}\left(a_{1}, a_{2} ; b_{1} ; z\right):=\sum_{k=0}^{\infty} \frac{\left(a_{1}\right)_{k}\left(a_{2}\right)_{k}}{\left(b_{1}\right)_{k} k!} z^{k} \tag{32}
\end{equation*}
$$

One can check by direct inspection that $\tau_{0}^{(\mathrm{L})}(N, 1, \nu)=N$ and $\tau_{1}^{(\mathrm{L})}(N, 1, \nu)=$ $N(N+\nu)$ as it should be.

### 2.3. Moments $\beta=2$

Combining (14) with (17), one easily obtains

$$
\begin{equation*}
\tau_{n}^{(\mathrm{L})}(N, 2, \nu)=2^{n} \sum_{m=0}^{N-1} \frac{\Gamma(m+1)}{\Gamma(m+\nu+1)} \mathcal{Q}(n+\nu ; m, m ; \nu) \tag{33}
\end{equation*}
$$

One can check by direct inspection that $\tau_{0}^{(\mathrm{L})}(N, 2, \nu)=N$ and $\tau_{1}^{(\mathrm{L})}(N, 2, \nu)=$ $2 N(N+\nu)$ as it should be.

$$
\text { 2.4. Moments } \beta=4
$$

Similarly, for $\beta=4$ one gets

$$
\begin{equation*}
\tau_{n}^{(\mathrm{L})}(N, 4, \nu)=\frac{1}{2} \tau_{n}^{(\mathrm{L})}(2 N, 2,2 \nu)-\sum_{m=0}^{N-1} \sum_{k=0}^{2 N} \sum_{k^{\prime}=0}^{2 m} \chi_{m, k, k^{\prime}} \tag{34}
\end{equation*}
$$

where:

$$
\begin{align*}
\chi_{m, k, k^{\prime}} & =f_{m}\left(N, \nu, n, k, k^{\prime}\right) c_{k}(2 N, 2 \nu) c_{k^{\prime}}(2 m, 2 \nu)  \tag{35}\\
f_{m}\left(N, \nu, n, k, k^{\prime}\right) & =\frac{\Gamma\left(1+k+k^{\prime}+n+2 \nu\right) \Gamma(N+1) \Gamma(m+1 / 2)}{2^{2 \nu+1-n} \Gamma(N+\nu+1 / 2) \Gamma(m+\nu+1)} \tag{36}
\end{align*}
$$

One can check by direct inspection that $\tau_{0}^{(\mathrm{L})}(N, 4, \nu)=N$ and $\tau_{1}^{(\mathrm{L})}(N, 4, \nu)=$ $4 N(N+\nu)$ as it should be.

### 2.5. Comparison with numerics

In Fig. 1 we plot the analytical formulae (17), (16) and (18) together with numerical diagonalization of matrices $\mathcal{W}$ from the Laguerre ensemble
with $\beta=1,2,4$ respectively ${ }^{1}$, obtained in Matlab as follows ${ }^{2}$ :

$$
\begin{array}{ll}
\beta=1 & \mathrm{X}=\mathrm{randn}(\mathrm{~N}, \mathrm{M}) ; \mathrm{W}=\mathrm{X} * \mathrm{X}^{\prime} ; \\
\beta=2 & \mathrm{X}=\mathrm{randn}(\mathrm{~N}, \mathrm{M})+\mathrm{i} * \operatorname{randn}(\mathrm{~N}, \mathrm{M}) ; \mathrm{W}=\mathrm{X} * \mathrm{X}^{\prime} ; \\
\beta=4 & \mathrm{X}=\operatorname{randn}(\mathrm{N}, \mathrm{M})+\mathrm{i} * \operatorname{randn}(\mathrm{~N}, \mathrm{M}) ; \\
& \mathrm{Y}=\operatorname{randn}(\mathrm{N}, \mathrm{M})+\mathrm{i} * \operatorname{randn}(\mathrm{~N}, \mathrm{M}) ; \\
& \mathrm{A}=[\mathrm{X} \mathrm{Y} ;-\operatorname{conj}(\mathrm{Y}) \operatorname{conj}(\mathrm{X})] ; \mathrm{W}=\mathrm{A} * \mathrm{~A}^{\prime} ;
\end{array}
$$





Fig. 1. Average spectral density of the Laguerre ensemble for $N=8, \nu=7$ and $\beta=1,2,4$ (top to bottom). Numerical diagonalization is given in (red) dots, while theoretical results are in solid (black).
${ }^{1}$ Note that for $\beta=4$ the matrices corresponding to a certain $N$ actually have size $2 N \times 2 N$ and thus have $2 N$ real and positive eigenvalues. Only $N$ of them are distinct though, and only those must be used when carrying out numerical simulations for moments.
${ }^{2}$ Alternatively, one can use the tridiagonal algorithm by Dumitriu and Edelman [21].

In Table I below, we compare the analytical results for moments of the Laguerre ensemble with numerical simulations via the Matlab algorithm given above.

TABLE I
Comparison between theory and numerics for moments of the Laguerre ensemble $(N=5, M=7 \rightarrow \nu=2)$ (the numerical results are obtained averaging over $\mathcal{O}\left(10^{5}\right)$ samples $)$.

| $\beta$ | $n$ | Theory | Numerics |
| :---: | :---: | :---: | :---: |
| 1 | 1 | 35 | 34.994 |
| 1 | 2 | 455 | 455.27311 |
| 1 | 3 | 7665 | 7660.026 |
| 2 | 1 | 70 | 69.981 |
| 2 | 2 | 1680 | 1680.238 |
| 2 | 3 | 50400 | 50409.428 |
| 4 | 1 | 140 | 139.982 |
| 4 | 2 | 6440 | 6437.074 |
| 4 | 3 | 362880 | 362857.134 |

## 3. Jacobi ensemble

### 3.1. Spectral densities

The eigenvalue densities for the Jacobi Ensemble $(x \in[-1,1])$ read $[29,30]^{3}$
$\rho_{N, 1, a, b}^{(\mathrm{J})}(x)=\frac{1}{N} \sum_{m=0}^{N / 2-1}\left(g_{2 m}\right)^{-1}\left[\phi_{2 m}(x) \psi_{2 m+1}(x)-\phi_{2 m+1}(x) \psi_{2 m}(x)\right]$,
$\rho_{N, 2, a, b}^{(\mathrm{J})}(x)=\frac{w_{a, b}(x)}{N} \sum_{n=0}^{N-1}\left(h_{n}^{a, b}\right)^{-1}\left[P_{n}^{(a, b)}(x)\right]^{2}$,
$\rho_{N, 4, a, b}^{(\mathrm{J})}(x)=\rho_{2 N, 2,2 \hat{a}+1,2 \hat{b}+1}^{(\mathrm{J})}(x)-P_{2 N}^{(2 \hat{a}+1,2 \hat{b}+1)}(x) \sum_{m=0}^{N-1} K_{2 N, m}^{(\hat{a}, \hat{b})} P_{2 m}^{(2 \hat{a}+1,2 \hat{b}+1)}(x)$,
where

$$
\begin{equation*}
h_{n}^{a, b}=\frac{2^{a+b+1}}{2 n+a+b+1} \frac{\Gamma(n+a+1) \Gamma(n+b+1)}{n!\Gamma(n+a+b+1)}, \tag{40}
\end{equation*}
$$

[^0]\[

$$
\begin{align*}
g_{2 m}= & g_{2 m+1}=h_{2 m}^{2 a+1,2 b+1}  \tag{41}\\
\phi_{2 m}(x)= & w_{a, b}(x) P_{2 m}^{(2 a+1,2 b+1)}(x)  \tag{42}\\
\psi_{2 m+1}(x)= & w_{a+1, b+1}(x) P_{2 m}^{(2 a+1,2 b+1)}(x)  \tag{43}\\
\phi_{2 m+1}(x)= & w_{a, b}(x)\left[A_{2 m+1} P_{2 m+1}^{(2 a+1,2 b+1)}(x)-B_{2 m-1} P_{2 m-1}^{(2 a+1,2 b+1)}(x)\right],  \tag{44}\\
\psi_{2 m}(x)= & \frac{1}{2} \int_{-1}^{+1} d y \operatorname{sign}(x-y) \phi_{2 m}(y)  \tag{45}\\
& \times \frac{\Gamma(m+1 / 2) \Gamma(m+a+b+3 / 2)}{\Gamma(m+a+3 / 2) \Gamma(m+b+3 / 2)}, \\
K_{N, m}^{(a, b)}= & \frac{(4 m+2 a+2 b+3) \Gamma((N+2) / 2) \Gamma((N+2 a+2 b+4) / 2)}{2^{2 a+2 b+3} \Gamma((N+2 a+2) / 2) \Gamma((N+2 b+2) / 2)}  \tag{46}\\
\hat{a}= & \frac{a-2}{2},  \tag{47}\\
\hat{b}= & \frac{b-2}{2}, \tag{48}
\end{align*}
$$
\]

where $\operatorname{sign}(z)=z /|z|$ and

$$
\begin{align*}
A_{n} & =-\frac{n(n+2 a+2 b+2)}{2 n+2 a+2 b+1} \\
B_{n} & =-\frac{(n+2 a+2)(n+2 b+2)}{2 n+2 a+2 b+5}, \quad\left(B_{-1}=0\right) \tag{49}
\end{align*}
$$

In the formulas above, $w_{a, b}(x)=(1-x)^{a}(1+x)^{b}$ is the Jacobi weight function, and $P_{n}^{(a, b)}$ is the $n$-th order Jacobi polynomial with parameters $a$ and $b$, defined as:

$$
\begin{align*}
P_{n}^{(a, b)}(x) & =\sum_{j=0}^{n}\binom{n+a}{j}\binom{n+b}{n-j}\left(\frac{x-1}{2}\right)^{n-j}\left(\frac{x+1}{2}\right)^{j} \\
& =\frac{1}{2^{n}} \sum_{j=0}^{n} c_{j}^{(n)}(a, b)(x-1)^{n-j}(x+1)^{j} \tag{50}
\end{align*}
$$

where we set:

$$
\begin{equation*}
c_{j}^{(n)}(a, b)=\frac{\Gamma(n+a+1)}{\Gamma(j+1) \Gamma(n+a-j+1)} \frac{\Gamma(n+b+1)}{\Gamma(n-j+1) \Gamma(b+j+1)} . \tag{51}
\end{equation*}
$$

We now turn to the problem of computing integer moments of Jacobi matrices. It is easy to derive first the following relations between integer moments in the shifted-Jacobi and ordinary Jacobi ensembles:

$$
\begin{align*}
\tau_{n}^{(\mathrm{sJ})}(N, \beta, \mathfrak{a}, \mathfrak{b}) & =2^{-n} \sum_{k=0}^{n}\binom{n}{k}(-1)^{k} \tau_{k}^{(\mathrm{J})}(N, \beta, \mathfrak{a}, \mathfrak{b}),  \tag{52}\\
\tau_{n}^{(\mathrm{J})}(N, \beta, a, b) & =\sum_{k=0}^{n}\binom{n}{k}(-2)^{k} \tau_{k}^{(\mathrm{sJ})}(N, \beta, a, b) \tag{53}
\end{align*}
$$

In the following, we will therefore focus on the ordinary Jacobi case.

$$
\text { 3.2. Moments } \beta=1
$$

The $n$-th moment of the distribution is therefore given by

$$
\begin{equation*}
\tau_{n}^{(\mathrm{J})}(N, 1, a, b)=\sum_{m=0}^{N / 2-1}\left(g_{2 m}\right)^{-1}\left(I_{2 m, 2 m+1}^{(n)}-I_{2 m+1,2 m}^{(n)}\right) \tag{54}
\end{equation*}
$$

where

$$
\begin{equation*}
I_{j, k}^{(n)}=\int_{-1}^{+1} d x \phi_{j}(x) \psi_{k}(x) x^{n} \tag{55}
\end{equation*}
$$

are computed explicitly in (58) and (63) below. As it is clear from equation (54), we only need to compute two different kinds of integrals. Let us then start from the first kind $\left(I_{2 m, 2 m+1}^{(n)}\right)$.

Using (42), (43) and (50), we therefore get (in the following the $c_{j}^{(n)}$ coefficients will always depend on the pair $(2 a+1,2 b+1)$, so we shall omit their explicit dependence on those parameters for the rest of the subsection)

$$
\begin{align*}
I_{2 m, 2 m+1}^{(n)}= & \int_{-1}^{+1} d x \phi_{2 m}(x) \psi_{2 m+1}(x) x^{n}=\frac{1}{2^{4 m}} \sum_{i, j=0}^{2 m}(-1)^{4 m-i-j} c_{i}^{(2 m)} c_{j}^{(2 m)} \\
& \times \underbrace{\int_{-1}^{+1} d x(1-x)^{4 m+2 a-i-j+1}(1+x)^{2 b+i+j+1} x^{n}}_{\mathcal{L}^{(n)}(4 m+2 a-i-j+1,2 b+i+j+1)} \tag{56}
\end{align*}
$$

where we have introduced the following integral:

$$
\begin{align*}
\mathcal{L}^{(n)}(w, z) & =\int_{-1}^{+1} d x(1-x)^{w}(1+x)^{z} x^{n} \\
& =2^{w+z+1} \sum_{k=0}^{n}\binom{n}{k}(-2)^{k} B(w+k+1, z+1) \tag{57}
\end{align*}
$$

introducing Euler's Beta function $B(x, y)=\Gamma(x) \Gamma(y) / \Gamma(x+y)$. Inserting this result into equation (56) we obtain

$$
\begin{align*}
& >I_{2 m, 2 m+1}^{(n)}=2^{2(a+b+1)+1} \sum_{i, j=0}^{2 m}(-1)^{4 m-i-j} c_{i}^{(2 m)} c_{j}^{(2 m)} \\
& \times \sum_{k=0}^{n}\binom{n}{k}(-2)^{k} B(2(2 m+a+1)+k-i-j, 2(b+1)+i+j) \tag{58}
\end{align*}
$$

In order to compute the second type of integrals $\left(I_{2 m+1,2 m}^{(n)}\right)$ in (54) and (55), we use (44), (45) and (50) to get

$$
\begin{align*}
I_{2 m+1,2 m}^{(n)}= & \frac{1}{2^{2 m+1}} \sum_{i=0}^{2 m} c_{i}^{(2 m)} \int_{-1}^{+1} d y w_{a, b}(y)(y-1)^{2 m-i}(y+1)^{i} \\
& \times \underbrace{\int_{-1}^{+1} d x \operatorname{sign}(x-y) \phi_{2 m+1}(x) x^{n}}_{\mathcal{K}_{2 m+1}^{(n)}(y)} . \tag{59}
\end{align*}
$$

The integral $\mathcal{K}_{2 m+1}^{(n)}(y)$ can be rewritten as follows by exploiting (44)

$$
\begin{equation*}
\mathcal{K}_{2 m+1}^{(n)}(y)=A_{2 m+1} \mathcal{G}_{a, b}^{(n)}(2 m+1 ; y)-B_{2 m-1} \mathcal{G}_{a, b}^{(n)}(2 m-1 ; y) \tag{60}
\end{equation*}
$$

where we have

$$
\begin{align*}
& \mathcal{G}_{a, b}^{(n)}(m ; y)=\int_{-1}^{+1} d x \operatorname{sign}(x-y) w_{a, b}(x) P_{m}^{(2 a+1,2 b+1)}(x) x^{n} \\
& =\frac{1}{2^{m}} \sum_{j=0}^{m}(-1)^{m-j} c_{j}^{(m)} \int_{-1}^{+1} d x \operatorname{sign}(x-y)(1-x)^{a+m-j}(1+x)^{b+j} x^{n} \\
& =2^{a+b+1} \sum_{j=0}^{m}(-1)^{m-j} c_{j}^{(m)} \sum_{k=0}^{n}\binom{n}{k}(-2)^{k} \widetilde{B}_{\frac{1-y}{2}}(a+m+k-j+1, b+j+1), \tag{61}
\end{align*}
$$

where $\widetilde{B}_{y}(w, z)=2 B_{y}(w, z)-B(w, z)$ and $B_{y}(w, z)=\int_{0}^{y} d t t^{w-1}(1-t)^{z-1}$ is the incomplete Beta function. Therefore we get

$$
\begin{align*}
& \mathcal{K}_{2 m+1}^{(n)}(y)=2^{a+b+1} \sum_{k=0}^{n}\binom{n}{k}(-2)^{k} \\
& \times\left[A_{2 m+1} \sum_{j=0}^{2 m+1}(-1)^{2 m-j+1} c_{j}^{(2 m+1)} \widetilde{B}_{\frac{1-y}{2}}(2(m+1)+a+k-j, b+j+1)\right. \\
& \left.-B_{2 m-1} \sum_{j=0}^{2 m-1}(-1)^{2 m-j-1} c_{j}^{(2 m-1)} \widetilde{B}_{\frac{1-y}{2}}(2 m+a+k-j, b+j+1)\right] \tag{62}
\end{align*}
$$

Eventually, inserting this result into (59) and computing the integral in $d y$, we obtain

$$
\begin{align*}
& >I_{2 m+1,2 m}^{(n)}=\frac{2^{a+b}}{2^{2 m}} \sum_{i=0}^{2 m} c_{i}^{(2 m)} \sum_{k=0}^{n}\binom{n}{k}(-2)^{k} \\
& \times\left[A_{2 m+1} \sum_{j=0}^{2 m+1}(-1)^{2 m-j+1} c_{j}^{(2 m+1)} \Omega_{a, b}(2 m-i, i ; 2(m+1)+a+k-j, b+j+1)\right. \\
& \left.-B_{2 m-1} \sum_{j=0}^{2 m-1}(-1)^{2 m-j-1} c_{j}^{(2 m-1)} \Omega_{a, b}(2 m-i, i ; 2 m+a+k-j, b+j+1)\right] \tag{63}
\end{align*}
$$

where we have

$$
\begin{align*}
& \Omega_{a, b}(h, \ell ; w, z)=\int_{-1}^{+1} d x w_{a, b}(x)(x-1)^{h}(x+1)^{\ell} \widetilde{B}_{\frac{1-x}{2}}(w, z) \\
& =(-1)^{h} 2^{a+b+h+\ell+1}\left[w^{-1} B(1+b+\ell, 1+a+h+w)\right. \\
& \times{ }_{3} F_{2}(w, 1+a+h+w, 1-z ; 1+w, 2+a+b+h+\ell+w ; 1) \\
& -B(w, z) B(1+a+h, 1+b+\ell)] \tag{64}
\end{align*}
$$

In (64), we have used the following generalized hyper geometric function:

$$
\begin{equation*}
{ }_{3} F_{2}\left(a_{1}, a_{2}, a_{3} ; b_{1}, b_{2} ; z\right):=\sum_{k=0}^{\infty} \frac{\left(a_{1}\right)_{k}\left(a_{2}\right)_{k}\left(a_{3}\right)_{k}}{\left(b_{1}\right)_{k}\left(b_{2}\right)_{k} k!} z^{k} \tag{65}
\end{equation*}
$$

### 3.3. Moments $\beta=2$

We wish now to compute the generic $n$-th moment of the density in (38), which reads

$$
\begin{equation*}
\tau_{n}^{(\mathrm{J})}(N, 2, a, b)=\left\langle\sum_{i=1}^{N} x_{i}^{n}\right\rangle=N \int_{-1}^{+1} d x \rho_{N, 2, a, b}^{(\mathrm{J})}(x) x^{n} \tag{66}
\end{equation*}
$$

By making use of the representation of Jacobi polynomials in (50), the $n$-th moment in (66) can be written as follows (throughout all the present subsection the $c_{j}^{(k)}$ coefficients will depend on the pair $\left.(a, b)\right)$ :

$$
\begin{align*}
\tau_{n}^{(\mathrm{J})}(N, 2, a, b) & =\sum_{k=0}^{N-1}\left(h_{k}^{a, b}\right)^{-1} \int_{-1}^{+1} d x w_{a, b}(x)\left[P_{k}^{(a, b)}(x)\right]^{2} x^{n} \\
& =\sum_{k=0}^{N-1} \sum_{i, j=0}^{k} \frac{c_{i}^{(k)} c_{j}^{(k)}}{2^{2 k} h_{k}^{a, b}} I^{(n)}(k, i, j, a, b) \tag{67}
\end{align*}
$$

where

$$
\begin{equation*}
I^{(n)}(k, i, j, a, b)=(-1)^{2 k-i-j} \int_{-1}^{+1} d x(1-x)^{2 k+a-i-j}(1+x)^{b+i+j} x^{n} \tag{68}
\end{equation*}
$$

and can be computed by changing variables and setting $y=(1-x) / 2$. When doing so, one obtains

$$
\begin{equation*}
I^{(n)}(k, i, j, a, b)=(-1)^{2 k-i-j} 2^{2 k+a+b+1} \int_{0}^{1} d y y^{2 k+a-i-j}(1-y)^{b+i+j}(1-2 y)^{n} \tag{69}
\end{equation*}
$$

and expanding $(1-2 y)^{n}$ according to the binomial Theorem, one gets

$$
\begin{align*}
I^{(n)}(k, i, j, a, b)= & (-1)^{2 k-i-j} 2^{2 k+a+b+1} \sum_{\ell=0}^{n}(-2)^{\ell}\binom{n}{\ell} \\
& \times \int_{0}^{1} d y y^{2 k+a+\ell-i-j}(1-y)^{b+i+j} \tag{70}
\end{align*}
$$

The integral appearing in the previous equation is of the following kind

$$
\begin{equation*}
\int_{0}^{1} d t t^{x-1}(1-t)^{y-1}=B(x, y) \tag{71}
\end{equation*}
$$

Thus we get

$$
\begin{align*}
I^{(n)}(k, i, j, a, b)= & (-1)^{2 k-i-j} 2^{2 k+a+b+1} \sum_{\ell=0}^{n}(-2)^{\ell}\binom{n}{\ell} \\
& \times B(2 k+a+\ell-i-j+1, b+i+j+1) \tag{72}
\end{align*}
$$

Plugging this result in equation (67), one can easily see that the $n$-th moment has the compact expression

$$
\begin{equation*}
\tau_{n}^{(\mathrm{J})}(N, 2, a, b)=\sum_{k=0}^{N-1} \sum_{i, j=0}^{k} \sum_{\ell=0}^{n} \mathfrak{y}_{k, i, j, \ell} B(2 k+a+\ell-i-j+1, b+i+j+1) \tag{73}
\end{equation*}
$$

where

$$
\begin{equation*}
\mathfrak{y}_{k, i, j, \ell}=\frac{(-1)^{2 k-i-j+\ell} c_{i}^{(k)} c_{j}^{(k)} 2^{a+b+1+\ell}}{h_{k}^{a, b}}\binom{n}{\ell} . \tag{74}
\end{equation*}
$$

After inserting (73) into (52), we have checked that the special case $\mathfrak{a}=\mu$, $\mathfrak{b}=0$ for $\tau_{n}^{(\mathrm{sJ})}(N, 2, \mathfrak{a}, \mathfrak{b})$ indeed agrees with Eq. (16) in [31] and with Eq. (13) in [32] as it should (see Appendix A for details).

### 3.4. Moments $\beta=4$

The $n$-th moment of the density in (39) reads:

$$
\begin{equation*}
\tau_{n}^{(\mathrm{J})}(N, 4, a, b)=\frac{1}{2} \tau_{n}^{(\mathrm{J})}(2 N, 2,2 \hat{a}+1,2 \hat{b}+1)-\frac{1}{2} \sum_{m=0}^{N-1} \Theta_{N, m} \tag{75}
\end{equation*}
$$

where:

$$
\begin{align*}
\Theta_{N, m} & =K_{2 N, m}^{(2 \hat{a}+1,2 \hat{b}+1)} Y_{2 N, 2 m}^{(n)}(2 \hat{a}+1,2 \hat{b}+1),  \tag{76}\\
Y_{k, \ell}^{(n)}(a, b) & =\int_{-1}^{+1} d x w_{a, b}(x) P_{k}^{(a, b)}(x) P_{\ell}^{(a, b)}(x) x^{n} . \tag{77}
\end{align*}
$$

Exploiting the Jacobi polynomial expansion (50) and the binomial Theorem we get (again, we omit the explicit dependence of the $c_{i}^{(k)}$ coefficients on the pair $(a, b))$

$$
\begin{align*}
Y_{k, \ell}^{(n)}(a, b)= & \frac{1}{2^{k+\ell}} \sum_{i=0}^{k} \sum_{j=0}^{\ell}(-1)^{k+\ell-i-j} c_{i}^{(k)} c_{j}^{(\ell)} \\
& \times \int_{-1}^{+1} d x(1-x)^{a+k+\ell-i-j}(1+x)^{b+i+j} x^{n} \\
= & 2^{a+b+1} \sum_{i=0}^{k} \sum_{j=0}^{\ell} \sum_{s=0}^{n}\binom{n}{s}(-1)^{k+\ell-i-j}(-2)^{s} c_{i}^{(k)} c_{j}^{(\ell)} \\
& \times B(a+k+\ell+s-i-j+1, b+i+j+1) \tag{78}
\end{align*}
$$

### 3.5. Comparison with numerics

In Fig. 2 we plot the analytical formulae (37), (38) and (39) together with numerical diagonalization of matrices $\mathcal{J}$ from the Jacobi ensemble with $\beta=1,2,4$ respectively, obtained in Matlab using the algorithm by Edelman and Sutton [22].


Fig. 2. Average spectral density of the Jacobi ensemble for $N=6, a=5, b=9$ and $\beta=1,2,4$ (clockwise from top left ). Numerical diagonalization is given in (blue) triangles, while theoretical results are in solid (black).

In the following tables, we compare the analytical results for moments of the Jacobi and shifted-Jacobi ensembles with numerical simulations via the algorithm by Edelman and Sutton [22].

TABLE II
Comparison between theory and numerics for moments of the Jacobi ensemble $(N=6, a=5, b=9)$ (the numerical results are obtained averaging over $\mathcal{O}\left(10^{5}\right)$ samples).

| $\beta$ | $n$ | Theory | Numerics |
| :---: | :---: | :---: | :---: |
| 1 | 1 | 1.1428 | 1.1454 |
| 1 | 2 | 1.1536 | 1.1573 |
| 1 | 3 | 0.5108 | 0.5125 |
| 2 | 1 | 0.9230 | 0.9259 |
| 2 | 2 | 1.4872 | 1.4892 |
| 2 | 3 | 0.5470 | 0.5485 |
| 4 | 1 | 0.6666 | 0.6674 |
| 4 | 2 | 1.9300 | 1.9294 |
| 4 | 3 | 0.5602 | 0.5602 |

TABLE III
Comparison between theory and numerics for moments of the shifted-Jacobi ensemble $(N=6, \mathfrak{a}=5, \mathfrak{b}=9)$ (the numerical results are obtained averaging over $\mathcal{O}\left(10^{5}\right)$ samples).

| $\beta$ | $n$ | Theory | Numerics |
| :---: | :---: | :---: | :---: |
| 1 | 1 | 2.4286 | 2.4300 |
| 1 | 2 | 1.2170 | 1.2169 |
| 1 | 3 | 0.6902 | 0.6902 |
| 2 | 1 | 2.5385 | 2.5390 |
| 2 | 2 | 1.4103 | 1.4102 |
| 2 | 3 | 0.8932 | 0.8919 |
| 4 | 1 | 2.6667 | 2.6669 |
| 4 | 2 | 1.6492 | 1.6493 |
| 4 | 3 | 1.1605 | 1.1612 |

## 4. Conclusions

In conclusions, we have first collected well-known formulae for the spectral density of Laguerre and Jacobi ensembles of random matrices with orthogonal, unitary and symplectic symmetry. We feel that this paper might serve as a quick reference point for formulae that might be hard to dig out in the literature. Using these results, we have computed the average of integer moments, reducing the complexity from a $N$-fold integration to a single integral over the spectral density. In all cases, expressions different from ours and derived through different methods already exist [24, 31, 32] (see also Appendix A). It would be interesting to prove mathematically the equiv-
alence of various formulae which are now available for the same objects. The present paper offers a different and possibly simpler way of deriving transport moments for all symmetry classes and finite number of open electronic channels. The obtained results have been checked numerically with high accuracy, and the corresponding Matlab codes are freely available on demand.

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## Appendix A

## Electronic transport in open cavities

A cavity of submicron dimensions etched in a semiconductor can be connected to the external world by two leads supporting $N_{1}$ and $N_{2}$ electronic channels. A voltage difference $V$ applied between the two leads lets an electronic current flow through the cavity, whose intensity presents timedependent fluctuations which persist down to zero temperature [34]. These are associated with the granularity of the electron charge $e$. Typical phenomena observed in experiments include weak localization [35], universality in conductance fluctuations [36] and constant Fano factor [37]. According to the Landauer-Büttiker scattering approach [34, 38, 39], the statistics of quantum transport observables depend in a rather simple way on the scattering matrix $\mathcal{S}$ of the cavity. This is a unitary $N_{0} \times N_{0}$ matrix (where $N_{0}=N_{1}+N_{2}$ ) which relates the wave function coefficients of the incoming and outgoing electrons in a natural basis:

$$
\mathcal{S}=\left(\begin{array}{cc}
\boldsymbol{r} & \boldsymbol{t}^{\prime}  \tag{A.1}\\
\boldsymbol{t} & \boldsymbol{r}^{\prime}
\end{array}\right)
$$

The transmission $\left(\boldsymbol{t}, \boldsymbol{t}^{\prime}\right)$ and reflection $\left(\boldsymbol{r}, \boldsymbol{r}^{\prime}\right)$ blocks are submatrices encoding the transmission and reflection coefficients among different channels ${ }^{4}$. The eigenvalues of the hermitian transport matrix $\boldsymbol{T}=\boldsymbol{t} \boldsymbol{t}^{\dagger}$ are of primary importance: for example, the dimensionless conductance and the shot noise are given respectively by $G=\operatorname{Tr}(\boldsymbol{T})$ [38] and $P=\operatorname{Tr}[\boldsymbol{T}(\mathbf{1}-\boldsymbol{T})$ ] [40], while other transport properties are encoded in higher moments $\tau_{n}=\operatorname{Tr}\left[\boldsymbol{T}^{n}\right]$.

[^1]When the classical electronic motion inside the cavity can be regarded as chaotic, Random Matrix Theory has been very successful in describing the statistics of universal fluctuations in such systems: the scattering matrix $\mathcal{S}$ is assumed to be drawn at random from a suitable ensemble of matrices, with the overall constraint of unitarity [41, 42, 43]. A maximum entropy approach (under the assumption of ballistic point contacts [34]) forces the probability distribution of $\mathcal{S}$ to be uniform within the unitary group, i.e. $\mathcal{S}$ belongs to one of Dyson's Circular Ensembles [27, 44].

The uniformity of $\mathcal{S}$ within the unitary group induces the following remarkably simple j.p.d. of transmission eigenvalues $\left\{T_{i}\right\}$ of the matrix $\boldsymbol{T}$ [34, 43, 45]:

$$
\begin{equation*}
\mathcal{P}_{\beta}^{(\boldsymbol{T})}\left(T_{1}, \ldots, T_{N}\right) \propto \prod_{j<k}\left|T_{j}-T_{k}\right|^{\beta} \prod_{i=1}^{N} T_{i}^{\frac{\beta}{2}(\mu+1)-1} \tag{A.2}
\end{equation*}
$$

with $N=\min \left(N_{1}, N_{2}\right), \mu=\left|N_{1}-N_{2}\right|$ and $\beta=1,2,4$ in the case of preserved time-reversal symmetry, broken time-reversal symmetry and spin-flip symmetry respectively. The j.p.d. (A.2) is precisely of the shifted-Jacobi form (4) with $\mathfrak{a}=(\beta / 2)(\mu+1)-1$ and $\mathfrak{b}=0$.

The transmission eigenvalues $T_{i}$ are thus correlated real random variables between 0 and 1, whose j.p.d. (A.2) in principle allows for a complete characterization of statistical properties of experimental observables. For most recent analytical results, we refer to $[30,31,32,33,46,47,48,49,50$, 51, 52, 53, 54, 55].

In particular, the study of higher moments of the transmission matrix $\tau_{n}^{\star}=\operatorname{Tr}\left[\boldsymbol{T}^{n}\right]$ has recently seen many analytical progresses ${ }^{5}$ [24, 25, 30, 31, $32,33,48,49,53]$. In particular, we now have two different (but equivalent) formulae for higher moments for $\beta=2$ and arbitrary $N_{1}, N_{2}$ :

$$
\begin{align*}
\tau_{n}^{\star} & =\sum_{p=0}^{N-1}(2 p+\mu+1) \sum_{k, \ell=0}^{p} \frac{g_{p}(k) g_{p}(\ell)}{\mu+n+k+\ell+1}, \quad \text { see [31] }  \tag{A.3}\\
\tau_{n}^{\star} & =\sum_{p=0}^{n-1} \frac{(-1)^{p}}{n!}\binom{n-1}{p} \frac{\left(N_{1}-p\right)_{n}\left(N_{2}-p\right)_{n}}{\left(N_{1}+N_{2}-p\right)_{n}}, \quad \text { see [32] } \tag{A.4}
\end{align*}
$$

where

$$
\begin{equation*}
g_{p}(\kappa):=(-1)^{\kappa}\binom{p}{\kappa}\binom{p+\mu+\kappa}{\mu+\kappa} \tag{A.5}
\end{equation*}
$$

[^2]Novaes [33] has also computed the leading $\mathcal{O}(N)$ term in the large $N_{1}, N_{2} \gg 1$ expansion of the moments $\tau_{n}^{(\mathrm{sJ})}(\infty, 2, \mu, 0)$ (Eq. (A.4)) for the quantum transport problem at $\beta=2$. His formula in our notation reads

$$
\begin{equation*}
\tau_{n}^{(\mathrm{sJ})}(\infty, 2, \mu, 0) \sim\left(N_{1}+N_{2}\right) \sum_{p=1}^{n}\binom{n-1}{p-1}(-1)^{p-1} c_{p-1} \xi^{p} \tag{A.6}
\end{equation*}
$$

where $c_{p}=\frac{1}{p+1}\binom{2 p}{p}$ and $\xi=N_{1} N_{2} /\left(N_{1}+N_{2}\right)^{2}$. In the next appendix, we will show how this result (for the case $\mu \sim \mathcal{O}(1) \rightarrow \xi=1 / 4)$ can be derived by integration of the asymptotic spectral density of shifted Jacobi ensemble with $\mathfrak{a}=\mathfrak{b} \sim \mathcal{O}(1)$ when $N \rightarrow \infty$.

In Fig. 3 we plot the transport moments $\tau_{n}\left(N_{1}, \beta,(\beta / 2)(\mu+1)-1,0\right)$ for $n=3$ and $n=4$ as a function of $N_{1}$, while the number of channels $N_{2}=N_{1}+2$ is held fixed. All cases yield expressions formally different from others [24] (see equation (52) and the equations in Subsections 3.2 and 3.4), but equivalent.


Fig. 3. Transport moments $\tau_{n}\left(N_{1}, \beta,(\beta / 2)(\mu+1)-1,0\right)$ for $n=3$ (circles) and $n=4$ (triangles) as a function of $N_{1}$. The number of channels $N_{2}=N_{1}+2$ is held fixed. The results are for $\beta=1$ (short-dashed, red), $\beta=2$ (long-dashed, blue) and $\beta=4$ (solid, black).

## Appendix B

Large $N$ asymptotics for moments
The general integral formula (14) offers a neat and explicit way to obtain the large $N$ asymptotic behavior for moments and possibly the full $1 / N$ expansion as follows. Suppose that the spectral density of the ensemble under discussion admits the following series expansion:

$$
\begin{equation*}
\rho(x)=\rho^{(\infty)}(x)+\frac{1}{N} \rho^{(1)}(x)+\frac{1}{N^{2}} \rho^{(2)}(x)+\ldots \tag{B.1}
\end{equation*}
$$

then the corresponding series for moments can be formally obtained as:

$$
\begin{equation*}
\tau_{n}=N \int d x \rho^{(\infty)}(x) x^{n}+\int d x \rho^{(1)}(x) x^{n}+\frac{1}{N} \int d x \rho^{(2)}(x) x^{n}+\ldots \tag{B.2}
\end{equation*}
$$

A first check of the validity of the putative formula (B.2) goes as follows. Novaes [33] has computed the leading $\mathcal{O}(N)$ term in the large $N_{1}, N_{2} \gg 1$ expansion of the moments $\tau_{n}^{(\mathrm{sJ})}(\infty, 2, \mu, 0)$ for the quantum transport problem in chaotic cavities with unitary symmetry supporting $N_{1}, N_{2}$ electronic channels in the leads $\left(\mu=\left|N_{1}-N_{2}\right|\right.$ and $\left.N=\min \left(N_{1}, N_{2}\right)\right)$, see Eq. (A.6). Now, the $N \rightarrow \infty$ limit for the spectral density of the ordinary Jacobi ensemble with $a, b \sim \mathcal{O}(1)$ reads [29]

$$
\begin{equation*}
\rho_{\infty, \beta, a, b}^{(\mathrm{J})}(x)=\frac{1}{\pi \sqrt{1-x^{2}}} \tag{B.3}
\end{equation*}
$$

i.e. it is independent of $a, b, \beta$. Using (9) we get

$$
\begin{equation*}
\rho_{\infty, \beta, \mathfrak{a}, \mathfrak{b}}^{(\mathrm{sJ})}(x)=\frac{2}{\pi \sqrt{1-(1-2 x)^{2}}} \tag{B.4}
\end{equation*}
$$

and using (B.2), the $\mathcal{O}(N)$ term in the expansion of the $n$-th moment of the $\beta=2$ shifted-Jacobi ensemble should read

$$
\begin{equation*}
\lim _{N \rightarrow \infty} \frac{1}{N} \tau_{n}^{(\mathrm{sJ})}(N, 2, \mathfrak{a}, \mathfrak{b})=\int_{0}^{1} d x \frac{2 x^{n}}{\pi \sqrt{1-(1-2 x)^{2}}}=\frac{\Gamma(n+1 / 2)}{\sqrt{\pi} \Gamma(1+n)} \tag{B.5}
\end{equation*}
$$

Eq. (B.5) should then be compared with (A.6) for a special value of $\xi$, namely the one that guarantees that parameters $\mathfrak{a}, \mathfrak{b}$ are finite (i.e. of $\mathcal{O}(1))$ when $N \rightarrow \infty$ (otherwise Eq. (B.3) would not hold). Since $\mathfrak{a} \equiv \mu=N_{1}-N_{2}$, we need to impose that this difference is finite when $N_{1}, N_{2} \rightarrow \infty$. This immediately leads to $\xi=1 / 4$ and $N_{1}+N_{2}=2 N$. Inserting these values into (A.6), we indeed find the identity:

$$
\begin{equation*}
2 \sum_{p=1}^{n}\binom{n-1}{p-1}(-1)^{p-1} c_{p-1} 4^{-p}=\frac{\Gamma(n+1 / 2)}{\sqrt{\pi} \Gamma(1+n)} \tag{B.6}
\end{equation*}
$$

Therefore, we have proven that a simple one-dimensional integral over the asymptotic spectral density of the shifted-Jacobi ensemble correctly reproduces Novaes' formula for the asymptotics of transport moment (obtained by a totally different method) when the difference between the number of electronic channels in the two leads is of $\mathcal{O}(1)$.

## REFERENCES

[1] J. Wishart, Biometrika 20A, 32 (1928).
[2] A.T. James, Ann. Math. Stat. 35, 475 (1964).
[3] R.J. Muirhead, Aspects of Multivariate Statistical Theory, John Wiley \& Sons, Inc., 2005.
[4] S.S. Wilks, Mathematical Statistics, John Wiley \& Sons, New York 1962.
[5] I.M. Johnstone, Ann. Statist. 29, 295 (2001).
[6] R.W. Preisendorfer, Principal Component Analysis in Meteorology and Oceanography, Elsevier, New York 1988.
[7] J.-P. Bouchaud, M. Potters, Theory of Financial Risks, Cambridge University Press, Cambridge 2001; L. Laloux, P. Cizeau, J.-P. Bouchaud, M. Potters, Phys. Rev. Lett. 83, 1467 (1999).
[8] Z. Burda, J. Jurkiewicz, Physica A 344, 67 (2004); Z. Burda, J. Jurkiewicz, B. Wacław, Acta Phys. Pol. B 36, 2641 (2005).
[9] A.M. Sengupta, P.P. Mitra, Phys. Rev. E60, 3389 (1999); E. Telatar, European Transactions on Telecommunications 10, 585 (1999);
P. Kazakopoulos, P. Mertikopoulos, A.L. Moustakas, G. Caire, IEEE Transactions on Information Theory 57, 1984 (2011).
[10] Y.V. Fyodorov, H.-J. Sommers, J. Math. Phys. 38, 1918 (1997);
Y.V. Fyodorov, B.A. Khoruzhenko, Phys. Rev. Lett. 83, 65 (1999).
[11] J.J.M. Verbaarschot, Handbook Article on "Applications of Random Matrix Theory to QCD", invited chapter in Handbook on Random Matrix Theory, Eds. G. Akemann, J. Baik, P. Di Francesco, Oxford University Press, 2011 [arXiv:0910. 4134 [hep-th]].
[12] K. Johansson, Commun. Math. Phys. 209, 437 (2000).
[13] S. Maslov, Y.C. Zhang, Phys. Rev. Lett. 87, 248701 (2001).
[14] Y.K. Yu, Y.C. Zhang, Physica A 312, 1 (2002).
[15] R.A. Janik, M.A. Nowak, J. Phys. A36, 3629 (2003).
[16] P. Vivo, S.N. Majumdar, O. Bohigas, J. Phys. A40, 4317 (2007).
[17] E. Katzav, I.P. Castillo, Phys. Rev. E82, 040104 (2010).
[18] C. Nadal, S.N. Majumdar, Phys. Rev. E79, 061117 (2009).
[19] S.N. Majumdar, Handbook Article on "Extreme Eigenvalues of Wishart Matrices: Application to Entangled Bipartite System", invited chapter in Handbook on Random Matrix Theory, Eds. G. Akemann, J. Baik, P. Di Francesco, Oxford University Press, 2011 [arXiv:1005.4515].
[20] B. Collins, Probab. Theory Rel. Fields 133, 315 (2005).
[21] I. Dumitriu, A. Edelman, J. Math. Phys. 43, 5830 (2002).
[22] A. Edelman, B.D. Sutton, Found. Comput. Math. 8, 259 (2008).
[23] C.E.I. Redelmeier, J. Theor. Probab. (2010), http://dx.doi.org/10.1007/s10959-010-0278-7.
[24] F. Mezzadri, N.J. Simm, arXiv:1103. 6203 [math-ph]; See also Poster at VI BRUNEL Workshop on Random Matrix Theory, December 17-18, 2010; Talk given at "Random Matrix Theory and Its Applications I", MSRI Berkeley, Sept. 2010.
[25] G. Berkolaiko, J. Kuipers, arXiv:1012.3526v1 [nlin.CD].
[26] J. Ambjørn, C.F. Kristjansen, Y.M. Makeenko, Mod. Phys. Lett. A7, 3187 (1992).
[27] M.L. Mehta, Random Matrices, 3rd edition, Elsevier-Academic Press, 2004.
[28] T. Nagao, M. Wadati, J. Phys. Soc. Jpn. 60, 3298 (1997).
[29] S. Ghosh, A. Pandey, Phys. Rev. E65, 046221 (2002).
[30] S. Kumar, A. Pandey, to be published.
[31] P. Vivo, E. Vivo, J. Phys. A41, 122004 (2008).
[32] M. Novaes, Phys. Rev. B78, 035337 (2008).
[33] M. Novaes, Phys. Rev. B75, 073304 (2007).
[34] C.W.J. Beenakker, Rev. Mod. Phys. 69, 731 (1997).
[35] A.M. Chang et al., Phys. Rev. Lett. 73, 2111 (1994).
[36] C.M. Marcus et al., Phys. Rev. Lett. 69, 506 (1992).
[37] S. Oberholzer et al., Phys. Rev. Lett. 86, 2114 (2001).
[38] R. Landauer, IBM J. Res. Dev. 1, 223 (1957); Phil. Mag. 21, 863 (1970); D.S. Fisher, P.A. Lee, Phys. Rev. B23, 6851 (1981).
[39] M. Büttiker, Phys. Rev. Lett. 57, 1761 (1986).
[40] Ya.M. Blanter, M. Büttiker, Phys. Rep. 336, 1 (2000).
[41] K.A. Muttalib, J.L. Pichard, A.D. Stone, Phys. Rev. Lett. 59, 2475 (1987).
[42] A.D. Stone et al., in Mesoscopic Phenomena in Solids, ed. B.L. Altshuler, P.A. Lee and R.A. Webb, North Holland, Amsterdam 1991.
[43] P.A. Mello, P. Pereyra, N. Kumar, Ann. Phys. 181, 290 (1988).
[44] F.J. Dyson, J. Math. Phys. 3, 140 (1962); J. Math. Phys. 3, 157 (1962); J. Math. Phys. 3, 166 (1962).
[45] P.J. Forrester, J. Phys. A39, 6861 (2006).
[46] P.W. Brouwer, C.W.J. Beenakker, Math. Phys. 37, 4904 (1996).
[47] D.V. Savin, H.-J. Sommers, Phys. Rev. B73, 081307(R) (2006).
[48] D.V. Savin, H.-J. Sommers, W. Wieczorek, Phys. Rev. B77, 125332 (2008).
[49] J.-G. Luque, P. Vivo, J. Phys. A43, 085213 (2010); C. Carré, M. Deneufchâtel, J.-G. Luque, P. Vivo, J. Math. Phys. 51, 123516 (2010); M. Novaes, Ann. Phys. 326, 828 (2011).
[50] H.-J. Sommers, W. Wieczorek, D.V. Savin, Acta Phys. Pol. A 112, 691 (2007).
[51] B.A. Khoruzhenko, D.V. Savin, H.-J. Sommers, Phys. Rev. B80, 125301 (2009).
[52] V.Al. Osipov, E. Kanzieper, Phys. Rev. Lett. 101, 176804 (2008); J. Phys. A42, 475101 (2009).
[53] P. Vivo, S.N. Majumdar, O. Bohigas, Phys. Rev. Lett. 101, 216809 (2008); Phys. Rev. B81, 104202 (2010).
[54] S.H. Simon, A.L. Moustakas, Phys. Rev. Lett. 96, 136805 (2006).
[55] J.P. Dahlhaus, B. Béri, C.W.J. Beenakker, Phys. Rev. B82, 014536 (2010).


[^0]:    ${ }^{3}$ For $\beta=1, N$ is even. The expression for $N$ odd is slightly more complicated [29].

[^1]:    ${ }^{4}\left(\boldsymbol{t}, \boldsymbol{t}^{\prime}\right)$ are respectively of size $N_{1} \times N_{2}$ and $N_{2} \times N_{1}$, while $\left(\boldsymbol{r}, \boldsymbol{r}^{\prime}\right)$ are of size $N_{2} \times N_{2}$ and $N_{1} \times N_{1}$.

[^2]:    ${ }^{5}$ In our notation, $\tau_{n}^{\star} \equiv \tau_{n}^{(\mathrm{sJ})}(N, 2, \mu, 0)$.

