HIDDEN-SECTOR CORRECTION TO COULOMB POTENTIAL THROUGH THE PHOTONIC PORTAL*

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We show that in the model of hidden sector of the Universe, interacting with the Standard-Model sector through the photonic portal, the Standard-Model Coulomb potential gets a tiny hidden-sector additive correction that might turn out to be either exciting or fatal for the verification of this model.

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1. Introduction

A model of hidden sector of the Universe, proposed in previous papers [1, 2], consists of sterile spin-1/2 Dirac fermions ("sterinos"), sterile spin-0 bosons ("sterons") and sterile non gauge mediating bosons ("A bosons") described by an antisymmetric-tensor field (of dimension one) weakly coupled to steron-photon pairs as well as to antisterino-sterino pairs, giving the new weak interaction

$$-\frac{1}{2}\sqrt{f}\left(\varphi F_{\mu\nu} + \zeta \bar{\psi} \sigma_{\mu\nu} \psi\right) A^{\mu\nu} \,. \tag{1}$$

Here, $F_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu}$ is the Standard-Model electromagnetic field (of dimension two), while \sqrt{f} and $\sqrt{f}\zeta$ denote two dimensionless small coupling constants. We presume that

$$\varphi = \langle \varphi \rangle_{\rm vac} + \varphi_{\rm ph} \tag{2}$$

with a spontaneously non zero vacuum expectation value $\langle \varphi \rangle_{\text{vac}} \neq 0$. We have called such a coupling of photons to the hidden sector "photonic portal" (to the hidden sector), and considered it as an alternative to the popular "Higgs portal" [3].

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In the present note, we show that the first term in the interaction (1) implies a small correction to the Standard-Model Coulomb potential. It can be estimated, if the coupling constant f, the vacuum expectation value $\langle \varphi \rangle_{\text{vac}}$ and the masses of sterino, steron and A boson, m_{ψ}, m_{φ} and M, are somehow established.

The new interaction Lagrangian (1), jointly with the A-boson kinetic and Standard-Model electromagnetic Lagrangians, lead to the following field equations for $F_{\mu\nu}$ and $A_{\mu\nu}$

$$\partial^{\nu} \left[F_{\mu\nu} + \sqrt{f} \left(\langle \varphi \rangle_{\text{vac}} + \varphi_{\text{ph}} \right) A_{\mu\nu} \right] = -j_{\mu} , \qquad F_{\mu\nu} = \partial_{\mu} A_{\nu} - \partial_{\nu} A_{\mu} \quad (3)$$

and

$$\left(\Box - M^2\right) A_{\mu\nu} = -\sqrt{f} \left[(\langle \varphi \rangle_{\rm vac} + \varphi_{\rm ph}) F_{\mu\nu} + \zeta \bar{\psi} \sigma_{\mu\nu} \psi \right] , \qquad (4)$$

where j_{μ} denotes the Standard-Model electric current and M stands for a mass scale of A bosons, expected typically to be large.

The field equations (3) are Maxwell's equations modified in the presence of hidden sector interacting with the Standard-Model sector through our photonic portal. This modification is of magnetic character, because the hidden-sector contribution to the total electric source-current

$$j_{\mu} + \partial^{\nu} \left[\sqrt{f} \left(\langle \varphi \rangle_{\rm vac} + \varphi_{\rm ph} \right) A_{\mu\nu} \right] \tag{5}$$

of the electromagnetic field A_{μ} is a four-divergence giving no contribution to the total electric charge

$$\int d^3x \left\{ j_0 + \partial^k \left[\sqrt{f} \left(\langle \varphi \rangle_{\text{vac}} + \varphi_{\text{ph}} \right) A_{0k} \right] \right\} = \int d^3x j_0 = Q.$$
 (6)

It can be seen that the vacuum expectation value $\langle \varphi \rangle_{\text{vac}} \neq 0$ generates spontaneously a small sterino magnetic moment

$$\mu_{\psi} = \frac{f\zeta}{M^2} \langle \varphi \rangle_{\rm vac} \,, \tag{7}$$

although sterinos are electrically neutral. This is a consequence of an effective sterino magnetic interaction

$$-\frac{1}{2}\mu_{\psi}\bar{\psi}\sigma_{\mu\nu}\psi F^{\mu\nu} \tag{8}$$

appearing, when the low-momentum-transfer approximation

$$A_{\mu\nu} \simeq \frac{\sqrt{f} \zeta}{M^2} \bar{\psi} \sigma_{\mu\nu} \psi \tag{9}$$

effectively implied by Eq. (4) with $F_{\mu\nu} \to 0$ is used in the second term of weak interaction (1).

It will turn out, however, that due to the photonic portal the Standard-Model electromagnetic interaction, in particular the Coulomb potential, gets a tiny hidden-sector correction.

2. Hidden-sector correction to Coulomb potential

From Eq. (4) we can see that

$$\left(\Box - M^2\right)\partial^{\nu}A_{\mu\nu} = -\sqrt{f}\,\partial^{\nu}\left(\varphi F_{\mu\nu} + \zeta\bar{\psi}\sigma_{\mu\nu}\psi\right)\,.\tag{10}$$

Hence, in the case of $\varphi_{\rm ph} \to 0, \, \psi \to 0$ and $A_{\mu\nu} \to A^{\rm (vac)}_{\mu\nu}$ we get

$$\left(\Box - M^2\right)\partial^{\nu}A^{(\text{vac})}_{\mu\nu} = -\sqrt{f}\langle\varphi\rangle_{\text{vac}}\partial^{\nu}F_{\mu\nu}\,,\qquad(11)$$

where from Eq. (3)

$$\partial^{\nu} F_{\mu\nu} = -j_{\mu} - \sqrt{f} \langle \varphi \rangle_{\rm vac} \partial^{\nu} A^{(\rm vac)}_{\mu\nu} \,. \tag{12}$$

Thus,

$$\left(\Box - \tilde{M}^2\right)\partial^{\nu}A^{(\text{vac})}_{\mu\nu} = \sqrt{f}\,\langle\varphi\rangle_{\text{vac}}j_{\mu} \tag{13}$$

with

$$\tilde{M}^2 = M^2 + f \langle \varphi \rangle_{\rm vac}^2 \,. \tag{14}$$

For a Standard-Model point-charge at rest at \vec{x}_0 we have

$$j_{\mu}(\vec{x}) = g_{\mu 0} e_0 \delta^3(\vec{x} - \vec{x}_0) , \qquad (15)$$

and then Eq. (13) gives

$$\partial^{\nu} A^{(\text{vac})}_{\mu\nu}(x) = -g_{\mu0}\sqrt{f} \langle \varphi \rangle_{\text{vac}} \frac{e_0}{4\pi |\vec{x} - \vec{x}_0|} e^{-\tilde{M}|\vec{x} - \vec{x}_0|} \,. \tag{16}$$

Therefore, a hidden-sector additive correction to electrostatic energy of two point-charges at rest at $\vec{x_0}$ and \vec{x}'_0 , following from the first term of weak interaction (1), arises and is equal to

$$\delta V^{(\text{vac})} = -\frac{1}{2} \int d^3x \sqrt{f} \langle \varphi \rangle_{\text{vac}} F^{\mu\nu}(x)' A^{(\text{vac})}_{\mu\nu}(x) = -\int d^3x \sqrt{f} \langle \varphi \rangle_{\text{vac}} A^{\mu}(x)' \partial^{\nu} A^{(\text{vac})}_{\mu\nu}(x), \qquad (17)$$

where $F^{\mu\nu}(x)' = \partial^{\mu}A^{\nu}(x)' - \partial^{\nu}A^{\mu}(x)'$ and

$$A^{\mu}(x)' = g^{\mu 0} \left[\frac{e'_0}{4\pi |\vec{x} - \vec{x}'_0|} + O(f) \right], \qquad (18)$$

the latter is due to Eq. (12) (with $\partial_{\nu}A^{\nu}(x)'=0$) and (15) (with e'_0 and \vec{x}'_0). This implies together with Eq. (16) that

$$\delta V^{(\text{vac})} = f \langle \varphi \rangle_{\text{vac}}^2 \int d^3 x \frac{e_0 \, e'_0}{(4\pi)^2 |\vec{x} - \vec{x}_0| |\vec{x} - \vec{x}'_0|} e^{-\tilde{M} |\vec{x} - \vec{x}_0|} + O\left(f^2\right) = \frac{f \langle \varphi \rangle_{\text{vac}}^2}{\tilde{M}^2} \frac{e_0 \, e'_0}{4\pi |\vec{x}_0 - \vec{x}'_0|} \left(1 - e^{-\tilde{M} |\vec{x}_0 - \vec{x}'_0|}\right) + O\left(f^2\right) .$$
(19)

We can see that such a hidden-sector correction $\delta V^{(\text{vac})}$ to the Coulomb potential $V = e_0 e'_0 / (4\pi |\vec{x}_0 - \vec{x}'_0|)$ is smaller than the latter by the factor

$$\frac{f\langle\varphi\rangle_{\rm vac}^2}{\tilde{M}^2} \left(1 - e^{-\tilde{M}|\vec{x}_0 - \vec{x}_0'|}\right) \,,\tag{20}$$

where $f\langle \varphi \rangle^2_{\text{yac}}/\tilde{M}^2$ is expected to be tiny. This factor approaches $f\langle \varphi \rangle^2_{\text{yac}}$ $|\vec{x}_0 - \vec{x}'_0|/\tilde{M}$ or $f\langle \varphi \rangle^2_{\text{vac}}/\tilde{M}^2$ if $\tilde{M}|\vec{x}_0 - \vec{x}'_0|$ tends to 0 or ∞ , respectively.

3. A suggestion for sterile masses consistent with thermal sterinos

In our model of hidden sector, thermal sterinos are candidates for cold dark matter after their freeze out. Thus, with the use of the abundance of cold dark matter observed by WMAP Collaboration, $\Omega_{\rm DM}h^2 \simeq 0.11$ [4], we infer that in the case of our weak interaction (1) the thermal average of total annihilation cross-section of an antisterino–sterino pair, multiplied by the sterino relative velocity, can be estimated as

$$\left\langle \sigma_{\rm ann} \left(\bar{\psi} \psi \right) 2 v_{\psi} \right\rangle \simeq {\rm pb} \simeq \frac{8}{\pi} \frac{10^{-3}}{{\rm TeV}^2} \,,$$
 (21)

where

$$\sigma_{\rm ann} \left(\bar{\psi} \psi \right) 2v_{\psi} \sim \left[\sigma \left(\bar{\psi} \psi \to \gamma \varphi_{\rm ph} \right) + \sum_{f} \sigma \left(\bar{\psi} \psi \to \bar{f} f \right) \right] 2v_{\psi} \tag{22}$$

with fs denoting charged leptons e^-, μ^-, τ^- as well as quarks u, c, t and d, s, b (among quarks the top t is considered or omitted if $m_t < m_{\psi}$ or $m_t > m_{\psi}$, respectively). Here, the interaction (1) plus the Standard-Model electromagnetic coupling $-j_{\mu}A^{\mu}$ gives (*cf.* the third Ref. [1] and the second Ref. [2])

$$\sigma \left(\bar{\psi} \psi \to \gamma \varphi_{\rm ph} \right) 2v_{\psi} = \frac{1}{6\pi} \left(\frac{f \zeta}{M^2} \right)^2 \left(E_{\psi}^2 + 2m_{\psi}^2 \right) \left(1 - \frac{m_{\varphi}^2}{4E_{\psi}^2} \right)$$
(23)

and

$$\sigma\left(\bar{\psi}\psi\to\bar{f}f\right)2v_{\psi} = \frac{1}{12\pi}\left(\frac{e_f f\,\zeta\langle\varphi\rangle_{\rm vac}}{M^2}\right)^2\frac{E_{\psi}^2+2m_{\psi}^2}{E_{\psi}^2}\,,\tag{24}$$

where the fermion mass m_f is neglected versus sterino mass m_{ψ} . The annihilation processes $\bar{\psi}\psi \to \gamma \varphi_{\rm ph}$ and $\bar{\psi}\psi \to \bar{f}f$ go virtually as $\bar{\psi}\psi \to A^* \to \gamma \varphi_{\rm ph}$ and $\bar{\psi}\psi \to A^* \to \gamma^* \to \bar{f}f$, respectively, where $A^* \to \gamma^*$ involves the action of $\langle \varphi \rangle_{\rm vac}$. Thus,

$$\sigma_{\rm ann}\left(\bar{\psi}\psi\right)2v_{\psi}\sim \left[1+\left(\frac{20}{3}\,\operatorname{for}\,m_{\psi}< m_{t}\,\operatorname{or}\,8\,\operatorname{for}\,m_{\psi}>m_{t}\right)\frac{\sigma\left(\bar{\psi}\psi\rightarrow e^{+}e^{-}\right)}{\sigma\left(\bar{\psi}\psi\rightarrow\gamma\varphi_{\rm ph}\right)}\right]\sigma\left(\bar{\psi}\psi\rightarrow\gamma\varphi_{\rm ph}\right)2v_{\psi}\,.$$

$$(25)$$

Here, $3+3\times 3(4/9+1/9) = 8$ for $m_{\psi} > m_t$ and $3+2\times 3(4/9+1/9)+3(1/9) = 3$ 20/3 for $m_{\psi} < m_t$.

Now, we assume tentatively that

$$m_{\psi}^2 \sim m_{\varphi}^2 \sim \left(10^{-2} \operatorname{to} 1\right) M^2 \sim \langle \varphi \rangle_{\mathrm{vac}}^2$$
 (26)

and conjecture boldly that

$$f \sim e^2 = 4\pi \alpha \simeq 0.0917, \qquad \zeta \sim 1.$$
 (27)

Then, from Eqs. (25) and (26)

$$\sigma_{\rm ann} (\psi \psi) 2v_{\psi} \sim \left[1 + \left(\frac{20}{3} \, \text{for} \, m_{\psi} < m_t \, \text{or} \, 8 \, \text{for} \, m_{\psi} > m_t \right) \frac{2}{3} e^2 \right] \frac{3f^2}{8\pi m_{\psi}^2} \times \left(10^{-4} \, \text{to} \, 1 \right)$$
(28)

with $f^2 = e^4$ due to Eq. (27), where $E_{\psi} \sim m_{\psi}$ and the electron mass m_e is neglected. Note that due to Eq. (28) $\langle \sigma_{\rm ann}(\bar{\psi}\psi)2v_{\psi}\rangle \sim \sigma_{\rm ann}(\bar{\psi}\psi)2v_{\psi}$. From Eqs. (21) and (28) with $f^2 = e^4$ we estimate

$$m_{\psi} \sim (7.5 \text{ to } 770) \text{ GeV}$$
 (29)

and

$$M \sim (10 \text{ to } 1) m_{\psi} \sim (75 \text{ to } 770) \text{ GeV}.$$
 (30)

In particular, for $M \sim 10 m_{\psi} \sim 75$ GeV the numerical coefficient in the factor (20) becomes

$$\frac{f\langle\varphi\rangle_{\rm vac}^2}{\tilde{M}^2} = \frac{f\langle\varphi\rangle_{\rm vac}^2/M^2}{1 + f\langle\varphi\rangle_{\rm vac}^2/M^2} \sim \frac{e^2 \, m_\psi^2/M^2}{1 + e^2 m_\psi^2 M^2} \sim 0.00092\,.$$
(31)

Of course, this estimate is simply a guess, but it is an example consistent with the thermal sterinos playing the role of cold dark matter after their freeze-out. For a value of f smaller than $f \sim e^2$, the coefficient (31) is decreased proportionally to f when f^2/m_{ψ}^2 (due to Eq. (21) and (28)) as well as m_{ψ}^2/M^2 (due to Eqs. (26)) are fixed. The not observed yet hidden-sector additive correction to Coulomb po-

The not observed yet hidden-sector additive correction to Coulomb potential, generated through the photonic portal, might turn out to be either exciting or fatal for the verification of our model of the hidden sector of the Universe.

Thus, if such a tiny correction existed, photons would be coupled not only electromagnetically to the Standard-Model electric current, but also weakly to the hidden sector through the photonic portal (1). Then, the problem of embedding this portal into an electroweakly symmetric form, broken spontaneously by the Standard-Model Higgs mechanism, would be actual (for a proposal cf. the last Ref. [1]). Note finally, that the photonic portal (1) would also imply this tiny correction to the Standard-Model Coulomb potential, when in contrast to the present paper our sterile mediating field A (of dimension one) were gauged (for the gauge option cf. Appendix in the second Ref. [2]).

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