# COHERENT STATES AND QUANTUM NUMBERS FOR TWIST-DEFORMED OSCILLATOR MODEL 

MARCIN DASZKIEWICZ<br>Institute of Theoretical Physics, University of Wrocław<br>pl. Maxa Borna 9, 50-206 Wrocław, Poland<br>marcin@ift.uni.wroc.pl<br>Cezary J. Walczyk<br>Department of Physics, University of Białystok<br>Lipowa 41, 15-424 Białystok, Poland<br>c.walczyk@alpha.uwb.edu.pl

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The coherent states for twist-deformed oscillator model provided in article by M. Daszkiewicz, C.J. Walczyk [Acta Phys. Pol. B 40, 293 (2009)] are constructed. Besides, it is demonstrated that the energy spectrum of considered model is labeled by two quantum numbers - by the so-called main and azimutal quantum numbers respectively.

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The suggestion to use noncommutative coordinates goes back to Heisenberg and was firstly formalized by Snyder in [1]. Recently, there were also found formal arguments based mainly on Quantum Gravity [2, 3] and String Theory models $[4,5]$, indicating that space-time at Planck scale should be noncommutative, i.e. it should have a quantum nature. Consequently, there appeared a lot of papers dealing with noncommutative classical and quantum mechanics (see e.g. [6, 7]) as well as with field theoretical models (see e.g. $[8,9]$ ), in which the quantum space-time is employed.

In accordance with the Hopf-algebraic classification of all deformations of relativistic [10] and nonrelativistic [11] symmetries, one can distinguish three basic types of space-time noncommutativity:

1. The canonical (soft) deformation

$$
\begin{equation*}
\left[x_{\mu}, x_{\nu}\right]=i \theta_{\mu \nu} \tag{1}
\end{equation*}
$$

with constant and antisymmetric tensor $\theta_{\mu \nu}$. The explicit form of corresponding Poincare Hopf algebra has been provided in [12, 13], while its nonrelativistic limit has been proposed in [14].
2. The Lie-algebraic case

$$
\begin{equation*}
\left[x_{\mu}, x_{\nu}\right]=i \theta_{\mu \nu}^{\rho} x_{\rho} \tag{2}
\end{equation*}
$$

with particularly chosen constant coefficients $\theta_{\mu \nu}^{\rho}$. Particular kind of such space-time modification has been obtained as representations of $\kappa$-Poincare [15, 16] and $\kappa$-Galilei [17] Hopf algebras. Besides, the Liealgebraic twist deformations of relativistic and nonrelativistic symmetries have been provided in [18, 19] and [14], respectively.
3. The quadratic deformation

$$
\begin{equation*}
\left[x_{\mu}, x_{\nu}\right]=i \theta_{\mu \nu}^{\rho \tau} x_{\rho} x_{\tau} \tag{3}
\end{equation*}
$$

with constant coefficients $\theta_{\mu \nu}^{\rho \tau}$. Its Hopf-algebraic realization was proposed in [20, 21] and [19] in the case of relativistic symmetry, and in [22] for its nonrelativistic counterpart.
Besides, it has been demonstrated in [23], that in the case of the so-called $N$-enlarged Newton-Hooke Hopf algebras $\mathcal{U}_{0}^{(N)}\left(N H_{ \pm}\right)$, the twist deformation provides the new space-time noncommutativity of the form ${ }^{1,2}$
4.

$$
\begin{equation*}
\left[t, x_{i}\right]=0, \quad\left[x_{i}, x_{j}\right]=i f_{\kappa \pm}(t) \theta_{i j}(x) \tag{4}
\end{equation*}
$$

with time-dependent functions

$$
\begin{aligned}
f_{\kappa+}(t) & =\kappa f\left(\sinh \left(\frac{t}{\tau}\right), \cosh \left(\frac{t}{\tau}\right)\right) \\
f_{\kappa-}(t) & =\kappa f\left(\sin \left(\frac{t}{\tau}\right), \cos \left(\frac{t}{\tau}\right)\right)
\end{aligned}
$$

$\theta_{i j}(x) \sim \theta_{i j}=$ const or $\theta_{i j}(x) \sim \theta_{i j}^{k} x_{k}$ and $\tau$ as well as $\kappa$ denoting the cosmological constant and deformation parameter respectively. It should be also noted that different relations between all mentioned above quantum spaces $1,2,3$ and 4 have been summarized in article [23].

[^0]Let us now turn to the quantum oscillator model defined on the twistdeformed phase space [24] ${ }^{3}$

$$
\begin{equation*}
\left[t, \bar{x}_{i}\right]=0, \quad\left[\bar{x}_{1}, \bar{x}_{2}\right]=i f_{\kappa}(t), \quad\left[\bar{x}_{i}, \bar{p}_{j}\right]=i \hbar \delta_{i j}, \quad\left[\bar{p}_{i}, \bar{p}_{j}\right]=0 . \tag{5}
\end{equation*}
$$

Its dynamic is given by the following Hamiltonian function with constant mass $m$ and frequency $\omega$

$$
\begin{equation*}
\bar{H}(\bar{p}, \bar{x})=\frac{1}{2 m}\left(\bar{p}_{1}^{2}+\bar{p}_{2}^{2}\right)+\frac{1}{2} m \omega^{2}\left(\bar{x}_{1}^{2}+\bar{x}_{2}^{2}\right) . \tag{6}
\end{equation*}
$$

In order to analyze the above system, we represent the noncommutative variables $\left(\bar{x}_{i}, \bar{p}_{i}\right)$ on classical phase space ( $x_{i}, p_{i}$ ) as follows (see e.g. [25, 26])

$$
\begin{equation*}
\bar{x}_{1}=\hat{x}_{1}-\frac{f_{\kappa}(t)}{2 \hbar} \hat{p}_{2}, \quad \bar{x}_{2}=\hat{x}_{2}+\frac{f_{\kappa}(t)}{2 \hbar} \hat{p}_{1} \tag{7}
\end{equation*}
$$

where

$$
\begin{equation*}
\left[\hat{x}_{i}, \hat{x}_{j}\right]=0=\left[\hat{p}_{i}, \hat{p}_{j}\right], \quad\left[\hat{x}_{i}, \hat{p}_{j}\right]=i \hbar \delta_{i j} . \tag{8}
\end{equation*}
$$

Then, the Hamiltonian (6) takes the form ${ }^{4}$

$$
\begin{equation*}
H_{f}(t)=\frac{\left(\hat{p}_{1}^{2}+\hat{p}_{2}^{2}\right)}{2 M_{f}(t)}+\frac{1}{2} M_{f}(t) \Omega_{f}^{2}(t)\left(\hat{x}_{1}^{2}+\hat{x}_{2}^{2}\right)-\frac{f_{\kappa}(t)}{2 \hbar} m \omega^{2} \hat{L}, \tag{9}
\end{equation*}
$$

with symbol

$$
\begin{equation*}
\hat{L}=\hat{x}_{1} \hat{p}_{2}-\hat{x}_{2} \hat{p}_{1} \tag{10}
\end{equation*}
$$

denoting angular momentum of particle. Besides, the coefficients $M_{f}(t)$ and $\Omega_{f}(t)$ present in the above formula denote the time-dependent functions given by

$$
\begin{equation*}
M_{f}(t)=\frac{m}{1+\frac{m^{2} \omega^{2} f_{\kappa}^{2}(t)}{4 \hbar^{2}}}, \quad \Omega_{f}(t)=\omega \sqrt{1+\frac{m^{2} \omega^{2} f_{\kappa}^{2}(t)}{4 \hbar^{2}}}, \tag{11}
\end{equation*}
$$

respectively, such that

$$
\begin{equation*}
M_{f}(t) \Omega_{f}^{2}(t)=m \omega^{2}=\text { const } . \tag{12}
\end{equation*}
$$

Further, we introduce a set of time-dependent creation $\left(a_{A}^{\dagger}(t)\right)$ and annihilation $\left(a_{A}(t)\right)$ operators

$$
\begin{equation*}
\hat{a}_{ \pm}(t)=\frac{1}{2 \sqrt{ } \hbar}\left[\frac{\left(\hat{p}_{1} \pm i \hat{p}_{2}\right)}{\sqrt{M_{f}(t) \Omega_{f}(t)}}-i \sqrt{M_{f}(t) \Omega_{f}(t)}\left(\hat{x}_{1} \pm i \hat{x}_{2}\right)\right] \tag{13}
\end{equation*}
$$

[^1]satisfying the standard commutation relations
\[

$$
\begin{equation*}
\left[\hat{a}_{A}, \hat{a}_{B}\right]=0, \quad\left[\hat{a}_{A}^{\dagger}, \hat{a}_{B}^{\dagger}\right]=0, \quad\left[\hat{a}_{A}, \hat{a}_{B}^{\dagger}\right]=\delta_{A B} ; \quad A, B= \pm \tag{14}
\end{equation*}
$$

\]

Then, one can easily check that in terms of the operators (13) the Hamiltonian function (9) looks as follows

$$
\begin{equation*}
\hat{H}_{f}(t)=\Omega_{+}(t)\left(\hat{N}_{+}(t)+\frac{1}{2}\right)+\Omega_{-}(t)\left(\hat{N}_{-}(t)+\frac{1}{2}\right) \tag{15}
\end{equation*}
$$

with

$$
\begin{equation*}
\Omega_{ \pm}(t)=\Omega_{f}(t) \mp \frac{f_{\kappa}(t) m \omega^{2}}{2 \hbar} \tag{16}
\end{equation*}
$$

and number operators in $\pm$ direction given by

$$
\begin{equation*}
\hat{N}_{ \pm}(t)=\hat{a}_{ \pm}^{\dagger}(t) \hat{a}_{ \pm}(t) \tag{17}
\end{equation*}
$$

respectively. Moreover, we see that the energy eigenvectors can be generated in a standard way as follows

$$
\begin{equation*}
\left|n_{+}, n_{-}, t\right\rangle=\frac{1}{\sqrt{n_{+}!}} \frac{1}{\sqrt{n_{-}!}}\left(\hat{a}_{+}^{\dagger}(t)\right)^{n_{+}}\left(\hat{a}_{-}^{\dagger}(t)\right)^{n_{-}}|0\rangle \tag{18}
\end{equation*}
$$

while the corresponding (parameterized by $n_{+}$and $n_{-}$) eigenvalues are

$$
\begin{equation*}
E_{n_{+}, n_{-}}(t)=\Omega_{+}(t)\left(n_{+}+\frac{1}{2}\right)+\Omega_{-}(t)\left(n_{-}+\frac{1}{2}\right), \quad n_{+}, n_{-}=0,1,2, \ldots \tag{19}
\end{equation*}
$$

Besides, using operator representation (13), one finds

$$
\begin{equation*}
\left(\Delta \hat{x}_{i}\right)_{\left|n_{+}, n_{-}, t\right\rangle}^{2}\left(\Delta \hat{p}_{i}\right)_{\left|n_{+}, n_{-}, t\right\rangle}^{2}=\frac{\hbar^{2}}{4}\left(1+n_{+}+n_{-}\right)^{2} \tag{20}
\end{equation*}
$$

where symbol $(\Delta \hat{a})_{|\varphi\rangle}$ denotes the uncertainty of observable $\hat{a}$ in quantum state $|\varphi\rangle$. The above result means that momentum-position uncertainty relations for eigenstates (18) become saturated only for $n_{+}=n_{-}=0$, i.e. only for vacuum vector $|0\rangle$. Apart from that, it is easy to see that the momentum operator (10) can be written as follows

$$
\begin{equation*}
\hat{L}=\hbar\left(\hat{a}_{-}^{\dagger}(t) \hat{a}_{-}(t)-\hat{a}_{+}^{\dagger}(t) \hat{a}_{+}(t)\right) \tag{21}
\end{equation*}
$$

while its action on quantum states (18) is given by

$$
\begin{equation*}
\hat{L}\left|n_{+}, n_{-},\right\rangle=\hbar\left(n_{-}-n_{+}\right)\left|n_{+}, n_{-}, t\right\rangle \tag{22}
\end{equation*}
$$

Consequently, the energy spectrum (19) can be written in terms of eigenvalues (22) as follows

$$
\begin{equation*}
E_{n_{+}, n_{-}}(t)=\hbar \Omega_{f}(t)\left(n_{+}+n_{-}+1\right)+\frac{f_{\kappa}(t) M_{f}(t) \Omega_{f}^{2}(t)}{2}\left(n_{-}-n_{+}\right) \tag{23}
\end{equation*}
$$

Let us now solve two problems. First of them concerns the construction of the so-called coherent states for considered model, i.e. the quantum vectors which saturate the momentum-position Heisenberg uncertainty relations. The second problem applies to the proper interpretation of quantum numbers $n=n_{+}+n_{-}$and $l=n_{-} n_{+}$labeling the energy spectrum (23).

Hence, let us consider the quantum states of the form

$$
\begin{equation*}
\left|c_{+}, c_{-}, t\right\rangle=\sum_{n_{+}, n_{-}} \frac{c_{+}^{n_{+}} e^{-\frac{1}{2}\left|c_{+}\right|^{2}}}{\sqrt{n_{+}!}} \frac{c_{-}^{n_{-}} e^{-\frac{1}{2}\left|c_{-}\right|^{2}}}{\sqrt{n_{-}!}}\left|n_{+}, n_{-}, t\right\rangle \tag{24}
\end{equation*}
$$

which play the role of the eigenfunctions for annihilation operators (13)

$$
\begin{equation*}
\hat{a}_{ \pm}(t)\left|c_{+}, c_{-}, t\right\rangle=c_{ \pm}\left|c_{+}, c_{-}, t\right\rangle . \tag{25}
\end{equation*}
$$

By direct calculation, one may check that

$$
\begin{equation*}
\left(\Delta p_{i}\right)_{\left|c_{+}, c_{-}, t\right\rangle}^{2}=\frac{\hbar M_{f}(t) \Omega_{f}(t)}{2}, \quad\left(\Delta x_{i}\right)_{\left|c_{+}, c_{-}, t\right\rangle}^{2}=\frac{1}{2} \frac{\hbar}{M_{f}(t) \Omega_{f}(t)}, \quad i=1,2 \tag{26}
\end{equation*}
$$

what leads to the saturated momentum-position Heisenberg uncertainty relations

$$
\begin{equation*}
\left(\Delta p_{i}\right)_{\left|c_{+}, c_{-}, t\right\rangle}^{2}\left(\Delta x_{i}\right)_{\left|c_{+}, c_{-}, t\right\rangle}^{2}=\frac{\hbar^{2}}{4}, \quad i=1,2 \tag{27}
\end{equation*}
$$

Consequently, we see that the vectors (24) are, in fact, nothing else than the coherent states for twist-deformed oscillator model, satisfying

$$
\begin{equation*}
\left\langle\hat{H}_{f}\right\rangle_{\left|c_{+}, c_{-}, t\right\rangle}=E_{|0,0, t\rangle}(t)+\frac{\Omega_{f}(t)}{\hbar}(\Delta L)_{\left|c_{+}, c_{-}, t\right\rangle}^{2}+\frac{M_{f}(t) \Omega_{f}^{2}(t) f_{\kappa}(t)}{2 \hbar}\langle L\rangle_{\left|c_{+}, c_{-}, t\right\rangle}, \tag{28}
\end{equation*}
$$

with

$$
\begin{align*}
\langle L\rangle_{\left|c_{+}, c_{-}, t\right\rangle} & =\hbar\left(\left|c_{-}\right|^{2}-\left|c_{+}\right|^{2}\right)  \tag{29}\\
(\Delta L)_{\left|c_{+}, c_{-}, t\right\rangle}^{2} & =\hbar^{2}\left(\left|c_{-}\right|^{2}+\left|c_{+}\right|^{2}\right) \tag{30}
\end{align*}
$$

In the case of second problem, one should solve the eigenvalue equation for Hamiltonian (9) written in terms of polar coordinates

$$
\begin{equation*}
\hat{H}_{f}(t) \psi(r, \varphi, t)=E(t) \psi(r, \varphi, t) \tag{31}
\end{equation*}
$$

where

$$
\begin{align*}
\hat{H}_{f}(t)= & -\frac{\hbar^{2}}{2 M_{f}(t)}\left(\frac{\partial^{2}}{\partial r^{2}}+\frac{1}{r} \frac{\partial}{\partial r}-\frac{1}{\hbar^{2}} \frac{\hat{L}^{2}}{r^{2}}\right) \\
& +\frac{M_{f}(t) \Omega_{f}^{2}(t)}{2} r^{2}-\frac{f_{\kappa}(t) M_{f}(t) \Omega_{f}^{2}(t)}{2 \hbar} \hat{L} \tag{32}
\end{align*}
$$

and

$$
\begin{equation*}
\hat{L}=-i \hbar \frac{\partial}{\partial \varphi}, \quad[\hat{H}, \hat{L}]=0 \tag{33}
\end{equation*}
$$

To this aim, it is convenient to take the corresponding eigenfunctions in the form

$$
\begin{equation*}
\psi(r, \varphi, t)=\phi(\varphi) R(r, t) \tag{34}
\end{equation*}
$$

with its azimutal part $\phi(\varphi)$ satisfying

$$
\begin{equation*}
\hat{L} \phi_{l}(\varphi)=\hbar l \phi(\varphi), \quad \phi_{l}(\varphi)=\frac{1}{\sqrt{2 \pi}} e^{i l \varphi}, \quad l=0, \pm 1, \pm 2, \ldots \tag{35}
\end{equation*}
$$

Then, the proper equation for radial function $R(r, t)$ looks as follows

$$
\begin{align*}
& \left(-\frac{\partial^{2}}{\partial \rho^{2}}-\frac{1}{\rho} \frac{\partial}{\partial \rho}+\frac{l^{2}}{\rho^{2}}+\frac{\rho^{2}}{4}-\mathcal{E}_{l}(t)\right) R_{l}(\rho(t))=0 \\
& \mathcal{E}_{l}(t)=\frac{E(t)-\frac{f_{\kappa}(t) M_{f}(t) \Omega_{f}^{2}(t)}{2} l}{\hbar \Omega_{f}(t)} \tag{36}
\end{align*}
$$

where $\rho(t)=r \sqrt{2 M_{f}(t) \Omega_{f}(t) / \hbar}$ plays the role of dimensionless variable. Its physical solution can be written as

$$
\begin{equation*}
R_{l}^{(n)}(\rho(t))=w_{l}^{(n)}(\rho(t)) e^{-\rho^{2}(t) / 4} \tag{37}
\end{equation*}
$$

with $w_{l}^{(n)}(\rho(t))$ denoting the polynomial of degree $n$. Then, equation (36) reduces to the following one

$$
\begin{align*}
& -\frac{\partial^{2} w_{l}^{(n)}(\rho(t))}{\partial \rho^{2}}+\frac{\rho^{2}-1}{\rho} \frac{\partial w_{l}^{(n)}(\rho(t))}{\partial \rho} \\
& +\frac{l^{2}}{\rho^{2}} w_{l}^{(n)}(\rho(t))-\left(\mathcal{E}_{l}(t)-1\right) w_{l}^{(n)}(\rho(t))=0 \tag{38}
\end{align*}
$$

for which the solution (this time) is given by ${ }^{5}$

$$
\begin{equation*}
w_{l}^{(n)}(\rho(t))=a_{l}^{(n)}\left(1+\sum_{k=1}^{(n-|l|) / 2}\left[\prod_{s=1}^{k} \frac{n+2-(2 s+|l|)}{l^{2}-(2 s+|l|)^{2}}\right] \rho^{2 k}(t)\right) \rho^{|l|}(t) \tag{39}
\end{equation*}
$$

[^2]only when
$\mathcal{E}_{l}(t) \rightarrow \mathcal{E}_{l}^{(n)}(t)=n+1, \quad l \in\{-n,-n+2, \ldots, n-2, n\}, \quad n=0,1,2,3, \ldots$,
or (equivalently)
\[

$$
\begin{equation*}
E_{l}^{(n)}(t)=\hbar \Omega_{f}(t)(n+1)+\frac{f_{\kappa}(t) M_{f}(t) \Omega_{f}^{2}(t)}{2} l \tag{41}
\end{equation*}
$$

\]

Consequently, after substitution $n=n_{+}+n_{-}$and $l=n_{-} n_{+}$into eigenvalues (41), we get, in fact, the energy spectrum (23) labeled by $n_{+}$and $n_{-}$parameters. For this reason as well as due to the formulas (35), (37) and (41), the quantities $n$ and $l$ may be called the "main" and "azimutal" quantum numbers respectively.

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[^0]:    ${ }^{1} x_{0}=c t$.
    ${ }^{2}$ The discussed space-times have been defined as the quantum representation spaces, the so-called Hopf modules (see e.g. [12, 13]), for quantum $N$-enlarged Newton-Hooke Hopf algebras.

[^1]:    ${ }^{3}$ See type 4 of quantum space-time.
    ${ }^{4}$ It should be noted that for $f_{\kappa}(t)=\theta$, we get the canonically deformed oscillator model provided in [26].

[^2]:    ${ }^{5}$ The symbol $a_{l}^{(n)}$ denotes the normalization factor.

