

COMMISSIONING OF THE INNER TRACKER OF THE KLOE-2 EXPERIMENT

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The KLOE-2 experiment is undergoing commissioning at the DAΦNE e^+e^- collider of the Frascati National Laboratory of the INFN, after the integration of new detectors in the former KLOE apparatus. The Inner Tracker, a very light detector (material budget $< 2\%X_0$), is one of the new subdetectors and it is composed of 4 cylindrical triple-GEM with a stereo XV-strips/pads readout. The commissioning phase of the Inner Tracker consists in the characterization of the detector response and in its performance evaluation. The method used to evaluate its detection efficiency is reported, together with some preliminary results.

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1. Introduction

After the integration of the new subdetectors [1–5] in the KLOE apparatus [6], the commissioning phase of the KLOE-2 Inner Tracker (IT) has started. This detector is composed of 4 cylindrical triple-GEM coaxial layers [7], equipped each with a double-view XV-strips/pads readout circuit. The dimension of the active area ($\sim 700 \times 300 \text{ mm}^2$) pushed the GEM deliverer (CERN TE-MPE-EM workshop), in collaboration with INFN-LNF groups, to proceed with a new manufacturing technique [8], while the assembly procedure has been totally developed at the Frascati National Laboratory (LNF) of the INFN. The Inner Tracker has been installed between the beam pipe and the inner wall of the KLOE Drift Chamber (DC) and will increase the acceptance for low transverse momentum tracks and improve charged vertex reconstruction (especially useful for events with more than

two tracks originating from Interaction Point [9]). The whole KLOE-2 [10] apparatus works in a 0.52 T magnetic field. Although each layer has been tested in a dedicated area at LNF, before their integration in the IT, new tests with the detector in its final position are required. The IT commissioning benefits from the presence of the DC and its excellent tracking performance.

2. Track reconstruction with the Inner Tracker

Starting from DC hits, tracks are reconstructed including IT hits by using the Kalman filter. Inner Tracker must be aligned and calibrated in order to get the best tracking performance. The presence of the B-field influences the reconstruction since the signal electron cloud experiences the Lorentz force

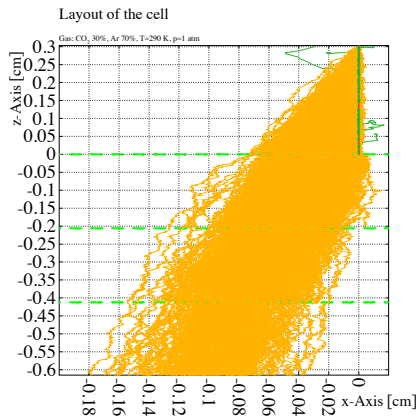


Fig. 1. GARFIELD simulation of electrons drift in a 0.52 T magnetic field with Ar:CO₂ 70 : 30 gas mixture: the track is at the position $x = 0$.

and, therefore, there is a shift in the position of the fired readout strips with respect to the true position of the intersection between the particle track and the readout plane, Fig. 1. Spatial resolution is also affected by non-radial tracks crossing the detector as shown in Fig. 2 [7]. The combination of the two effects produces a focusing or a defocusing of the electron cloud according to the impact parameter of the track on the cylindrical detectors. These effects must be studied independently: cosmic-ray muon data will be used to evaluate the non-radial correction (without B-field) and the magnetic field influence (with B-field), and Bhabha scattering events will be used to check the effect of the two corrections. The calibration and alignment procedure are presently ongoing. In the next paragraphs, the monitoring of the detector status and the efficiency evaluation will be described.

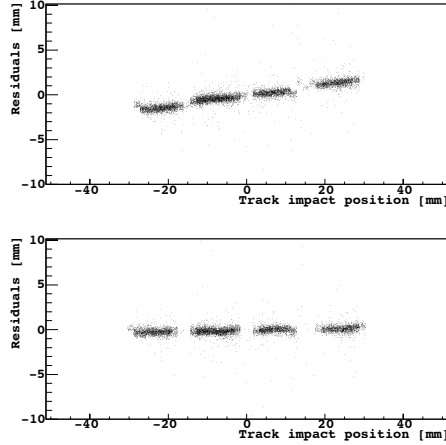


Fig. 2. Non-radial tracks effect on the residuals measured with a cylindrical GEM prototype.

3. Noisy and inefficient strips

The detector status must be monitored in time and taken into account at the analysis stage. The definition of *noisy* or *dead* strips is related to their average occupancy for a single view and for each layer of the Inner Tracker. The technique used to evaluate the average occupancy and the definition of the strips status is briefly reported.

X-view. The X-view is given by the axial strips which provide coordinates in the r - φ plane (transverse plane with respect to the beam line). All these strips have the same length and width (and, consequently, the same parasitic capacitance) so that, assuming a uniform exposition to cosmic-ray muons, the profile of the occupancy distribution is expected to be a Gaussian function. In Fig. 3 the occupancy for the X-view strips is shown,

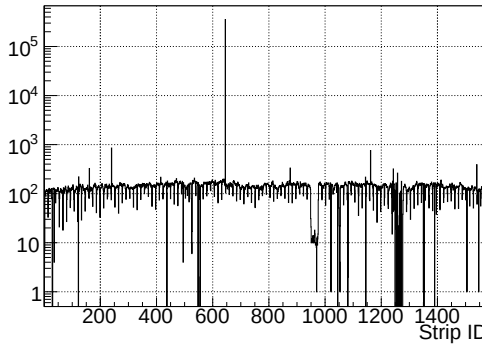


Fig. 3. Layer 2 X-view occupancy distribution.

while its profile distribution is displayed in Fig. 4. This latter is fitted with a combination of a second-degree polynomial (whose parameters are p_0 , p_1 and p_2) and a Gaussian function. The mean value, $\bar{\xi}$, of the Gaussian fit is considered to be the average occupancy of the strips belonging to the layer under study, while the sigma, σ , is used to identify the strips status as follows: let ξ_i be the occupancy of the i^{th} strip, then strips with $\xi_i = 0$ will be considered *dead*, while strips with $\xi_i > \bar{\xi} + 5\sigma$ will be considered *noisy*. The same definitions are applied to V-view strips, although an intermediate step is required as further explained.

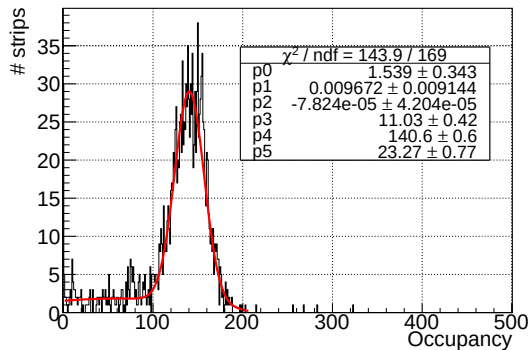


Fig. 4. Layer 2 X-view occupancy profile.

V-view. V-view strips have different lengths and, thus, a non-uniform occupancy distribution. In order to use the tool previously described, it seems natural to normalize the occupancy of the i^{th} strip to its length and then use the same definitions as for the axial strips.

The list of noisy strips is then recorded in the database in order to mask them directly at the DAQ initialization stage for data taking and to monitor the time evolution of the noisy and dead strips. From a reference cosmic-ray run, we have $\sim 6\%$ of noisy strips and $\sim 17\%$ of dead strips considering both views and all four layers.

4. Inner Tracker efficiency

The Inner Tracker efficiency is evaluated by using cosmic-ray muon tracks reconstructed in the Drift Chamber in 0.52 T magnetic field.

Normalization sample. Tracks reconstructed in the DC are extrapolated to the Inner Tracker (IT) assuming straight lines, approximation valid if tracks with transverse momentum $p_T > 500$ MeV are selected. Two crossing points with each IT layer and a Point of Closest Approach (PCA) of the track to the beam-line are required, with $z_{\text{PCA}} < 35$ cm and $R_{\text{PCA}} < 5$ cm, the latter

being the PCA radius in the bending plane. The two crossing points represents the track expected positions on the IT and their angular φ -position distribution in the bending plane will be the normalization sample.

Efficiency evaluation. The efficiency is evaluated as the ratio between the expected and the measured φ -position distributions. The latter distribution is obtained by using an algorithm which produces the 3-dimensional coordinates of the detected clusters. With the position of the reconstructed clusters, we can compute the φ -position in the bending plane. The numerator of the efficiency is, therefore, the distribution of the measured clusters in the IT (Fig. 5). A normalization sample of about 6×10^6 events in which

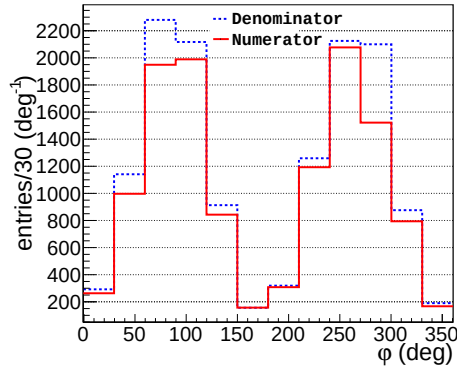


Fig. 5. Layer 2 efficiency as a function of the φ -position in the bending plane.

at least a reconstructed track is present has been used to evaluate efficiency. Selection criteria reduced the normalization sample to $\sim 3.4 \times 10^3$ events, while $\sim 3.3 \times 10^3$ events enter the numerator. As an example, in Fig. 6 (top) the efficiency for layer 2 as a function on the φ -position is shown. The low-efficiency values in some of the φ -regions, as reported in Fig. 6 (bottom), are

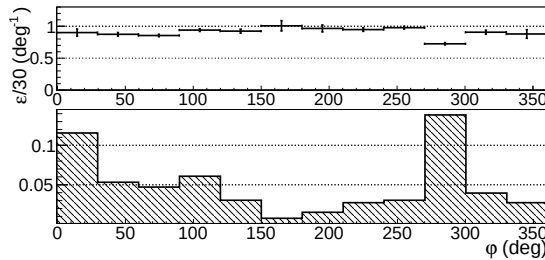


Fig. 6. Top: Layer 2 efficiency as a function of the φ -position in the bending plane. Bottom: Fraction of masked strips as a function of the φ -position in the bending plane.

correlated to the presence of a non-negligible fraction of dead/noisy strips, which are masked for evaluating efficiency. The fraction is computed as the ratio between the number of masked strips per bin and the total number of strips in that bin. The present status is quite similar for all the layers.

5. Conclusions

The commissioning of the Inner Tracker for the KLOE-2 experiment is in progress. The detector operational condition is under optimization and tools to monitor the detector status and measure its efficiency have been developed. Preliminary efficiency measurements have been reported. Calibration and alignment procedures are under development and will use both cosmic-ray muon and Bhabha scattering events.

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