MASSLESS FOUR-LOOP FORM FACTORS*

MATTHIAS STEINHAUSER

Institut für Theoretische Teilchenphysik, Karlsruhe Institute of Technology (KIT) 76128 Karlsruhe, Germany

(Received November 8, 2017)

In this contribution we briefly discuss the status of the calculation of the four-loop massless form factors.

DOI:10.5506/APhysPolB.48.2167

1. Introduction

In QCD, form factors are building blocks for many physical quantities, most prominently for Higgs boson production and the Drell–Yan process. Furthermore, they are the simplest infra-red divergent objects and are thus used to study general properties of QCD and to develop a deeper insight into the infra-red properties.

In this contribution, we consider the massless photon quark vertex, Γ_q^{μ} and QCD corrections up to four loops to the corresponding form factor which is obtained via

In this formula, $D = 4 - 2\epsilon$ is the space-time dimension, p_1 and p_2 are the incoming momenta of the quark and anti-quark, respectively, and $q = p_1 + p_2$ is the photon momentum.

Three-loop corrections to $F_q(q^2)$ have been computed for the first time in Ref. [1] and afterwards confirmed in [2]. In the final expression presented in these references, the highest ϵ coefficients of the three most complicated master integrals were only known numerically. Their analytic expressions have been provided in [3].

First results at four-loop order have been obtained in Refs. [4, 5] where all planar contributions to F_q have been computed. This provides a complete

^{*} Presented at the XLI International Conference of Theoretical Physics "Matter to the Deepest", Podlesice, Poland, September 3–8, 2017.

result in the large- N_c limit. Note that the corrections with three closed quark loops have been computed in Ref. [6]. Let us also mention the work [7], where the $1/\epsilon^2$ pole of the four-loop form factor within $\mathcal{N} = 4$ super Yang–Mills theory has been computed using numerical methods.

In this contribution, we concentrate on the fermionic terms to F_q with two closed fermion lines. This part receives planar and non-planar contributions which have been computed in Ref. [8].

In the next section, we briefly comment on the individual steps of the calculation and present analytic results in Section 3.

2. Calculation

Sample Feynman diagrams which contribute to the massless four-loop form factor are shown in Fig. 1. The first two diagrams represent purely gluonic planar contributions, which have been computed in Refs. [4, 5] and the last one represents one of the most complicated non-planar diagrams with two (massless) fermion insertions. A crucial step in the course to compute these diagrams is the identification of a (scalar) integral family for each Feynman amplitude. For the planar contributions, there are 38 families which have been classified in Refs. [4, 5]. In general, the non-planar families are more complicated and have not yet been identified by our group. However, in the case of the $n_{\rm f}^2$ terms, there are only two non-planar families which are shown in Fig. 2. Note that for the $n_{\rm f}^2$ amplitudes, many lines of the families in Fig. 2 are absent. However, in Ref. [8], all master integrals which belong to the two families have been computed. It is expected that all of them are needed for the $n_{\rm f}$ -independent four-loop form factor.



Fig. 1. Sample Feynman diagrams. For the $n_{\rm f}^2$ contribution, also non-planar contributions are considered.

Once the integral families are classified, the input files for FIRE [9, 10] and LiteRed [11, 12] can be generated which perform the reduction to master integrals. Then the main task is the analytic evaluation of the latter. In the following, we briefly summarize the basic ideas which enter this part of the calculation.

— Differential equations are a powerful tool to compute master integrals [13–16]. Since we have a one-scale problem at hand and thus



Fig. 2. Non-planar families needed for the $n_{\rm f}^2$ contribution.

no dimensionless ratio can be formed, a further mass scale has to be introduced in order to be able to use the differential equation method. One possible choice is $p_2^2 = xq^2 \neq 0$. Note that for x = 1, the integrals become massless two-point integrals which are available in the literature [17–19]. We will use this point for fixing the initial conditions. For the form factor, we need the master integrals for x = 0.

A non-vanishing q_2^2 increases the number of master integrals. Furthermore, a more complicated reduction problem (with two dimensionful scales) has to be solved.

— A further crucial ingredient is the choice of basis. It has been suggested in [16] (see also Ref. [20]) to choose a so-called canonical basis where the differential equation for the master integrals has the form of

$$g'(x,\epsilon) = \epsilon A(x) \cdot g(x,\epsilon), \qquad (2)$$

where the matrix A(x) is ϵ -independent and only has simple poles of the form $1/(x - x_i)$. In our case, we only have two such terms with $x_0 = 0$ and $x_1 = 1$.

We transform our primary basis (chosen by FIRE) to a canonical basis with the help of the algorithm presented in Ref. [21].

— It is simple to obtain the general solution of Eq. (2) in terms of Harmonic Polylogarithms (HPL) [22, 23]. Afterwards, we fix the boundary conditions for x = 1 using the analytic results from [18] and obtain analytic results for the (two-scale) master integrals. Next, the limit $x \to 0$ has to be considered. This is more complicated than one might expect since this limit has to be taken in the "naive" sense. This means that terms like $x^{k\epsilon}$ should be kept unexpanded and then terms with k = 0 have to be selected. Thus, simply expanding the analytic result does not work. Instead, we go back to the differential equation in Eq. (2). For $x \to 0$, it takes the form of

$$g'(x,\epsilon) = \epsilon \frac{a}{x} \cdot g(x,\epsilon),$$
 (3)

where all matrix elements of a are just numbers. The solution of this equation reads

$$g_{x\to 0} = x^{\epsilon a} h(\epsilon) \,, \tag{4}$$

where $x^{\epsilon a}$ is a matrix where each element is a linear combination of $x^{k\epsilon}$ with $k \leq 0$. The column vector $h(\epsilon)$ is obtained by comparing the original analytic results and $g_{x\to 0}$ after a Taylor expansion for $x \to 0$.

— The sought-after master integrals in the primary basis are obtained by substituting $g_{x\to 0}|_{x^{0\epsilon}}$ and taking the limit $x \to 0$.

3. Some selected results

After inserting the analytic results for the master integrals into the amplitude, we expand in ϵ and obtain results for the form factor up to the finite term. From the $1/\epsilon^2$ pole, it is possible to extract four-loop results for the cusp anomalous dimension

$$\gamma_{\rm cusp} = \sum_{n \ge 0} \left(\frac{\alpha_{\rm s}(\mu^2)}{4\pi} \right)^n \gamma_{\rm cusp}^n \,, \tag{5}$$

which reads

$$\gamma_{\rm cusp}^{3} = \left(\frac{128\pi^{2}\zeta_{3}}{9} + 224\zeta_{5} - \frac{44\pi^{4}}{27} - \frac{16252\zeta_{3}}{27} + \frac{13346\pi^{2}}{243} - \frac{39883}{81}\right) N_{\rm c}^{2} n_{\rm f} \\ + \left(-32\zeta_{3}^{2} - \frac{176\pi^{2}\zeta_{3}}{9} + \frac{20992\zeta_{3}}{27} - 352\zeta_{5} - \frac{292\pi^{6}}{315} + \frac{902\pi^{4}}{45} \right) \\ - \frac{44416\pi^{2}}{243} + \frac{84278}{81}\right) N_{\rm c}^{3} + n_{\rm f}^{2} \left[C_{\rm A}\left(\frac{2240\zeta_{3}}{27} - \frac{56\pi^{4}}{135} - \frac{304\pi^{2}}{243} + \frac{923}{81}\right) + C_{\rm F}\left(-\frac{640\zeta_{3}}{9} + \frac{16\pi^{4}}{45} + \frac{2392}{81}\right)\right] + \left(\frac{64\zeta_{3}}{27} - \frac{32}{81}\right) n_{\rm f}^{3}.$$

$$(6)$$

Note that in this expression, the $n_{\rm f}^3$ and $n_{\rm f}^2$ terms are complete and contain the colour factors $C_{\rm A}$ and $C_{\rm F}$. On the other hand, for the linear $n_{\rm f}$ and the $n_{\rm f}$ -independent terms, only the leading $N_{\rm c}$ terms are shown. The $n_{\rm f}^3$ terms have already been computed in Refs. [24, 25] and all other contributions in Eq. (6) have been obtained independently in Refs. [26, 27]. Numerical results for all other colour coefficients can be found in Ref. [27].

Analytic results for the collinear anomalous dimension and the finite part of the form factor can be found in Refs. [4, 5, 8].

I would like to thank the organizers for the kind invitation and the pleasant atmosphere at the workshop. Furthermore, I thank Johannes Henn, Roman Lee, Alexander Smirnov and Vladimir Smirnov for a fruitful collaboration on the subject presented in this contribution. This work was supported by DFG through the project "Infrared and threshold effects in QCD".

REFERENCES

- [1] P.A. Baikov et al., Phys. Rev. Lett. 102, 212002 (2009)
 [arXiv:0902.3519 [hep-ph]].
- [2] T. Gehrmann *et al.*, J. High Energy Phys. 1006, 094 (2010)
 [arXiv:1004.3653 [hep-ph]].
- [3] R.N. Lee, V.A. Smirnov, J. High Energy Phys. 1102, 102 (2011) [arXiv:1010.1334 [hep-ph]].
- [4] J.M. Henn, A.V. Smirnov, V.A. Smirnov, M. Steinhauser, J. High Energy Phys. 1605, 066 (2016) [arXiv:1604.03126 [hep-ph]].
- [5] J. Henn et al., J. High Energy Phys. 1703, 139 (2017)
 [arXiv:1612.04389 [hep-ph]].
- [6] A. von Manteuffel, R.M. Schabinger, *Phys. Rev. D* 95, 034030 (2017) [arXiv:1611.00795 [hep-ph]].
- [7] R.H. Boels, T. Huber, G. Yang, *Phys. Rev. Lett.* 119, 201601 (2017) [arXiv:1705.03444 [hep-th]].
- [8] R.N. Lee, A.V. Smirnov, V.A. Smirnov, M. Steinhauser, *Phys. Rev. D* 96, 014008 (2017) [arXiv:1705.06862 [hep-ph]].
- [9] A.V. Smirnov, V.A. Smirnov, Comput. Phys. Commun. 184, 2820 (2013) [arXiv:1302.5885 [hep-ph]].
- [10] A.V. Smirnov, Comput. Phys. Commun. 189, 182 (2015)
 [arXiv:1408.2372 [hep-ph]].
- [11] R.N. Lee, arXiv:1212.2685 [hep-ph].
- [12] R.N. Lee, J. Phys.: Conf. Ser. 523, 012059 (2014)
 [arXiv:1310.1145 [hep-ph]].
- [13] A.V. Kotikov, *Phys. Lett. B* **254**, 158 (1991).
- [14] E. Remiddi, Nuovo Cim. A 110, 1435 (1997) [arXiv:hep-th/9711188].

- [15] T. Gehrmann, E. Remiddi, Nucl. Phys. B 580, 485 (2000)
 [arXiv:hep-ph/9912329].
- [16] J.M. Henn, Phys. Rev. Lett. 110, 251601 (2013) [arXiv:1304.1806 [hep-th]].
- [17] P.A. Baikov, K.G. Chetyrkin, *Nucl. Phys. B* 837, 186 (2010)
 [arXiv:1004.1153 [hep-ph]].
- [18] A.V. Smirnov, M. Tentyukov, Nucl. Phys. B 837, 40 (2010) [arXiv:1004.1149 [hep-ph]].
- [19] R.N. Lee, A.V. Smirnov, V.A. Smirnov, Nucl. Phys. B 856, 95 (2012)
 [arXiv:1108.0732 [hep-th]].
- [20] J.M. Henn, J. Phys. A 48, 153001 (2015) [arXiv:1412.2296 [hep-ph]].
- [21] R.N. Lee, J. High Energy Phys. 1504, 108 (2015)
 [arXiv:1411.0911 [hep-ph]].
- [22] E. Remiddi, J.A.M. Vermaseren, Int. J. Mod. Phys. A 15, 725 (2000) [arXiv:hep-ph/9905237].
- [23] D. Maitre, Comput. Phys. Commun. 174, 222 (2006)
 [arXiv:hep-ph/0507152].
- [24] J.A. Gracey, *Phys. Lett. B* **322**, 141 (1994) [arXiv:hep-ph/9401214].
- [25] M. Beneke, V.M. Braun, Nucl. Phys. B 454, 253 (1995)
 [arXiv:hep-ph/9506452].
- [26] J. Davies et al., Nucl. Phys. B 915, 335 (2017) [arXiv:1610.07477 [hep-ph]].
- [27] S. Moch et al., J. High Energy Phys. 1710, 041 (2017)
 [arXiv:1707.08315 [hep-ph]].