## QUARK–NUCLEAR HYBRID EoS WITH EXCLUDED-VOLUME EFFECTS\*

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We employ a new quark–nuclear matter hybrid equation of state that takes into account both quark Pauli blocking in hadronic matter and screening of interactions in quark matter as excluded-volume effects. We obtain mass-radius relations for hybrid stars that fulfill the  $2M_{\odot}$  constraint while exhibiting the high-mass twin phenomenon.

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In the recent work [1], the question of whether a phase-transition from hadronic to deconfined quark matter occurs in the interior of compact stars has been considered within a new approach to the hybrid equation of state (EoS) within the two-phase description. For the hadronic phase, the wellknown DD2 model, developed by Typel *et al.* [2], was implemented with the excluded volume modification at supersaturation densities. For the quark matter phase, the string-flip model (SFM) [3,4] was revived, and the confining interactions between colored constituents were included by a densitydependent mean-field approximation. At low densities, the quark mass diverges, thus emulating confinement by the exclusion of the quark degrees of freedom. As the density increases, the Pauli-blocking effects lead first to a

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strong repulsion and then to the delocalization of the quark wave function, driving the system into a homogeneous quark matter phase. As high-mass twins (HMT), we denote two compact stars that have the same mass (including the presently best known high mass of pulsar PSR J0348+0432 with  $M = 2.01 \pm 0.04 M_{\odot}$  [5]), but very distinct radii. In order to support this phenomenon, the EoS of the compact star matter must have a strong first-order phase transition, while not violating causality, as seen in figure 1. The observation of HMTs would support a CEP in the QCD phase diagram.

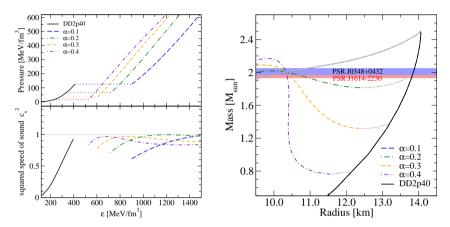


Fig. 1. (Left) Pressure versus energy density for different values of the available volume parameter  $\alpha$  in the quark matter phase. The hadronic phase is shown with a transition to the quark phase corresponding to the Maxwell-construction. The lower panel shows the corresponding squared speed of sound that shall be limited from above by the causality constraint to the squared speed of light  $c_{\rm s}^2 < 1$ . (Right) The M-R relations corresponding to the EoS curves in the left panel. The solid gray lines represent the unstable configurations of hybrid stars.

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