

# THE LIGHT SCALAR $K_0^*(700)$ IN THE VACUUM AND AT NONZERO TEMPERATURE\*

FRANCESCO GIACOSA

Institute of Physics, Jan Kochanowski University, 25-406 Kielce, Poland  
and

Institute for Theoretical Physics, Goethe University  
60438 Frankfurt am Main, Germany

(Received January 24, 2019)

There is mounting evidence toward the existence of a light scalar kaon  $\kappa \equiv K_0^*(700)$  with quantum numbers  $I(J^P) = 1/2(0^+)$ . Here, we recall the results of an effective model with both derivative and non-derivative terms in which only one scalar kaonic field is present in the Lagrangian (the standard quark–antiquark “seed” state  $K_0^*(1430)$ ): a second “companion” pole  $K_0^*(700)$  emerges as a dynamically generated state. A related question is the role of  $K_0^*(700)$  at nonzero  $T$ : since it is the lightest scalar strange state, one would naively expect that it is relevant for  $\pi$  and  $K$  multiplicities. However, a repulsion in the  $\pi K$  channel with  $I = 3/2$  cancels its effect.

DOI:10.5506/APhysPolBSupp.12.283

## 1. Introduction

The lightest scalar kaonic state listed in the PDG [1] is  $K_0^*(700)$  (previously called  $K_0^*(800)$ , see PDG 2016 [2] and older versions). This state, sometimes called  $\kappa$ , still “*needs confirmation*”, but many works do find a pole in that energy region, see Ref. [3] and references therein. The PDG reports at present the following result:

$$\text{pole } \kappa[\text{PDG}] : (630\text{--}730) - i(260\text{--}340) \text{ MeV}, \quad (1)$$

(hence, the pole width lies between 520–680 MeV), while the Breit–Wigner (BW) mass and widths are

$$\text{BW [PDG]} : m_{\kappa,\text{BW}} = 824 \pm 30 \text{ MeV}, \quad \Gamma_{\kappa,\text{BW}} = 478 \pm 50 \text{ MeV}. \quad (2)$$

The BW and the pole widths are compatible, but the BW mass is somewhat larger. There is, however, no friction, since BW and pole masses are different

---

\* Presented at the XIII Workshop on Particle Correlations and Femtoscopy, Kraków, Poland, May 22–26, 2018.

quantities which coincide only when a resonance is narrow. This is definitely not the case for the  $\kappa$ , which is a very broad state with a width-to-mass ratio larger than 0.5.

In a certain sense, the light  $\kappa$  can be regarded as the “brother” of the light  $\sigma \equiv f_0(500)$  meson [1]. This state is also very broad and for a long time, it was not clear if there is a pole on the complex plane. Now, its existence is confirmed by many studies and the state is listed in the PDG, see also the review paper [4]. The destiny of the light  $\kappa$  looks somewhat similar: its final confirmation is probably just a matter of time.

Yet, a different issue is the nature of the  $\kappa \equiv K_0^*(700)$  and the  $\sigma \equiv f_0(500)$ . According to mounting evidence, both states are not simple quark–antiquark states, but are rather four-quark objects, either in the form of a tetraquark nonet together with  $a_0(980)$  and  $f_0(980)$  [5] or as dynamically generated molecular-like states [6]. The  $\kappa$  can be then interpreted as a diquark–antidiquark state ( $[u, d][\bar{d}, \bar{s}], \dots$ ) and/or as  $K\pi$  state (mixing among these configurations is of course possible and rather probable to occur). If  $\kappa$  is not  $\bar{q}q$ , where should be the scalar strange quarkonium? According to the quark model [7] and modern chiral approaches [8], the lightest  $\bar{q}q$  kaonic state ( $u\bar{s}, \dots$ ) is the well-established  $K_0^*(1430)$  (similarly, the lightest scalar/isoscalar quarkonium is the state  $f_0(1370)$ ). The question that we review in this work is the link between the standard state  $K_0^*(1430)$  and the dynamically generated state  $K_0^*(700)$ . We find (see Sec. 2) that the  $\pi K$  loops dressing  $K_0^*(1430)$  generate  $K_0^*(700)$  as a companion pole (a peculiar four-quark object) [9] (similarly, the  $a_0(980)$  emerges as a companion pole of  $a_0(1450)$  [10]).

There is, however, a related important question: if the light  $\kappa$  is existent, should it be included into thermal hadronic models [11]? At a first sight, the answer is ‘yes’. In fact, the light  $\kappa$  is the second-lightest state with nonzero strangeness, thus potentially relevant. Yet, a detailed analysis of the problem [12] shows that one should better *not* include this state into a thermal model (see Sec. 3). Namely, also repulsive channels contribute to the thermodynamics [13–15]. Just as for the  $f_0(500)$  whose contribution is cancelled by  $\pi\pi$  scattering with  $I = 2$ , the contribution of the  $\kappa$  is cancelled by the repulsion in  $\pi K$  channel with  $I = 3/2$ . Thus, the easiest thing to do is to neglect both the  $f_0(500)$  and the  $K_0^*(700)$  when studying hadronic thermal models for the late stage of heavy-ion collisions.

## 2. The light $\kappa$ in the vacuum

As a first step, we write down a Lagrangian that contains *only one* scalar state  $K_0^*$ , to be identified with  $K_0^*(1430)$ , coupled to  $K\pi$  pairs

$$\mathcal{L}_{K_0^*} = a K_0^{*+} K^- \pi^0 + b K_0^{*+} \partial_\mu K^- \partial^\mu \pi^0 + \dots , \quad (3)$$

where dots refer to other isospin channels. Note, there is *no*  $\kappa \equiv K_0^*(700)$  in the model (yet). There are both derivative and non-derivative terms: the former naturally dominates in the context of chiral perturbation theory and also emerge from the extended Linear Sigma Model [8]. The decay width reads

$$\Gamma_{K_0^* \rightarrow K\pi}(m) = 3 \frac{|\vec{k}_1|}{8\pi m^2} \left[ a - b \frac{m^2 - M_K^2 - M_\pi^2}{2} \right]^2 F_\Lambda(m), \quad (4)$$

with the vertex function  $F_\Lambda(m) = \exp(-2\vec{k}_1^2/\Lambda^2)$ . Here,  $\Lambda$  is an energy scale describing the nonlocal nature of mesons [16],  $\vec{k}_1$  the three-momentum of one outgoing particle,  $M_K$  the kaon mass, and  $M_\pi$  the pion mass. (For details and phenomenology of the spectral function, see Ref. [17].)

The propagator of  $K_0^*$  is given by  $\Delta_{K_0^*}(m^2) = [m^2 - M_0^2 + \Pi(m^2) + i\varepsilon]^{-1}$ ,  $M_0$  being the bare mass of  $K_0^*(1430)$  and  $\Pi(m^2)$  the one-loop contribution. The spectral function  $d_{K_0^*}(m) = \frac{2m}{\pi} |\text{Im} \Delta_{K_0^*}(p^2 = m^2)|$  is the mass probability density (its integral is normalized to unity). Typically, for the Breit–Wigner value  $M_{\text{BW}}$  determined as  $M_{\text{BW}}^2 - M_0^2 + \text{Re} \Pi(M_{\text{BW}}^2) = 0$ , the spectral function has a peak's width  $\Gamma_{\text{BW}} = \text{Im} \Pi(M_{\text{BW}})/M_{\text{BW}}$ . A useful approximation, valid if the width is sufficiently small, is the relativistic Breit–Wigner expression

$$d_{K_0^*}(m) \approx d_{K_0^*}^{\text{BW}}(m) = N \left[ (m^2 - M_{\text{BW}}^2)^2 + M_{\text{BW}}^2 \Gamma_{\text{BW}}^2 \right]^{-1}. \quad (5)$$

Under this approximation, there is only one pole in the complex plane at  $m^2 \simeq M_{\text{BW}}^2 - iM_{\text{BW}}\Gamma_{\text{BW}}$  (hence,  $m \simeq M_{\text{BW}} - i\Gamma_{\text{BW}}/2$ ). However, when a resonance is broad, these approximations are not valid anymore.

We now turn to  $\pi K$  scattering. Within our framework, the pion–kaon phase shift is given by [9]

$$\delta_{\pi K, \text{S-wave}}(m) = \delta_{(I=1/2, J=0)}(m) = \frac{1}{2} \arccos [1 - \pi \Gamma_{K_0^*}(m) d_{K_0^*}(m)], \quad (6)$$

where  $\delta_{(I,J)}(m)$  is the general phase shift for a given isospin  $I$  and total spin  $J$ . The amplitude of the process and the phase-shift are linked by  $a_{(I,J)} = (e^{i\delta_{(I,J)}(m)} - 1)/(2i)$ . The parameters  $(a, b, M_0, \Lambda)$  entering in Eq. (3) were determined via a fit to  $\pi K$  phase-shift data [18], see Ref. [9] for details. A very good description of data is achieved. A study of the complex plane shows an interesting fact: besides the pole corresponding to the well-known  $K_0^*(1430)$  state  $(1.413 \pm 0.057) - i(0.127 \pm 0.011)$  GeV, there is a second pole which correspond to  $K_0^*(700)$

$$(0.745 \pm 0.029) - i(0.263 \pm 0.027) \text{ GeV}. \quad (7)$$

The numerical value is compatible with the PDG value of Eq. (1). A large- $N_c$  study confirms that, while the first pole tends to the real axis (and hence is a  $\bar{q}q$  state), the second one moves away from it, as it is expected for a dynamically generated state.

In conclusion, the simple model of Eq. (3) is able to describe  $\pi K$  scattering data and naturally gives rise to the pole of  $K_0^*(700)$  as a companion pole of the predominantly quark–antiquark resonance  $K_0^*(1430)$ .

### 3. The light $\kappa$ at nonzero temperature

The partition function of an hadronic gas can be expressed as the sum of the contributions of stable particles and their mutual interactions

$$\ln Z = \ln Z_{\text{pions}} + \ln Z_{\text{kaons}} + \cdots + \ln Z^{\text{int}}, \quad \ln Z^{\text{int}} = \sum_{I,J} \ln Z_{(I,J)}. \quad (8)$$

The first term  $\ln Z_{\text{pions}} = 3F_1(m_\pi)$  refers to pions and  $\ln Z_{\text{kaons}} = 4F_1(m)$  to kaons, where  $F_1(m) = \int \frac{d^3 p}{(2\pi)^3} \ln \left[ 1 - e^{-\sqrt{\vec{p}^2 + m_\pi^2}/T} \right]$  is the contribution of a free particle with mass  $m$ . The term  $\ln Z_{IJ}$  refers to the contribution of the interactions in the  $(I, J)$  channel [13]

$$\ln Z_{(I,J)} = (2I+1)(2J+1) \int_0^\infty \frac{dm}{\pi} \frac{d\delta_{(I,J)}(m)}{dm} F_1(m). \quad (9)$$

When in a certain channel a narrow resonance is present, one finds its standard contribution. For instance, for  $I = J = 1$ , the  $\rho$  meson is produced. In the nonrelativistic BW-limit,  $\frac{1}{\pi} \frac{d\delta_{(1,1)}(m)}{dm} \simeq \frac{\Gamma_\rho}{2\pi} [(m - M_\rho)^2 + \Gamma_\rho^2/4]^{-1}$ . (Moreover, for  $\Gamma_\rho \rightarrow 0$ ,  $\delta(m - M_\rho)$  emerges: the contribution of a stable  $\rho$  is obtained.)

However, Eq. (9) is very general and can describe also broad resonances as well as non-resonant channels, such as repulsive ones. This is important for the  $\kappa$ . In the resonant  $I = 1/2, J = 0$  channel in which the  $\kappa$  is formed, one has (upon integrating up to 1 GeV)  $\ln Z_{(1/2,0)} = \int_0^1 \frac{2dm}{\pi} \frac{d\delta_{(1/2,0)}(m)}{dm} F_1(m)$ . This is sizable. However, one should also consider the repulsion in the  $I = 3/2, J = 0$  channel. Remarkably, the sum

$$\ln Z_{(1/2,0)} + \ln Z_{(3/2,0)} = \int_0^{1 \text{ GeV}} dm \left( \frac{2}{\pi} \frac{d\delta_{(1/2,0)}(m)}{dm} + \frac{4}{\pi} \frac{d\delta_{(3/2,0)}(m)}{dm} \right) F_1(m) \quad (10)$$

is small. Namely, while  $\frac{d\delta_{(1/2,0)}(m)}{dm} > 0$  (attraction),  $\frac{4}{\pi} \frac{d\delta_{(3/2,0)}(m)}{dm} < 0$  (repulsion). Note:  $\frac{1}{\pi} \frac{d\delta_{(1/2,0)}(m)}{dm} \neq d_{K_0^*}(m)$ . (This would be true only in the BW-limit). In conclusion, the light  $\kappa$  can be safely neglected in the construction of thermal hadronic models.

#### 4. Conclusions

We have described the emergence of the state  $\kappa \equiv K_0^*(700)$  as a companion pole of  $K_0^*(1430)$  by using an effective hadronic model [9]. The numerical value of the pole (7) is in agreement with the present PDG estimate of Eq. (1). On the other hand, contrary to the naive expectations, the light  $\kappa$  is not relevant in a thermal hadronic gas. Namely, its influence on thermodynamical properties is cancelled by a repulsion in the  $I = 3/2$  channel. Either one includes both the light  $\kappa$  and the repulsion, or — even easier — neglects both of them.

The author thanks M. Piotrowska, T. Wolkanowski, W. Broniowski, V. Begun for cooperations. Financial support from the National Science Centre, Poland (NCN) through the OPUS project No. 2015/17/B/ST2/01625 is acknowledged.

#### REFERENCES

- [1] M. Tanabashi *et al.* [Particle Data Group], *Phys. Rev. D* **98**, 030001 (2018).
- [2] C. Patrignani *et al.* [Particle Data Group], *Chin. Phys. C* **40**, 100001 (2016).
- [3] S. Ishida *et al.*, *Prog. Theor. Phys.* **98**, 621 (1997) [[arXiv:hep-ph/9705437](#)]; D. Black, A.H. Fariborz, F. Sannino, J. Schechter, *Phys. Rev. D* **58**, 054012 (1998); P.C. Magalhaes *et al.*, *Phys. Rev. D* **84**, 094001 (2011) [[arXiv:1105.5120 \[hep-ph\]](#)]; S. Descotes-Genon, B. Moussallam, *Eur. Phys. J. C* **48**, 553 (2006) [[arXiv:hep-ph/0607133](#)]; J.R. Peláez, *Phys. Rev. Lett.* **92**, 102001 (2004) [[arXiv:hep-ph/0309292](#)]; J.R. Peláez, A. Rodas, *Eur. Phys. J. C* **77**, 431 (2017) [[arXiv:1703.07661 \[hep-ph\]](#)]; P. Buettiker, S. Descotes-Genon, B. Moussallam, *Eur. Phys. J. C* **33**, 409 (2004) [[arXiv:hep-ph/0310283](#)]; J. Sa Borges, J. Soares Barbosa, V. Oguri, *Phys. Lett. B* **412**, 389 (1997); H.Q. Zheng *et al.*, *Nucl. Phys. A* **733**, 235 (2004) [[arXiv:hep-ph/0310293](#)]; Z.Y. Zhou, H.Q. Zheng, *Nucl. Phys. A* **775**, 212 (2006) [[arXiv:hep-ph/0603062](#)]; A.H. Fariborz, E. Pourjafarabadi, S. Zarepour, S.M. Zerbarjad, *Phys. Rev. D* **92**, 113002 (2015) [[arXiv:1511.01623 \[hep-ph\]](#)]; S. Descotes-Genon, B. Moussallam, *Eur. Phys. J. C* **48**, 553 (2006) [[arXiv:hep-ph/0607133](#)]; M. Ablikim *et al.* [BES Collaboration], *Phys. Lett. B* **698**, 183 (2011) [[arXiv:1008.4489 \[hep-ex\]](#)].
- [4] J.R. Peláez, *Phys. Rep.* **658**, 1 (2016) [[arXiv:1510.00653 \[hep-ph\]](#)].
- [5] R.L. Jaffe, *Phys. Rev. D* **15**, 267 (1977); **15**, 281 (1977); *Phys. Rep.* **409**, 1 (2005) [*Nucl. Phys. Proc. Suppl.* **142**, 343 (2005)] [[arXiv:hep-ph/0409065](#)]; L. Maiani, F. Piccinini, A.D. Polosa, V. Riquer, *Phys. Rev. Lett.* **93**, 212002

- (2004) [[arXiv:hep-ph/0407017](#)]; F. Giacosa, *Phys. Rev. D* **74**, 014028 (2006) [[arXiv:hep-ph/0605191](#)]; **75**, 054007 (2007) [[arXiv:hep-ph/0611388](#)]; A.H. Fariborz, R. Jora, J. Schechter, *Phys. Rev. D* **72**, 034001 (2005) [[arXiv:hep-ph/0506170](#)]; M. Napsuciale, S. Rodríguez, *Phys. Rev. D* **70**, 094043 (2004) [[arXiv:hep-ph/0407037](#)].
- [6] E. van Beveren *et al.*, *Z. Phys. C* **30**, 615 (1986) [[arXiv:0710.4067 \[hep-ph\]](#)]; E. van Beveren, D.V. Bugg, F. Kleefeld, G. Rupp, *Phys. Lett. B* **641**, 265 (2006) [[arXiv:hep-ph/0606022](#)]; J.R. Peláez, *Phys. Rev. Lett.* **92**, 102001 (2004) [[arXiv:hep-ph/0309292](#)]; J.A. Oller, E. Oset, *Nucl. Phys. A* **620**, 438 (1997) [[arXiv:hep-ph/9702314](#)]; J.A. Oller, E. Oset, J.R. Peláez, *Phys. Rev. D* **59**, 074001 (1999) [[arXiv:hep-ph/9804209](#)]; *Phys. Rev. Lett.* **80**, 3452 (1998).
- [7] S. Godfrey, N. Isgur, *Phys. Rev. D* **32**, 189 (1985).
- [8] D. Paganlaja *et al.*, *Phys. Rev. D* **87**, 014011 (2012); S. Janowski, F. Giacosa, D.H. Rischke, *Phys. Rev. D* **90**, 114005 (2014) [[arXiv:1408.4921 \[hep-ph\]](#)].
- [9] T. Wolkanowski, M. Sołtysiak, F. Giacosa, *Nucl. Phys. B* **909**, 418 (2016) [[arXiv:1512.01071 \[hep-ph\]](#)].
- [10] T. Wolkanowski, F. Giacosa, D.H. Rischke, *Phys. Rev. D* **93**, 014002 (2016) [[arXiv:1508.00372 \[hep-ph\]](#)]; M. Boglione, M.R. Pennington, *Phys. Rev. D* **65**, 114010 (2002) [[arXiv:hep-ph/0203149](#)].
- [11] A. Andronic, P. Braun-Munzinger, J. Stachel, *Nucl. Phys. A* **772**, 167 (2006) [[arXiv:nucl-th/0511071](#)]; *Phys. Lett. B* **673**, 142 (2009) [*Erratum ibid. B* **678**, 516 (2009)] [[arXiv:0812.1186 \[nucl-th\]](#)]; P. Alba *et al.*, *Phys. Lett. B* **738**, 305 (2014) [[arXiv:1403.4903 \[hep-ph\]](#)]; G. Torrieriet *et al.*, *Comput. Phys. Commun.* **167**, 229 (2005) [[arXiv:nucl-th/0404083](#)].
- [12] W. Broniowski, F. Giacosa, V. Begun, *Phys. Rev. C* **92**, 034905 (2015) [[arXiv:1506.01260 \[nucl-th\]](#)].
- [13] R. Dashen, S.K. Ma, H.J. Bernstein, *Phys. Rev.* **187**, 345 (1969); R.F. Dashen, R. Rajaraman, *Phys. Rev. D* **10**, 694 (1974); W. Weinhold, B.L. Friman, W. Nörenberg, *Acta Phys. Pol. B* **27**, 3249 (1996); *Phys. Lett. B* **433**, 236 (1998) [[arXiv:nucl-th/9710014](#)].
- [14] W. Broniowski, W. Florkowski, B. Hiller, *Phys. Rev. C* **68**, 034911 (2003) [[arXiv:nucl-th/0306034](#)].
- [15] P.M. Lo, *Eur. Phys. J. C* **77**, 533 (2017) [[arXiv:1707.04490 \[hep-ph\]](#)]; P.M. Lo *et al.*, *Phys. Rev. C* **96**, 015207 (2017) [[arXiv:1703.00306 \[nucl-th\]](#)].
- [16] J. Terning, *Phys. Rev. D* **44**, 887 (1991); A. Faessler *et al.*, *Phys. Rev. D* **68**, 014011 (2003) [[arXiv:hep-ph/0304031](#)]; F. Giacosa, T. Gutsche, A. Faessler, *Phys. Rev. C* **71**, 025202 (2005) [[arXiv:hep-ph/0408085](#)].
- [17] F. Giacosa, G. Pagliara, *Phys. Rev. C* **76**, 065204 (2007) [[arXiv:0707.3594 \[hep-ph\]](#)]; S. Coito, F. Giacosa, *Nucl. Phys. A* **981**, 38 (2019) [[arXiv:1712.00969 \[hep-ph\]](#)].
- [18] D. Aston *et al.*, *Nucl. Phys. B* **296**, 493 (1988).