FEMTOSCOPY OF PROTON, LIGHT NUCLEI, AND STRANGE HADRONS IN Au+Au COLLISIONS AT STAR*

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In these proceedings, we present the measurements of proton, light nuclei, and strange particle with neutral kaons correlation functions in $\mathrm{Au}+\mathrm{Au}$ collisions at the BES program and top RHIC energy. The experimental results are compared with model calculations to extract the size of emitting source and the properties of the final-state interactions. The collision energy and centrality dependence of the source size are studied. Further, the implications for the production mechanism of light nuclei are discussed.

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1. Introduction

Measurements of the correlation function for a pair of particles with small relative momenta can provide insight into the geometry and lifetime of the particle-emitting source in relativistic heavy-ion collisions [1]. By measuring correlation at low relative velocities, one could access the smallest sizes in nature, which correspond to the size of a nucleon (1 fm), therefore such a two-particle correlation method is called Femtoscopy [2].

The fundamental observable in femtoscopy is the correlation function $C(k^*)$. The theoretical definition is expressed as a function of the relative distance between two particles \mathbf{r}^* and their reduced relative momentum, $k^* = \frac{1}{2} |\mathbf{p}_1^* - \mathbf{p}_2^*|$ in the pair rest frame, with $\mathbf{p}_1^* = -\mathbf{p}_2^*$, by the Koonin–Pratt formula [3]

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$$C(k^*) = \int S(\mathbf{r}^*) |\psi(\mathbf{r}^*, \mathbf{k}^*)|^2 d^3r, \qquad (1)$$

where the first term $S(\mathbf{r}^*)$ describes the source that emits the particles; the second term contains the interaction part via the two-particle wave function $\psi(\mathbf{r}^*, \mathbf{k}^*)$.

The experimental correlation function is defined as a ratio of the probability of registering two particles simultaneously (in the same event) to the product of registering probabilities of such particles independently (in the mixed event):

$$C\left(k^{*}\right) = \frac{N_{\text{same}}\left(k^{*}\right)}{N_{\text{mixed}}\left(k^{*}\right)}.$$
(2)

The shape and amplitude of the correlation function are determined by the quantum statistical effect (identical pair) and final-state interaction (strong and Coulomb interactions).

2. Particle identification and reconstruction

Particle identification (PID) of protons and deuterons is done by using the energy loss (dE/dx) information measured by the Time Projection Chamber (TPC) (Fig. 1(a)) and particles' m^2 information from the Time Of Flight (TOF) (Fig. 1(b)). A combination of dE/dx and m^2 criteria is used to identify the particles with purity higher than 96%.

The K_s^0 particles are reconstructed via a weak decay channel: $K_s^0 \to \pi^+ + \pi^-$ (Fig. 1(c)). A set of topological cuts was used to ensure the reconstructed K_s^0 purity higher than 95%.

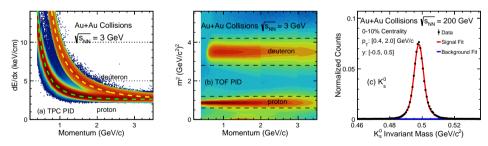


Fig. 1. (Color online) (a) The $\mathrm{d}E/\mathrm{d}x$ distribution of protons and deuterons versus total momentum in Au+Au collisions at $\sqrt{s_{NN}}=3\,\mathrm{GeV}$. The dashed curves are the corresponding Bichsel expectations. (b) Particle mass square (m^2) distribution versus total momentum in Au+Au collisions at $\sqrt{s_{NN}}=3\,\mathrm{GeV}$. The dashed lines represent the m^2 ranges used in the analysis. (c) The K_s^0 invariant mass distribution in 0–10% centrality at $\sqrt{s_{NN}}=200\,\mathrm{GeV}$. The red/light gray line represents the fit to the data including signal and background, and the blue/dark gray line to the background alone.

3. Results

3.1.
$$K_s^0$$
- K_s^0 correlation functions at $\sqrt{s_{NN}}=39$ GeV and 200 GeV

Figure 2 shows the K_s^0 – K_s^0 correlation functions as a function of $q_{\rm inv}$ ($q_{\rm inv}=2k^*$) at $\sqrt{s_{NN}}=39\,{\rm GeV}$ and 200 GeV with three centrality classes (0–10%, 10–70%, and 0–70%) in Au+Au collisions. The correlation functions show dip structures around $q_{\rm inv}=0.1\,{\rm GeV}/c$, which is caused by the near-threshold $f_0(980)$ and $a_0(980)$ resonances. The parameterization is done by using the Gaussian function which only includes quantum statistics (QS)

$$C\left(q_{\text{inv}}\right) = 1 + \lambda \ e^{\left[-R^2 q_{\text{inv}}^2\right]},\tag{3}$$

and the Lednicky–Lyuboshitz model [2] (L–L model) which includes both QS and strong interaction (SI)

$$C(q) = 1 + \lambda \times \left(e^{\left[-R^2 q^2 \right]} + \frac{1}{2} \left[\left| \frac{f(k^*)}{R} \right|^2 + \frac{4\Re f(k^*)}{\sqrt{\pi}R} F1(qR) - \frac{2\mathfrak{f}f(k^*)}{\sqrt{\pi}R} F2(qR) \right] \right), (4)$$

where R is the invariant radius, λ is correlation strength, $f(k^*)$ is the s-wave $K^0\bar{K}^0$ scattering amplitude. In this work, we use the resonance masses and couplings from Refs. [4–7]. The experimental data can be well described by

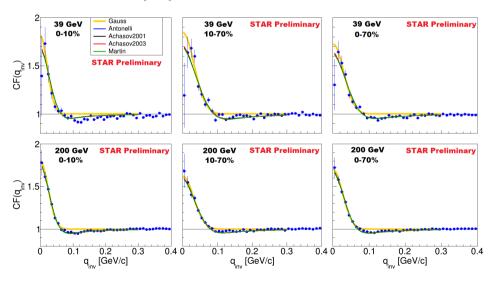


Fig. 2. The K_s^0 - K_s^0 correlation functions at $\sqrt{s_{NN}}=39\,\mathrm{GeV}$ (upper row) and 200 GeV (lower row) with three centrality classes (0–10%, 10–70%, and 0–70%) as a function of q_{inv} . The statistical errors are shown as vertical lines. The colored lines represent the Gaussian fit and L–L model fit results.

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the L–L model, while the Gaussian function failed to describe the dip structure. This implies that the strong final-state interaction has a significant effect between $K_s^0 - K_s^0$ pairs. Figure 3 shows the centrality dependence and energy dependence of the extracted radii. The results suggest a decreasing trend from central to peripheral collisions and an increasing trend from low to high energy. Besides, a significant difference in radii between data and QS and the L–L model is found.

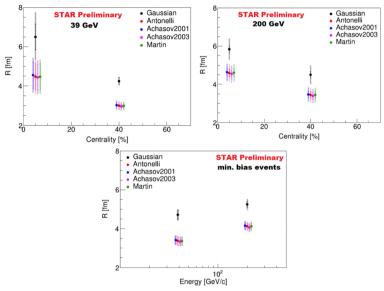


Fig. 3. Centrality and energy dependence of R extracted by fitting the Gaussian function and the L–L model to $K_s^0 - K_s^0$ correlation function at $\sqrt{s_{NN}} = 39 \, \text{GeV}$ and 200 GeV in Au+Au collisions. The statistical and systematic errors are shown as the vertical lines and shaded boxes, respectively.

3.2. Light-nuclei correlation functions at $\sqrt{s_{NN}} = 3 \text{ GeV}$

The production mechanisms of light nuclei are still under debate. Some experimental methods to resolve the problem have been proposed [8]. Femtoscopy of light nuclei may provide a unique tool for obtaining detailed information about the production mechanism. At low-beam energies ($\sqrt{s_{NN}}$ < 20 GeV), an enhancement in the light-nuclei yield is expected due to the high baryon density [9]. This provides a great opportunity for a precise femtoscopy that involves light nuclei [10].

Figure 4 shows the first measurements of p-d and d-d correlation functions at mid-rapidity with four centrality classes in Au+Au collisions at $\sqrt{s_{NN}} = 3 \,\text{GeV}$. Clear depletions below unity due to the final-state interaction at small k^* are observed in the measured correlations. Also no significant centrality dependence is found.

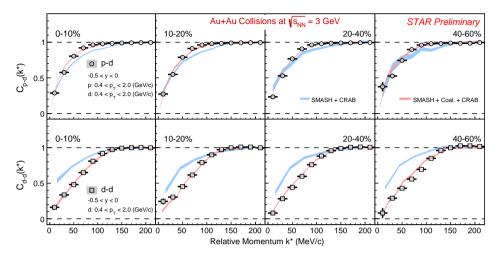


Fig. 4. (Color online) The p-d and d-d correlation functions in different collision centralities in Au+Au collisions at $\sqrt{s_{NN}} = 3 \,\text{GeV}$. Statistical and systematic errors are shown as vertical lines and grey boxes, respectively. The colored bands represent the p-d and d-d correlations obtained with the deuteron from nucleon coalescence (red/dark gray) in SMASH and directly produced from SMASH via hadronic scattering (blue/light gray), respectively.

To further understand light-nucleus formation, a transport model, Simulating Many Accelerated Strongly-interacting Hadrons Model (SMASH) [11], is used to simulate particle production in $\sqrt{s_{NN}}=3\,\mathrm{GeV}$ Au+Au collisions. We use the cascade mode in the SMASH model where the particles are propagated as in vacuum between collisions with other particles. In this work, two different versions of the SMASH model are chosen to generate deuterons. In the first version, SMASH can generate deuterons via hadron scattering, such as $p+n+\pi\leftrightarrow d+\pi$. These are so-called directly produced deuterons. The second version of the SMASH model produces nucleons and the deuterons are formed from nucleon coalescence with an afterburner package based on Wigner function [12]. The SMASH model itself does not contain femtoscopic correlation between two particles after kinetic freeze-out. The interaction (Coulomb potential, strong interaction, and quantum statistics) between two particles is introduced by the Correlation Afterburner (CRAB) package [13], and the input potentials of $p\!-\!d$ and $d\!-\!d$ pairs are taken from Ref. [14].

The resulting p-d and d-d correlation functions from the models are shown as color bands in Fig. 4, respectively. It is found that the experimental data are well reproduced by the SMASH plus coalescence calculations. On the other hand, the correlation functions with directly produced deuterons from SMASH can only qualitatively reproduce the overall trends,

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but over/underestimate the data depending on the particle pair. Those observations imply that at $\sqrt{s_{NN}} = 3 \,\text{GeV}$ Au+Au collisions, deuterons are likely formed via the nucleon coalescence processes.

4. Summary

In summary, we report the femtoscopy results of protons, light nuclei, and strange hadrons in Au+Au collisions from STAR. For K_s^0 – K_s^0 correlations, which are well described by the L–L model, it is found that the FSI plays an important role. The extracted radii of the particle emitting source show a clear centrality and energy dependence. As for the light-nuclei correlations, the SMASH model with coalescence can well reproduce data, which suggests that the coalescence mechanism may dominate the deuteron production at $\sqrt{s_{NN}} = 3 \, \text{GeV}$ Au+Au collisions.

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REFERENCES

- [1] M.A. Lisa et al., Annu. Rev. Nucl. Part. Sci. 55, 357 (2005).
- [2] R. Lednicky, Nucl. Phys. A 774, 189 (2006).
- [3] S. Pratt, Phys. Rev. D 33, 1314 (1986).
- [4] A.D. Martin, E.N. Ozmutlu, E.J. Squires, Nucl. Phys. B 121, 514 (1977).
- [5] V. Baru et al., Eur. Phys. J. A 23, 523 (2005).
- [6] A. Antonelli, eConf C020620, THAT06 (2002), arXiv:hep-ex/0209069.
- [7] N.N. Achasov, V.V. Gubin, *Phys. Rev. D* **63**, 094007 (2001).
- [8] St. Mrówczyński, Eur. Phys. J. Spec. Top. 229, 3559 (2020).
- [9] STAR Collaboration (J. Adam et al.), Phys. Rev. C 99, 064905 (2019).
- [10] St. Mrówczyński, P. Słoń, Phys. Rev. C 104, 024909 (2021).
- [11] J. Weil et al., Phys. Rev. C 94, 054905 (2016).
- [12] W. Zhao et al., Phys. Rev. C 98, 054905 (2018).
- [13] https://web.pa.msu.edu/people/pratts/freecodes/crab/home.html
- [14] B.K. Jennings et al., Phys. Rev. C 33, 1303 (1986).