

## MULTI-HIGGS PRODUCTION AND UNITARITY\*

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Multi-particle production has been under scrutiny in spontaneously broken scalar theories. In this article, the self-consistent Schwinger–Dyson equation is solved in the spectral representation, and the multi-scalar production rate is calculated. We find an amplitude growing quadratically with the energy, which leads to an asymptotically decreasing scalar propagator.

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## 1. Introduction

There are problems with perturbative results in self-interacting scalar theories at high multiplicities and high energies [1]. Due to the factorially growing number of diagrams, the  $h^* \rightarrow n \times h$  amplitude grows without limit and may violate unitarity. Recently, Khoze *et al.* [2] proposed the ‘Higgs-persion’ mechanism, where the highly excited Higgs appears only as an internal, virtual particle, and the off-shell propagator tames the fast-growing decay amplitude. They claim that if the propagator can be resummed, then it can suppress the growing high-energy contribution of the decay amplitude. However, the results of Ref. [2] were criticized in Ref. [3], where it was argued that the propagator diverges and cannot be resummed as proposed due to the exponential growth of the amplitude and self-energy.

In this paper, we calculate the multi-scalar transition rate by solving the Schwinger–Dyson equation (SDE) in the dispersion representation of the propagator and self-energy.

## 2. Multi-Higgs production

The physical self-interacting single real scalar field  $\varphi(x)$  is the perturbation around the vacuum expectation value  $v$  after symmetry breaking. The Lagrangian is

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$$\mathcal{L} = \frac{1}{2} \partial^\mu \varphi(x) \partial_\mu \varphi(x) - \frac{m^2}{2} \varphi^2(x) - \frac{\kappa}{3} \varphi^3(x) - \frac{\lambda}{4} \varphi^4(x), \quad (1)$$

where  $m = \sqrt{2\lambda v}$  and  $\kappa = 3\lambda v$ .

The  $R_n(p^2)$  transition rate characterizes the decay of a highly virtual Higgs boson ( $|p|^2 \gg m^2$ ) into  $n$  Higgses,  $1^* \rightarrow n$  defined as

$$\int d\Pi_n |\mathcal{M}(1 \rightarrow n)|^2 = R_n(p^2), \quad (2)$$

where the  $|\mathcal{M}|^2$  scattering amplitude squared is integrated over the  $n$ -particle phase space  $d\Pi_n$ . Recently, in [1, 2], the transition rate was calculated close to the threshold using the steepest descent method, valid only in the  $\lambda \rightarrow 0$  limit,  $\lambda n = \text{fixed} \gg 1$ , and  $\epsilon = (E - nm)/nm \ll 1$ . The peak rate grows exponentially with the energy  $E$  of the initial Higgs boson, with fixed  $\lambda$ . However, none of the assumed conditions  $\lambda \rightarrow 0$ ,  $\epsilon = \text{fixed} \ll 1$  are satisfied for  $\lambda$  and  $E$ .

### 3. Solving the Schwinger–Dyson equation

For the spontaneously broken  $\varphi^4$  theory given in (1), we can derive the following self-consistent equation for the self-energy  $\Pi(p^2)$ :

$$\begin{aligned} \Pi(p^2) &= \kappa \int_{k_1} G_2^c(-k_1) G_2^c(k_1 + p) \Gamma_3(k_1, p) \\ &+ \lambda \int_{k_1} \int_{k_2} G_2^c(k_1) G_2^c(-k_1 - k_2) G_2^c(k_2 + p) \Gamma_4(-k_1, k_1 + k_2, p) \\ &+ 3\lambda \int_{k_1} \int_{k_2} G_2^c(-k_1) G_2^c(-k_2) G_2^c(k_1 + k_2 + p) \Gamma_3(k_1 + p, k_2) G_2^c(k_1 + p) \Gamma_3(k_1, p), \end{aligned} \quad (3)$$

where  $G_2^c(k)$  is the connected two-point function,  $\Gamma_n(k_1, k_2, \dots, k_{n-1})$  is the  $n$ -point vertex function [5],  $\int_k \equiv \int \frac{d^4 k}{(2\pi)^4}$ . The generic spectral decomposition of the scalar two-point function reads

$$G(p^2) = \frac{iZ}{p^2 - m^2 + i\epsilon} + \int_{\omega_{\text{th}}}^{\infty} d\omega \frac{i\sigma(\omega)}{p^2 - \omega + i\epsilon}, \quad (4)$$

where  $Z$  is the wave function renormalization factor,  $\omega_{\text{th}} = 4m^2$  corresponds to the two particle threshold. In 3+1 dimensions, the bubble contribution

has a logarithmic divergence, while the setting sun has quadratic divergence. The dispersion relation of the renormalized self-energy is defined in the non-minimal momentum subtraction scheme

$$\Pi_R(p^2) = \int_{\omega_{\text{th}}}^{\infty} d\omega \frac{\rho(\omega)}{p^2 - \omega + i\epsilon} \left( \frac{p^2 - m^2}{\omega - m^2} \right)^2 = \int_{\omega_{\text{th}}}^{\infty} d\omega \frac{\tilde{\rho}(\omega, p^2; m^2)}{p^2 - \omega + i\epsilon}. \quad (5)$$

The spectral functions  $\sigma(\omega)$ ,  $\rho(\omega)$  are further related via the trivial identity  $G(p^2)G^{-1}(p^2) = 1$ , where  $G^{-1}(p^2) = p^2 - m^2 - \Pi_R(p^2)$  [5]

$$\sigma(p^2) = \frac{Z\rho(p^2)}{(p^2 - m^2)^2} + \frac{1}{p^2 - m^2} \text{P} \int_{\omega_{\text{th}}}^{\infty} d\omega \frac{\sigma(p^2) \tilde{\rho}(\omega, p^2; m^2) + \rho(p^2) \sigma(\omega)}{p^2 - \omega}, \quad (6)$$

where  $\text{P}$  denotes the Cauchy principal value integral and  $Z$  is the wave function renormalization factor.  $Z$  satisfies the sum rule  $Z + \int_{\omega_{\text{th}}}^{\infty} d\omega \sigma(\omega) = 1$ .

Using the LSZ reduction formula and the optical theorem, we can connect the transition rate to the renormalized self-energy [4]

$$\sum_n R_n(p^2) = 2\pi Z\rho(p^2). \quad (7)$$

There are two self-consistent equations for  $\sigma(\omega)$  and  $\rho(\omega)$ , Eq. (6) and Eq. (3), after applying the spectral representations of Eqs. (4) and (5) [4]. They can be solved by iteration [6], starting from the perturbative propagator with vanishing  $\sigma(\omega)$ . Dimensional analysis determines that at asymptotic energies in Eq. (3), the contribution of the second line, the setting-sun diagram, is the dominant. For the small coupling,  $\lambda = 0.125$ , the spectral function of the propagator sharply peaks after the threshold, then asymptotically goes to zero at high energies, see the left panel of Fig. 1.

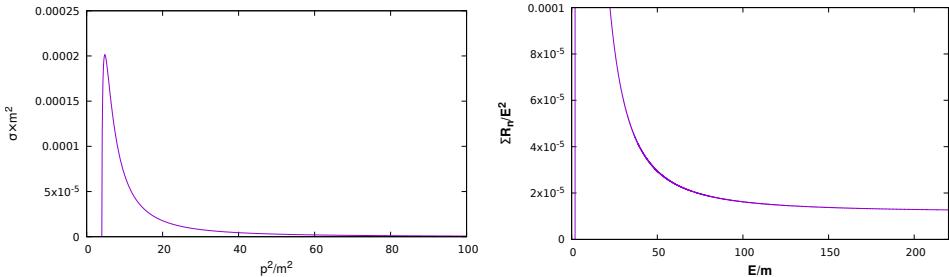


Fig. 1. The  $\sigma$  spectral function of the propagator (left); The transition rate  $\sum_n R_n(E)$  divided by  $E^2$  at high energies (right),  $\lambda = 0.125$ .

The numerical calculations show that the spectral function  $\rho(\omega)$ , related to the propagator, grows with  $p^2$  at  $E \gg m$  energies. Thus, the summed transition rate  $\sum_n R_n(E)$  goes with the square of the energy at asymptotically high energies  $E \gg m$ , see the right panel of Fig. 1.

The *Higgs propagator* usually can be written as a geometric series of the (one-particle irreducible, 1PI) self-energy  $\Pi(p^2)$

$$G(p^2) = \frac{i}{p^2 - m^2} \sum_{n=0}^{\infty} \left( -i\Pi(p^2) \frac{i}{p^2 - m^2} \right)^n = \frac{i}{p^2 - m^2 - \Pi(p^2)}. \quad (8)$$

However, if the series does not converge, it cannot be summed up, and the propagator becomes divergent. In Ref. [3], Belyaev *et al.* have already shown that in the solution of [2], the power series (8) diverges at high energies,  $E \gg m$ , as the self-energy grows exponentially with  $\sqrt{p^2}$ . However, considering our solution of the SDE, the spectral function of the self-energy grows only with  $p^2$ , thus the series is convergent at any  $E$  and one can resum the  $\Pi(p^2)$  into the denominator of the propagator. We end up with an asymptotically vanishing spectral density of the Higgs propagator, as seen in  $\sigma(\omega)$ , taming the cross section of the multi-Higgs production in gluon fusion via an excited, off-shell Higgs [4].

#### 4. Conclusions

We have solved numerically the SDE of the  $\varphi^4$  Higgs model in spectral representation to study the  $1^* \rightarrow n$  multi-Higgs production. A fast converging solution of the SDE was presented. The resulting propagator is well behaved, summable, and asymptotically vanishes at high energies. Thus, the cross section of the gluon fusion process ( $gg \rightarrow h^* \rightarrow nh$ ) goes to zero in the limit of infinite energy [4], meaning that perturbative unitarity is not violated in agreement with the renormalizability of the Standard Model.

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