# CHARGE DEPENDENCE AND SCALING PROPERTIES OF DYNAMICAL $K/\pi$ , $p/\pi$ , AND K/p FLUCTUATIONS FROM THE STAR EXPERIMENT\*

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Dynamical fluctuations in global conserved quantities such as baryon number, strangeness, or charge may be enhanced near a QCD critical point. Charge dependent results from new measurements of dynamical  $K/\pi$ ,  $p/\pi$ , and K/p ratio fluctuations are presented. The STAR experiment has performed a comprehensive study of the energy dependence of these dynamical fluctuations in Au+Au collisions at the energies  $\sqrt{s_{NN}} = 7.7-200$  GeV using the observable,  $\nu_{dyn}$ . These results are compared to previous measurements and to theoretical predictions. Various proposed scaling scenarios that attempt to remove the intrinsic volume dependence of  $\nu_{dyn}$  and to simplify comparisons between experimental measurements are also considered. Constructing an intensive quantity allows for a direct connection to thermodynamic predictions.

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## 1. Introduction

Fluctuations and correlations are well known signatures of phase transitions. In particular, the quark/gluon to hadronic phase transition may lead to significant fluctuations [1]. In 2010, the Relativistic Heavy Ion Collider (RHIC) began a program to search for the QCD critical point. This involves an "energy scan" of Au+Au collisions from top collision energy  $(\sqrt{s_{NN}} = 200 \text{ GeV})$  down to energies as low as  $\sqrt{s_{NN}} = 7.7 \text{ GeV}$  [2]. This

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critical point search will make use of the study of correlations and fluctuations, particularly those that could be enhanced during a phase transition that passes close to a critical point.

 $\nu_{\rm dyn}$  quantifies deviations in the particle ratios from those expected for an ideal statistical Poissonian distribution [3,4]. The definition of  $\nu_{\rm dyn, K/\pi}$ (describing fluctuations in the  $K/\pi$  ratio) is,

$$\nu_{\mathrm{dyn},K/\pi} = \frac{\langle N_K(N_K-1)\rangle}{\langle N_K\rangle^2} + \frac{\langle N_\pi(N_\pi-1)\rangle}{\langle N_\pi\rangle^2} - 2\frac{\langle N_KN_\pi\rangle}{\langle N_K\rangle\langle N_\pi\rangle}.$$
 (1)

A formula similar to (1) can be constructed for  $p/\pi$  and K/p ratio fluctuations. Additional information about  $\nu_{\rm dyn}$  can be found in [4, 5, 6]. An in-depth study of  $K/\pi$  fluctuations in Au+Au collisions at  $\sqrt{s_{NN}} = 200$  and 62.4 GeV was previously carried out by the STAR experiment [5].

## 2. Results and discussion

Current results on the energy dependence of measured dynamical particle ratio fluctuations from the STAR experiment are shown in figures 1, 2 (a), and 2 (b). These include previously presented results from the inclusive charged  $K/\pi$ ,  $p/\pi$ , and K/p ratios [9], as well as initial results on the charge dependence of these dynamical ratio fluctuations as a function of collision energy in the search for the QCD critical point at RHIC. Results from previous measurements of  $K/\pi$  fluctuations at  $\sqrt{s_{NN}} = 62.4$  and 200 GeV for both charge dependent and independent cases are discussed in [5].

Figure 1 shows  $\nu_{\mathrm{dyn},K/\pi}$  as a function of energy from the STAR experiment, measured in central 0-5% Au+Au collisions at the energies  $\sqrt{s_{NN}} = 7.7, 11.5, 19.6, 39, \text{ and } 200 \text{ GeV}$ . The total dynamical  $K/\pi$  fluctuations (filled stars) are constructed as  $K/\pi = (K^+ + K^-)/(\pi^+ + \pi^-)$ . Dynamical fluctuations can also be constructed for different charge combinations. The combinations include: ++, --, +-, and -+. The average of the same sign (open stars) and opposite sign (open crosses)  $K/\pi$  fluctuations are also shown in figure 1. The charge dependent fluctuations provide a more detailed picture of how particle production mechanisms affect the dynamical fluctuations. Also shown are published measurements of  $\sigma_{\text{dyn},K/\pi}$ from central 0-3.5% Pb+Pb collisions from the NA49 experiment [10], converted to  $\nu_{\text{dyn},K/\pi}$  using  $\sigma_{\text{dyn}}^2 \approx \nu_{\text{dyn}}$ . This was also done for  $p/\pi$  and K/p $\sigma_{\rm dyn}$  results from NA49. STAR has also directly calculated  $\sigma_{{\rm dyn},K/\pi}$  and experimentally verified the relationship between  $\nu_{dyn}$  and  $\sigma_{dyn}$ . The total charged dynamical  $K/\pi$  fluctuations measured by both STAR and NA49 are positive. STAR observes no large change in dynamical  $K/\pi$  fluctuations between the energies from  $\sqrt{s_{NN}} = 7.7-200 \,\text{GeV}.$ 



Fig. 1. Results for the measurement of inclusive charged (filled stars) and average same-sign (open stars) and opposite-sign (open crosses), charge dependent  $\nu_{dyn,K/\pi}$ .

The charge dependent dynamical  $K/\pi$  fluctuations measured by STAR are negative at all energies for both the same and opposite sign charge combinations. Both charged sign combinations are close in magnitude at  $\sqrt{s_{NN}} = 200 \,\text{GeV}$ , but the average opposite sign dynamical fluctuations become more negative with decreasing energy, faster, than the same sign fluctuations. Fluctuations in opposite-sign charge combinations tend to be more negative overall than their same-sign counterparts due to stronger cross-correlation (immediate local charge conservation) of two opposite sign particles produced by a single resonance decay. Additionally, the charge independent fluctuations are not a simple sum of the charge dependent components. This indicates that while the charge dependent results are always negative, the charge independent results can be positive.

Figure 2 (a) shows  $\nu_{\text{dyn},p/\pi}$  as a function of energy from the STAR experiment, measured in central 0–5% Au+Au collisions at the energies  $\sqrt{s_{NN}} =$ 7.7, 11.5, 19.6, 39, and 200 GeV. The total dynamical  $p/\pi$  fluctuations (filled stars) are constructed as  $p/\pi = (p^+ + p^-)/(\pi^+ + \pi^-)$ . As discussed for  $K/\pi$ fluctuations, dynamical fluctuations for  $p/\pi$  can also be constructed for different charge combinations. The average of the same sign (open stars) and opposite sign (open crosses)  $p/\pi$  fluctuations are also included in figure 2 (a). Also included are published measurements from the NA49 experiment [10].

The total charged dynamical  $p/\pi$  fluctuations measured by both STAR and NA49 are negative and become more negative with decreasing collision energy. The two experiments measure coincident values for dynamical  $p/\pi$ fluctuations at energies below  $\sqrt{s_{NN}} = 19.6$  GeV.

Unlike  $K/\pi$  fluctuations, both charge independent and charge dependent dynamical  $p/\pi$  fluctuations are always negative. This indicates a strong cross-correlation exists between produced protons and pions at all ener-



Fig. 2. Results for the measurement of inclusive charged (filled stars) and average same-sign (open stars) and opposite-sign (open crosses), charge dependent  $\nu_{\text{dyn},p/\pi}$  (left) and  $\nu_{\text{dyn},K/p}$  (right).

gies. The strong decay of a  $\Delta$  baryon is a large contributor to this cross-correlation. Examples of this decay that produce cross-correlations would be  $\Delta^{++} \rightarrow p^+ + \pi^+$  or  $\Delta^0 \rightarrow p^+ + \pi^-$ .

The charge dependent results for dynamical  $p/\pi$  fluctuations become more negative at about the same rate with decreasing collision energy, until below  $\sqrt{s_{NN}} = 19.6 \text{ GeV}$ . The anti-proton yield drops dramatically below this energy. The measured  $\overline{p}/p$  ratio decreases by  $\approx$  an order of magnitude between  $\sqrt{s_{NN}} = 17.3$  and 7.6, from 0.1 to 0.01 [11]. Additional study is required of fluctuations involving anti-protons at the lowest energies of  $\sqrt{s_{NN}} = 7.7$  and 11.5 GeV.

The measurement of  $\nu_{dyn,K/p}$  as a function of energy from the STAR experiment, measured in central 0–5% Au+Au collisions at the energies  $\sqrt{s_{NN}} = 7.7, 11.5, 19.6, 39, \text{ and } 200 \text{ GeV}$  is shown in figure 2 (b). The total K/p fluctuations dynamical (filled stars)are constructed as  $K/p = (K^+ + K^-)/(p^+ + p^-)$ . Different charge combinations of dynamical K/p fluctuations are also examined. The average of the same sign (open stars) and opposite sign (open crosses) K/p fluctuations are presented in figure 2(b). Also included are published measurements from the NA49 experiment [12]. Similar to  $p/\pi$  fluctuations, both charge independent and charge dependent dynamical K/p fluctuations are always negative. This indicates a strong cross-correlation exists between produced kaons and protons at all energies. The charge independent dynamical K/p fluctuations become more negative as the collision energy is decreased. Charge dependent dynamical K/p fluctuations become more negative from  $\sqrt{s_{NN}} = 200$  to 19.6 GeV, below which the fluctuations are almost constant. Results from the NA49 experiment for both charge independent and charge dependent K/p fluctuations demonstrate a rapid increase below energies of  $\sqrt{s_{NN}} = 8 \,\text{GeV}$  and actually cross zero and become positive at the lowest energies measured. The results from STAR are consistent with negative dynamical K/p fluctuations at all energies. However, as for dynamical  $p/\pi$  fluctuations, the K/p fluctuations involving anti-protons at the two lowest energies measured by STAR are still under study.

As defined,  $\nu_{\rm dyn}$  is an extensive variable, by virtue of an explicit systemsize (multiplicity) dependence. Therefore, as the system-size increases,  $\nu_{\rm dyn}$ should decrease (toward zero). This is seen in the centrality dependence of  $\nu_{\rm dyn}$ , where it approaches zero from peripheral to central collisions where the multiplicity increases [5]. This also accounts for a part of the overall trends in the energy dependence of  $\nu_{\rm dyn}$ . To create an intensive variable,  $\nu_{\rm dyn}$  scaled by the system-size (particle multiplicity) can be studied. In [5], the scaling used are values of  $dN/d\eta$  for each centrality, corrected for detector efficiency and acceptance. For this study, the uncorrected values of  $dN/d\eta$  from 0–5% central Au+Au collisions at each energy are used.

Figures 3 (a), and 3 (b) show the values of  $\nu_{\text{dyn},K/\pi}$ ,  $\nu_{\text{dyn},p/\pi}$ , and  $\nu_{\text{dyn},K/p}$ , respectively, scaled by uncorrected charged particle multiplicity  $dN/d\eta$ . Figures 3 (a) and 3 (b) also include the scaling result if average uncorrected number of protons + anti-protons ( $\langle p \rangle$ ) is used. Figures 3 (a) and 3 (b) indicates that the uncorrected  $dN/d\eta$  scaled fluctuations involving protons reach a minimum value between  $\sqrt{s_{NN}} = 7.7$  and 19.6 GeV, before increasing toward zero at higher energies. Since the net-baryon density ( $\mu_B$ ) also changes with energy, further investigation of this behavior is required. The smooth scaling with only  $\langle p \rangle$  could reflect the change in net-baryon number with energy, as the ratio of  $\mu_B$  at  $\sqrt{s_{NN}} = 7.7$  and 200 is similar to the ratio of  $\langle p \rangle * \nu_{\text{dyn},p/\pi}$  at those energies.



Fig. 3. Results for the measurement of  $\nu_{\text{dyn},p/\pi}$  (left) and  $\nu_{\text{dyn},K/p}$  (right) scaled by uncorrected charged particle multiplicity,  $dN/d\eta$ , from central 0–5% Au+Au collisions at  $\sqrt{s_{NN}} = 7.7$ –200 GeV.

### 3. Summary

Results from dynamical particle ratio fluctuations  $(K/\pi, p/\pi, \text{and } K/p)$ and new results on the charge dependence of dynamical fluctuations have been presented from data acquired in Au+Au collisions at energies from  $\sqrt{s_{NN}} = 7.7-200$  GeV. Also discussed are initial results from multiplicity scaled dynamical fluctuations of all three particle ratios at energies from  $\sqrt{s_{NN}} = 7.7-200$  GeV.

The charge dependent dynamical fluctuations provide additional insight into particle production as a function of energy. Further study is necessary at all energies, including plotting each charge combination separately. Uncorrected multiplicity scaled dynamical  $p/\pi$  and K/p fluctuations become more negative between  $\sqrt{s_{NN}} = 7.7$  and 19.6 GeV, before increasing toward zero at higher energies.

Additional data points at  $\sqrt{s_{NN}} = 27$  and 62.4 GeV are under study and will complete the excitation function for the first phase of the RHIC energy scan to search for the QCD critical point.

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